

DOSIMETRY AND MICRODOSIMETRY
ONBOARD ISS AND RELATED TOPICS
2004- 2005

**František SPURNÝ, Iva JADRNÍČKOVÁ,
Karel TUREK**

*Department of Radiation Dosimetry,
Nuclear Physics Institute,
Czech Academy of Sciences, Prague*

WRMISS Chiba, September 2005

Experiments and analysis 2004- 2005

- **On-Earth's calibrations**
 - **ICCHIBAN 6 (C, Ar, Kr; 24 - 600 keV/μm)**
 - **ICCHIBAN-NSRL (H, O, Fe; 0.2 – 150 keV/μm)**
 - **Dubna Nuclotron (C, Mg, Fe; 8 – 200 keV/μm)**
- **Onboard ISS exposures**
 - **January to October 2004; russian module**
- **Other analysis**
 - **Influence of sensitive volume dimensions on the microdosimetry distributions**

Thermoluminescent detectors (TLD's)

$\text{Al}_2\text{O}_3:\text{C}$

- $H^*(10) \geq 1 \mu\text{Sv}$
- rapid decrease of light conversion factor (relative response RR) with LET above $\sim 1 \text{ keV}/\mu\text{m}$

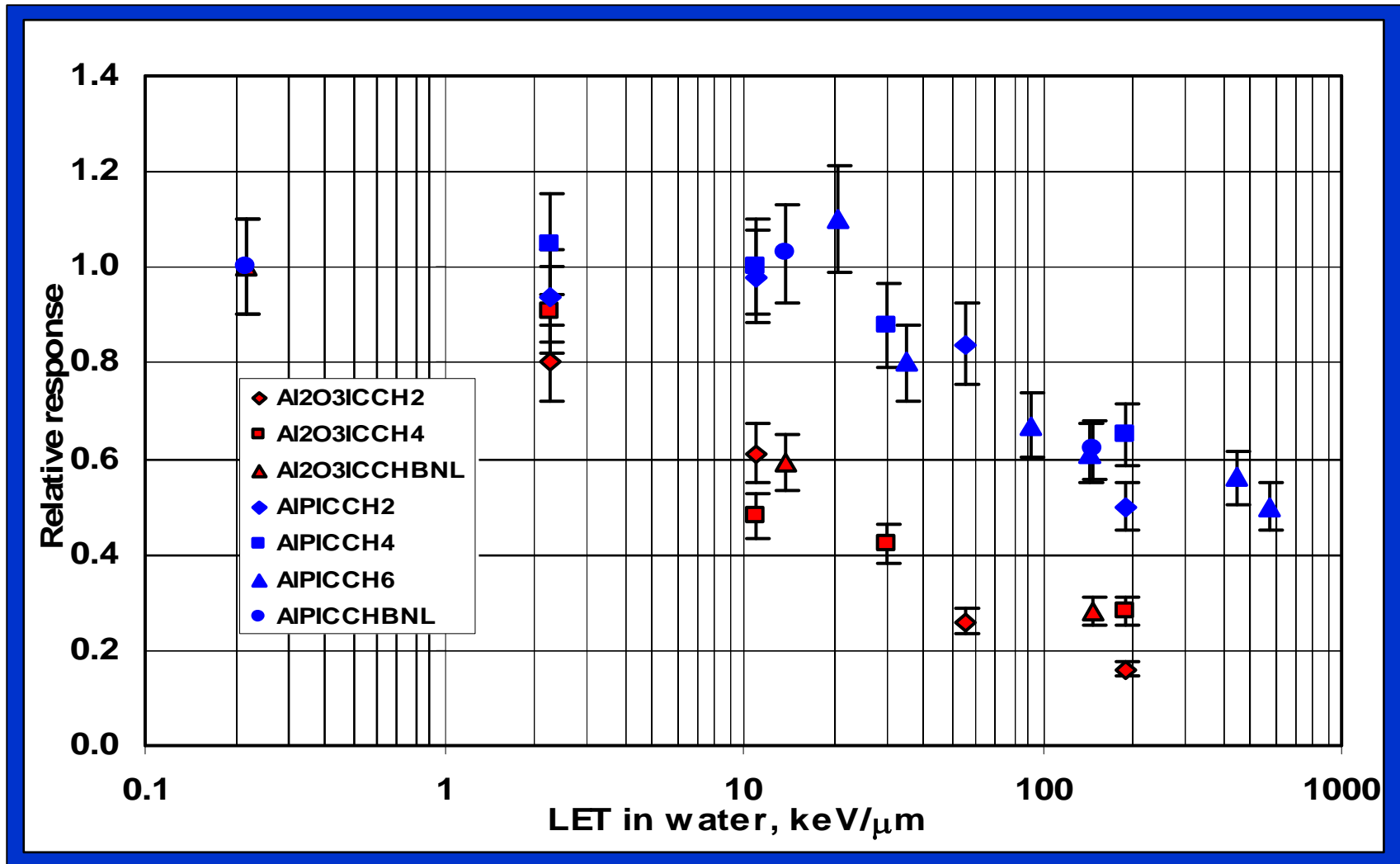
Czech alumophosphate (AIP) TL glass

- $H^*(10) \geq 10 \mu\text{Sv}$
- slower decrease of relative response RR with LET above $\sim 1 \text{ keV}/\mu\text{m}$

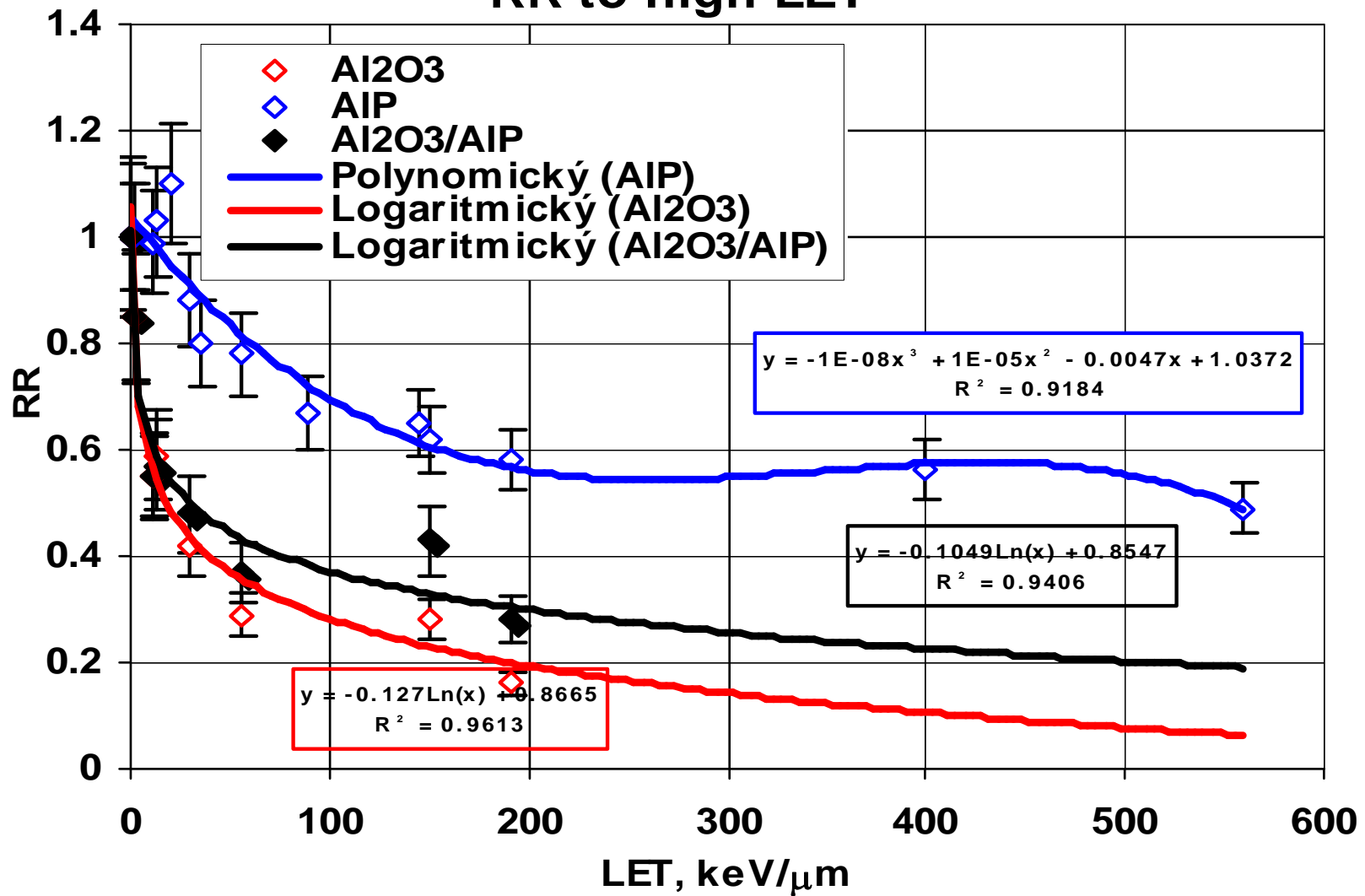
LiF's from IFJ Krakow

- MTS-6; MTS-7; MTT-7; MCP-N; MCP-7
- different decrease of relative response RR with LET above $\sim 1 \text{ keV}/\mu\text{m}$

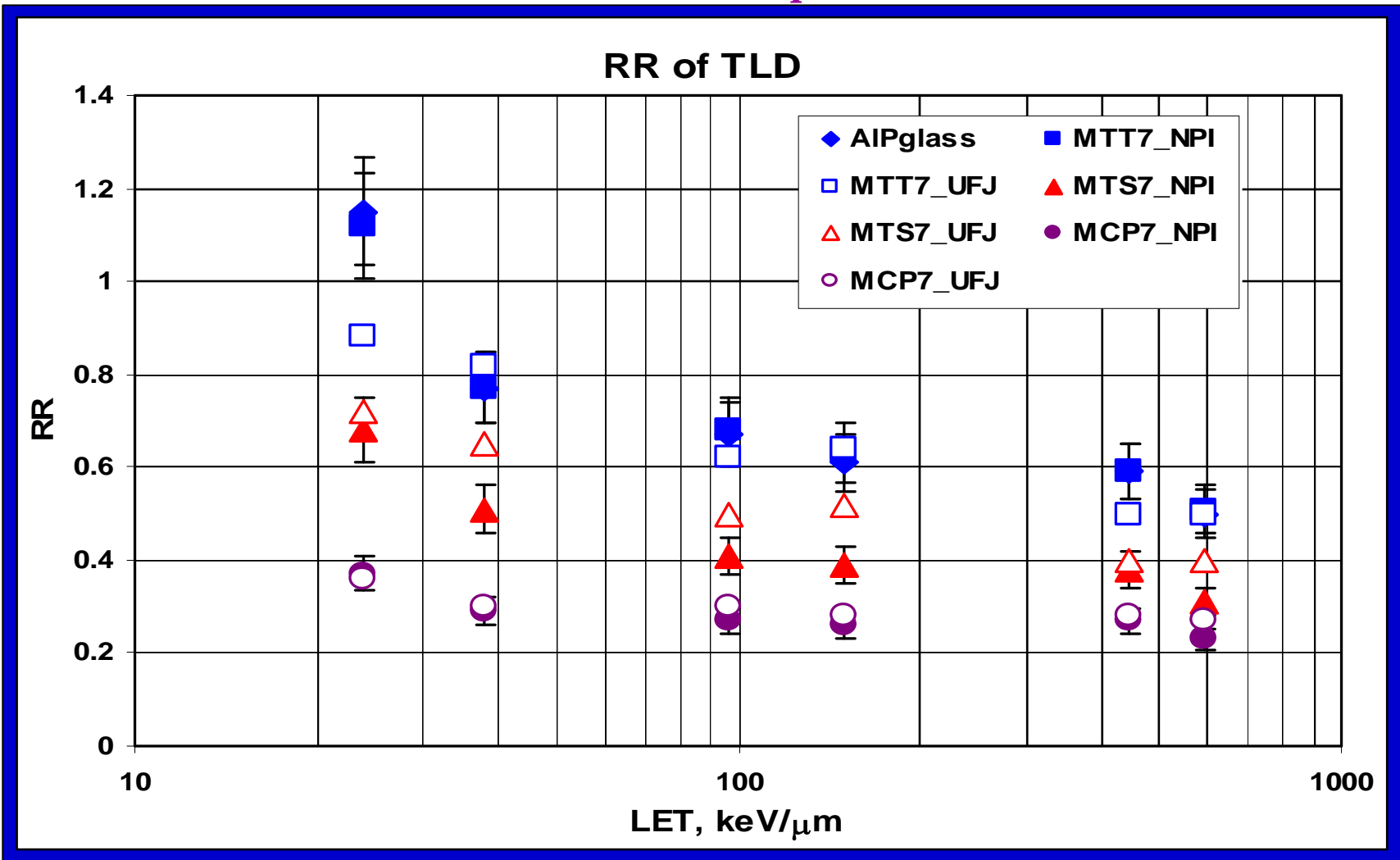
LET dependence of the TL relative response (RR)



RR to high LET



Comparison of relative responses obtained during ICCHIBAN 6 experiment



NPI – Prague (full symbols);

UFJ - Krakow [Bilski & Olko, WRMISS Vienna 2004], approx. (open symbols)

WRMISS Chiba, September 2005

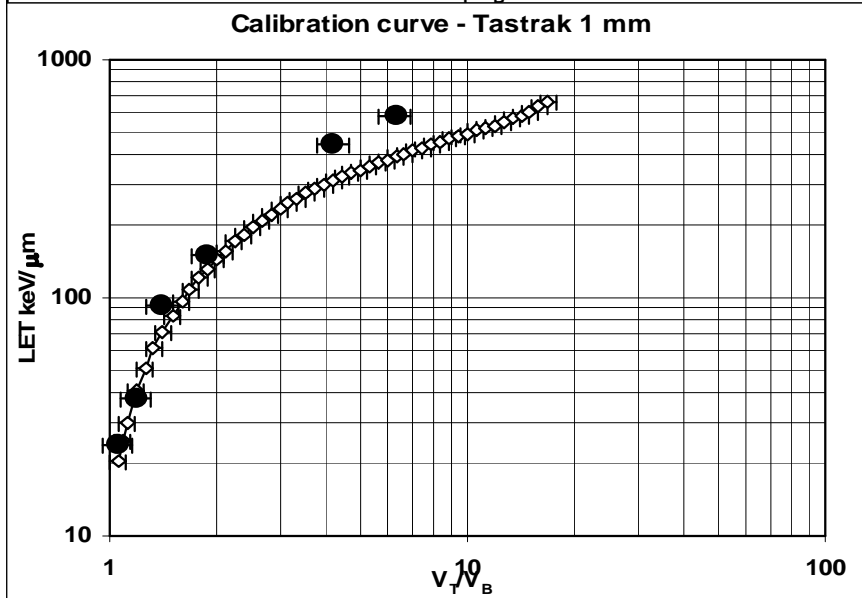
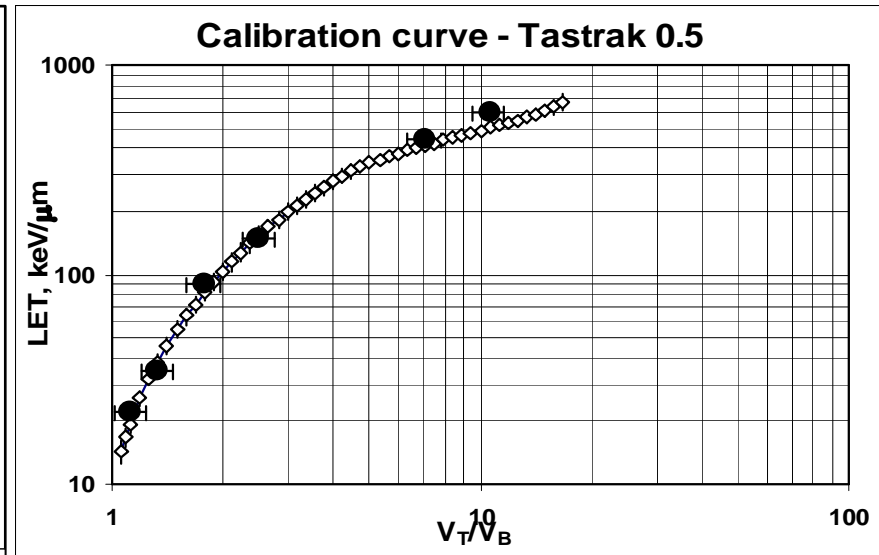
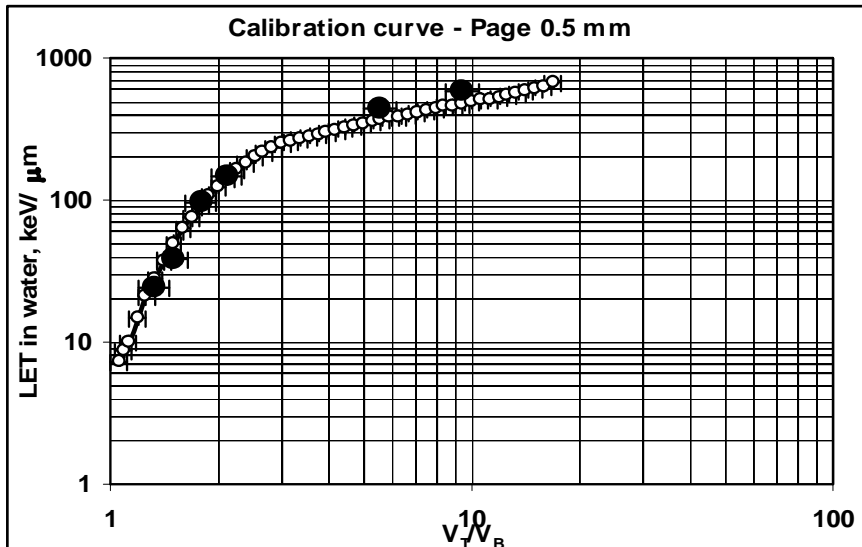
Track etch detectors (TED's)

LET spectrometer based on chemically etched PADC TED

Material	LET range keV/ μm	Range of H mSv
Page, 0.5 mm thick	7 – 700	1 - 100
Tastrak, 0.5 mm thick	15 – 700	
Tastrak, 1 mm thick	22 – 700	

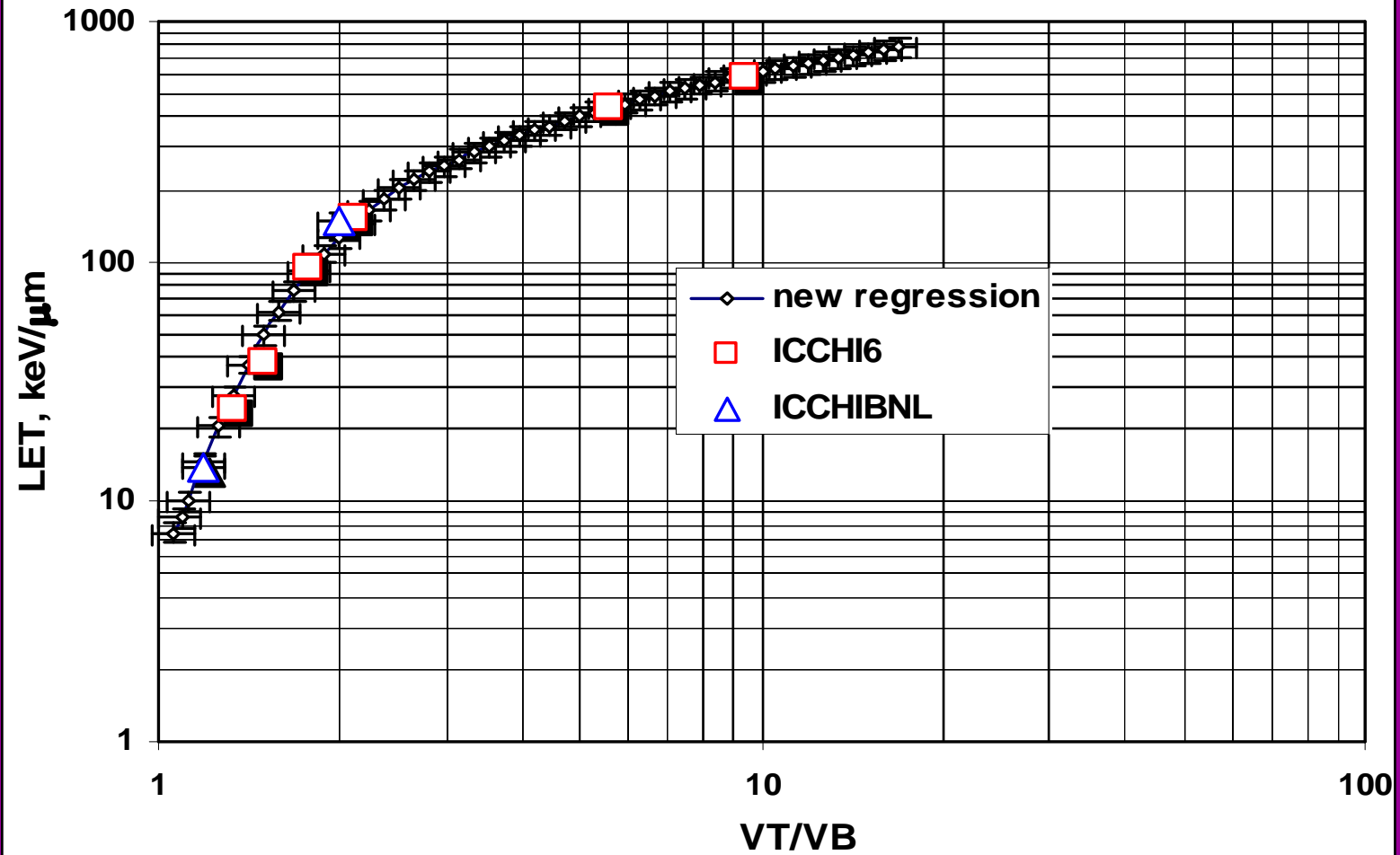
- **LET spectrometer: Etching** - 5 N NaOH, 70°C; 18 h, $\Delta h \approx 17 \mu\text{m}$; to determine **LET** - etching rate ratio $V=V_T/V_B$ established through the determination of track parameters;
- **PADC etched in 30% KOH**, both chemically and electrochemically – to determine angular responses

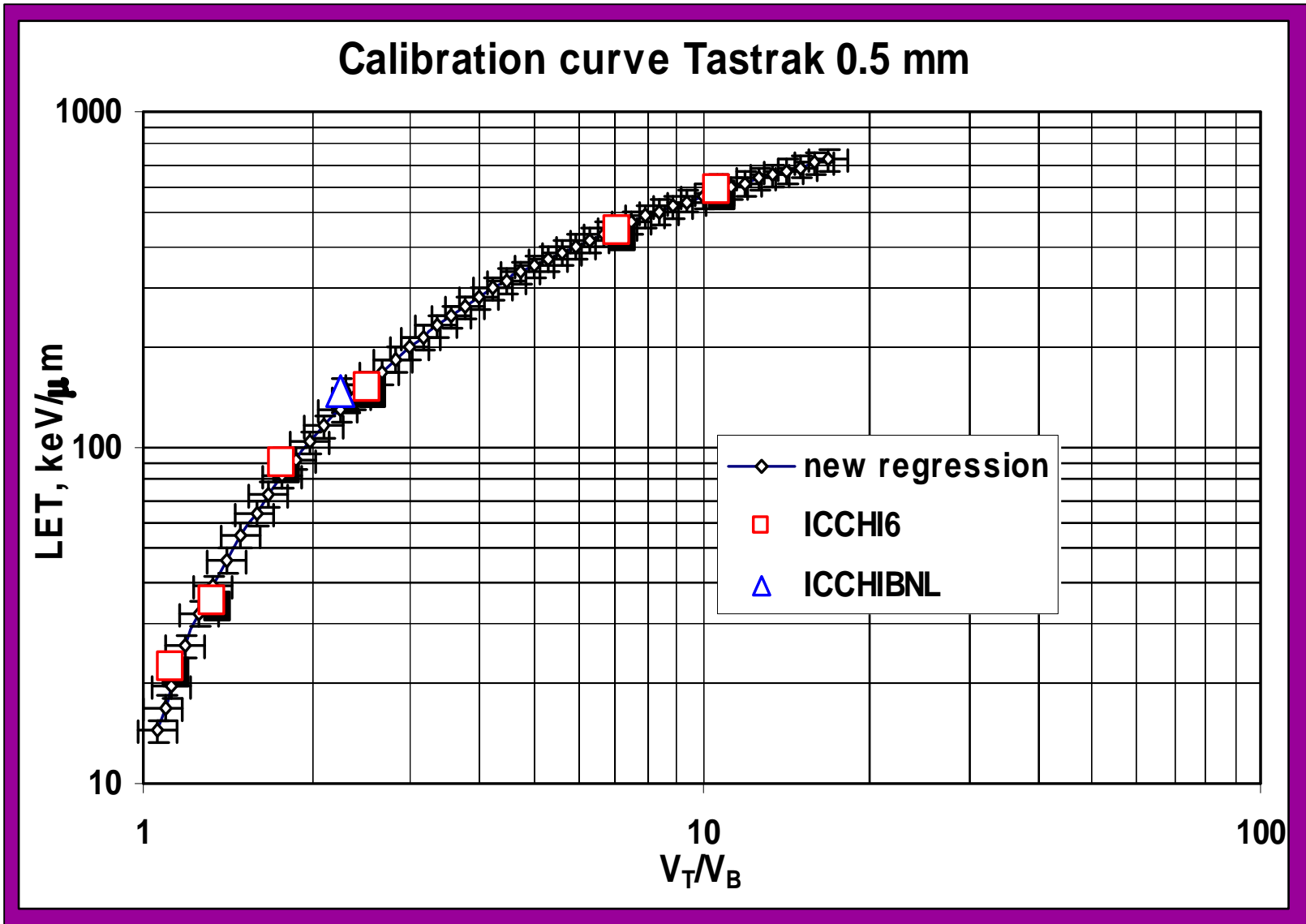
Previous regressions and ICCHI 6 & NSRL

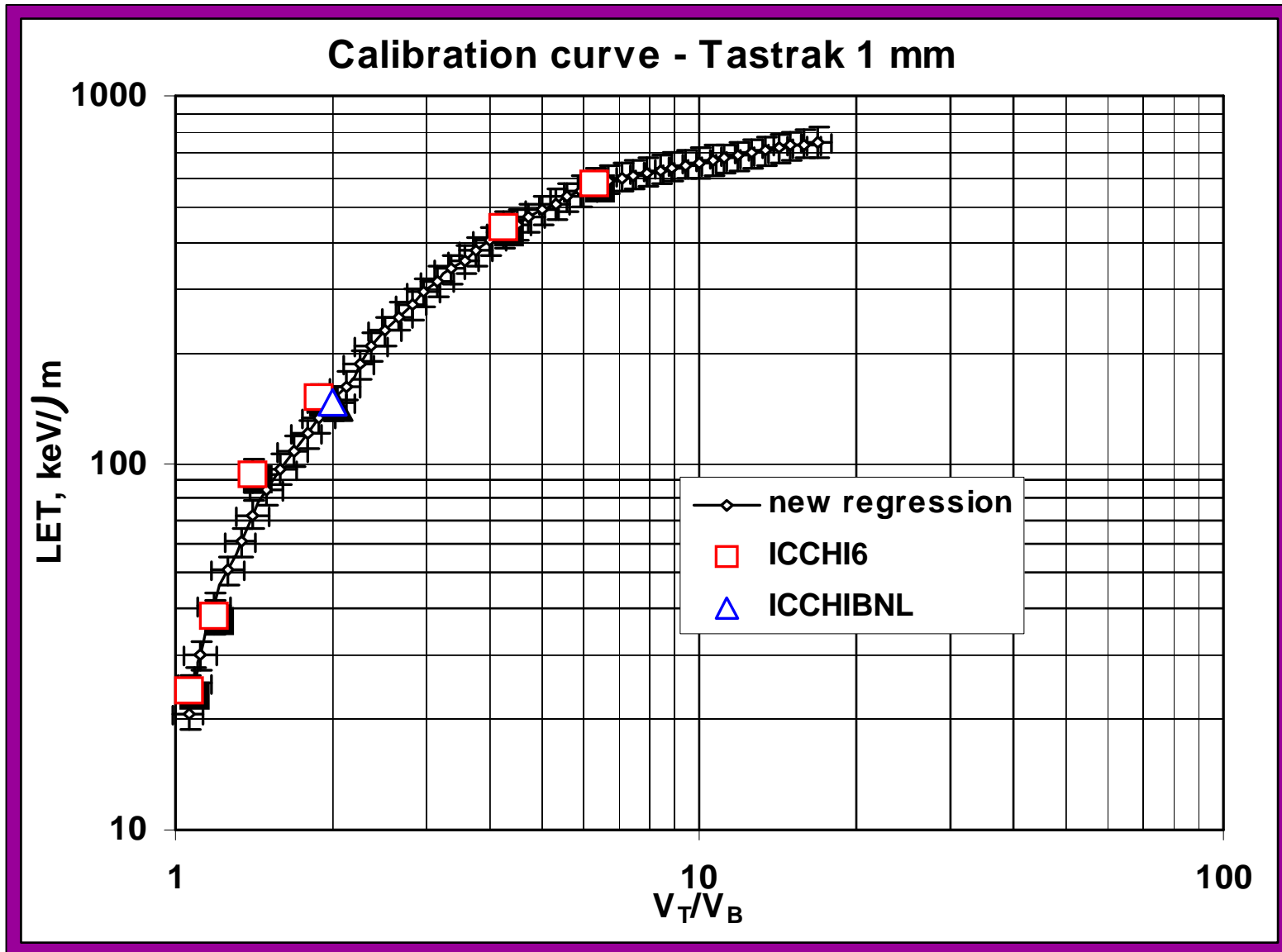


• and now:

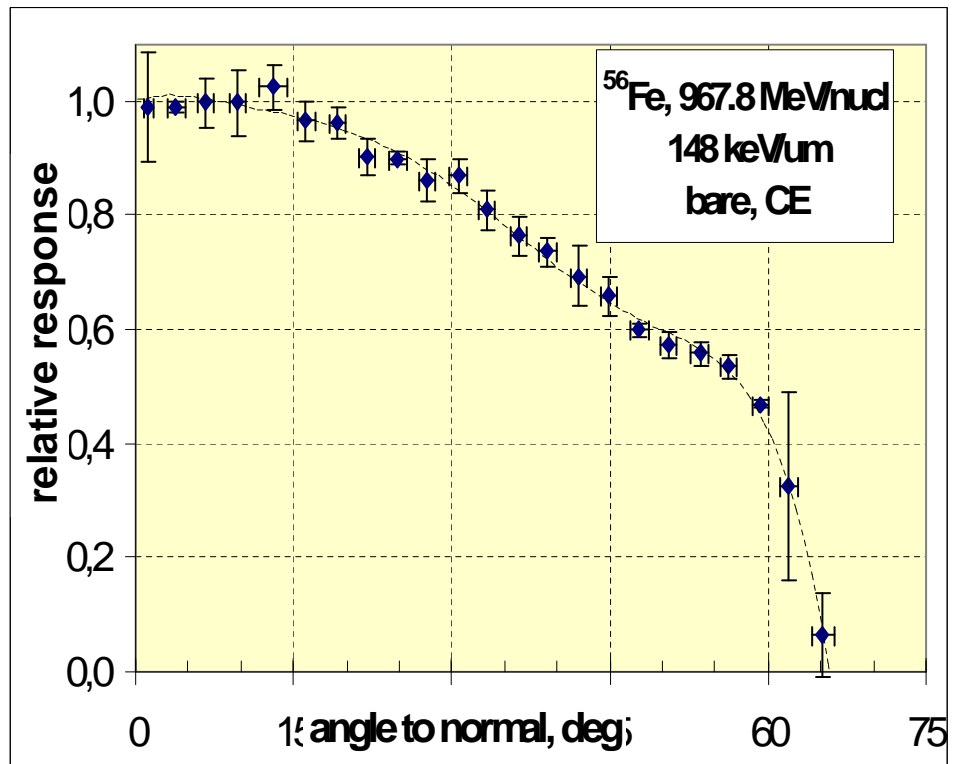
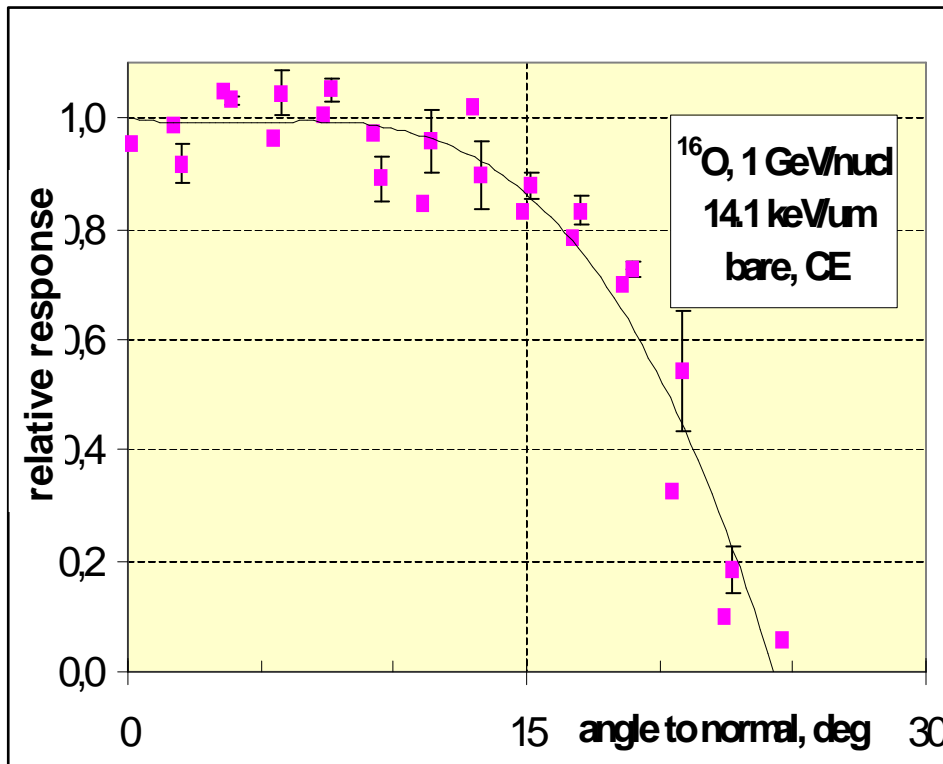
Calibration curve Page



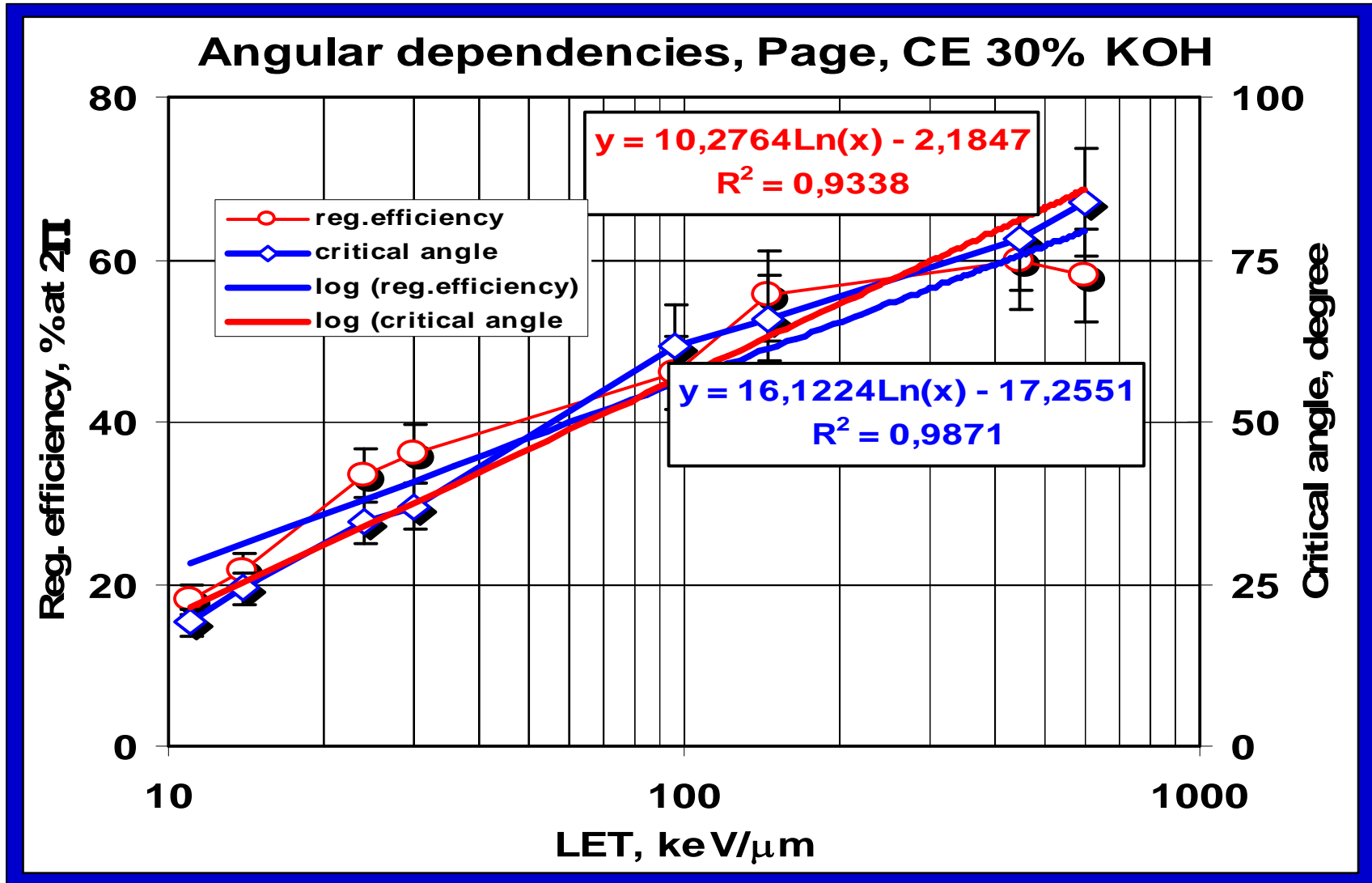




Angular dependences - ICCHIBAN NSRL 30% KOH



Angular dependences - ICCHIBAN's summary; CE 30% KOH



Registration efficiencies for chemically^{*)} etched bare PADC's and ICCHIBAN6 ions

Problem – systematically lower than 1.0 – never observed before, also not at NSRL

Ion	Page	Page	Tastrak 0.5	Tastrak 1.0
	etched 30%KOH	etched 5N NaOH	used for LET	spectrometry
C	0.73 ± 0.04	0.73 ± 0.04	0.77 ± 0.05	0.87 ± 0.05
Ar	0.67 ± 0.03	0.66 ± 0.04	0.73 ± 0.05	0.82 ± 0.04
Kr	0.59 ± 0.02	0.65 ± 0.05	0.67 ± 0.05	0.74 ± 0.05

^{*)} Only Kr-ions revealed by ECET, with the registration efficiency (0.46±0.01)

Registration efficiencies for chemically^{*)} etched bare PADC's and ICCHIBAN-NSRL ions

Ion	Page	Page	Tastrak 0.5	Tastrak 1.0
	etched 30%KOH	etched 5N NaOH	used for LET	spectrometry
O	0.79 ± 0.11	1.20 ± 0.08	0.80 ± 0.13	-
Fe	0.95 ± 0.03	1.10 ± 0.06	1.03 ± 0.04	1.10 ± 0.07

***) Neither O nor Fe-ions revealed by ECET**

ICCHIBAN6 blinds - D_{LET} (above ~ 10 keV/ μ m) and D_{TLD} (below ~ 10 keV/ μ m)

Blind No.	D_{LET} , mGy	D_{TLD} (NPI) mGy
1	0.45 ± 0.07	72.0
2	0.29 ± 0.06	71.2
3	0.44 ± 0.08	74.8
4	0.62 ± 0.18	60.9
5	≥ 0.42	73.0
6	0.77 ± 0.13	104.1

ICCHIBAN6 blinds - D measured with TLD's

Blind No.	D, mGy, as measured with TLD:			
	AIP glass ^{*)}	MTS 7 ^{**)}	MTT 7 ^{**)}	MCP 7 ^{**)}
1	72.0	96.1	98.4	104.3
2	71.2	94.3	96.6	100.4
3	74.8	98.0	99.4	105.0
4	60.9	86.5	91.8	80.0
5	73.0	99.4	99.8	105.6
6	104.1	159.4	105.6	136.7

*) 1 S.D. relative ~ 5%;

***) presented by Bilski & Olko [WRMISS 2004]; without correction for LET

Remark: MTT/MCP = (0.949 ± 0.005) for blinds 1,2,3, and 5;
= 1.15 (1.16) for blind 4 (6)

ICCHIBAN6 blinds - Remarks

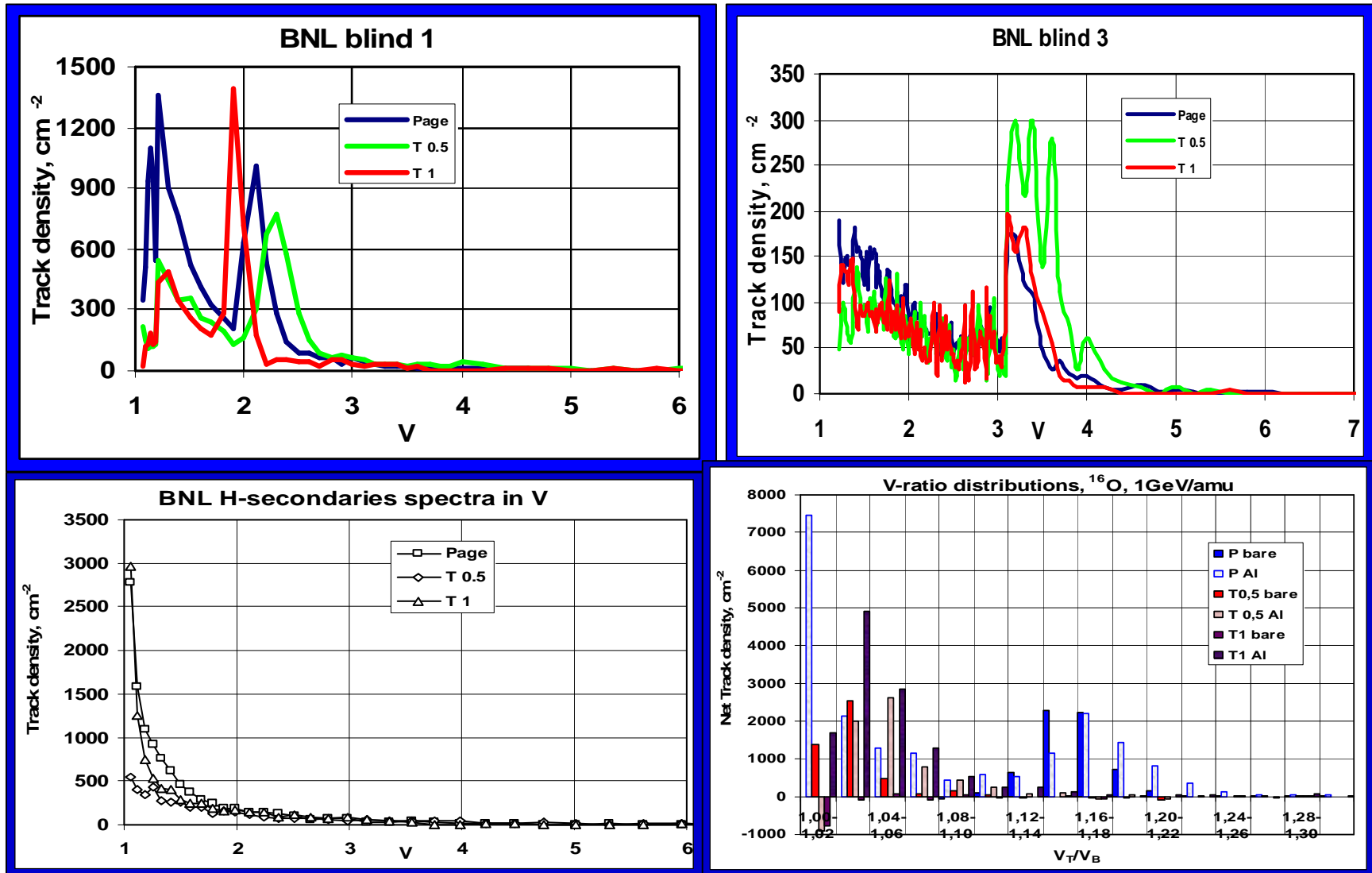
1. For all blinds – low LET radiation (<10 keV/ μm) dominating in the dose
2. When **TLD readings and theirs ratios** considered, it could be deduced that:
 - Average ratio AIP/MTT (similar dependence of $RR \Rightarrow LET$) equal to (0.738 ± 0.012) – systematic difference in exposure levels as for track detectors?
 - Blinds 1,2,3, and 5 were exposed mostly to the radiation with LET below $\sim 1-2$ keV/ μm ;
 - Exposure of blinds 4, and 6 to the radiation with LET between 2 and 10 keV/ μm (not registered by means of LET spectrometer, decreasing RR of MCP-7 as compared to MTT-7 or AIP glass).

ICCHIBAN-NSRL blinds - D estimated with TLD's

Blind No.	Direct reading, mGy		Corrected reading, mGy		Average value mGy
	AlP glass	Al ₂ O ₃ :C	AlP glass	Al ₂ O ₃ :C	
1	29.7 ^{*)}	29.5	29.7	29.5	29.6±4.4
2	29.5	30.4	29.5	30.4	30.0±4.5
3	1.41	0.48	2.32	1.96	2.14±0.32
4	0.36	0.13	0.59	0.48	0.53±0.08

***) 1s relative of read values about ± 5%;
uncertainty of correcting procedure estimated to 15%**

NSRL exposures: some of V-ratio spectra



ICCHIBAN-NSRL blinds - full evaluation

1. Considering the LET spectra of registered particles, it could be, it seems, to be deduced: sets for blinds Nos. 1 and 2 were exposed to Fe-ions; sets for blinds Nos. 1 and 4, and, perhaps also 3, were exposed to O-ions; in all sets secondary particles due to protons, and/or fragments?
2. Total doses, due to the particles with the LET above about $\sim 10 \text{ keV}/\mu\text{m}$, D_{LET} , have been calculated supposing that estimated particles are registered with efficiency ~ 1.0 , proton created secondary particles (and fragments?) have angular distribution expected when the isotropicity in the center-of-mass is preserved.
3. D_{LET} and D_{TLD} (\sim total) are presented in the Table.

Blind No.	Average D_{LET} , mGy	Average D_{TLD} , mGy
1	1.51 ± 0.28	29.6 ± 4.4
2	1.87 ± 0.35	30.0 ± 4.5
3	2.52 ± 0.60	2.14 ± 0.32
4	0.74 ± 0.32	0.53 ± 0.08

Direct TLD and TED LET spectrometer readings onboard Space Stations

Mission	TLD dose, $\mu\text{Gy/day}$	D -TED $\text{LET} \geq 10\text{keV}/\mu\text{m},$ $\mu\text{Gy/day}$	H60 -TED $\text{LET} \geq 10\text{keV}/\mu\text{m},$ $\mu\text{Sv/day}$
MIR 28 - 2000	$140 \pm 10$¹⁾	13.1 ± 0.9	85 ± 5
ISS – 11/01-11/02	212 ± 15	22 ± 2	202 ± 12
ISS – 01/04-10/04	166 ± 14	10 - 12	82 - 132

¹⁾ Here and in all other cases - 1 S.D.

Proton's and neutron's contribution

Mission	D-TED, LET \geq 10 keV/ μ m			H60-TED, LET \geq 10 keV/ μ m		
	total	protons	neutrons	total	protons	neutrons
	μ Gy/day			μ Sv/day		
MIR 28	13.1 \pm 0.9 ¹⁾	3.5 \pm 0.4	9.5 \pm 1.1	85 \pm 5	25.5 \pm 3.0	60 \pm 6
ISS - 01/02	22 \pm 2	5.3 \pm 0.6	16.7 \pm 1.9	202 \pm 12	61 \pm 8	141 \pm 16
ISS - 04	10 - 12	3.9 \pm 0.5	6 - 7	82 - 132	47 \pm 3	35 - 85

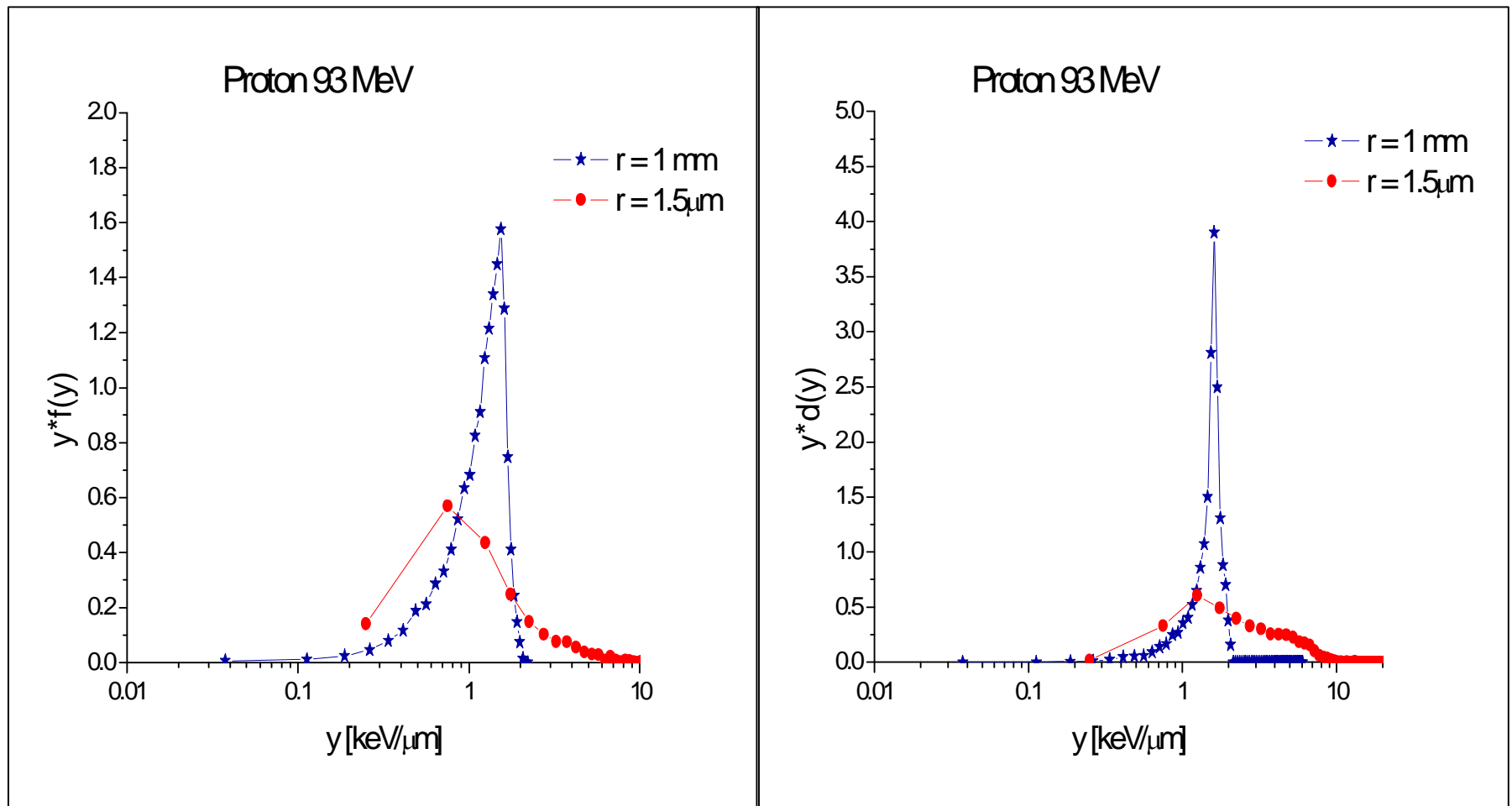
- Supposed:**
- 1) TLD dose \cong Dose due to high energy CP (protons)
 - 2) Relative response $D(> 10 \text{ keV}/\mu\text{m})/D(\text{ionization losses}) \sim 0.025$
 - 3) $D(\text{H})$ from neutrons – relative response ~ 1.0

Full dosimetric characteristics onboard of space stations

Mission	D, $\mu\text{Gy/day}$		Neutrons in % of D	H, $\mu\text{Sv/day}$		Neutrons in % of H
	>10 keV/ μm	Total		>10 keV/ μm	Total	
MIR 28	17 ¹⁾	157	6.0	129	269	22
ISS - 01/02	28	240	7.0	306	518	27
ISS - 04	13 - 15	171- 186	~ 4	124-200	280-360	13-24

1. TLD's data characterize the contribution mostly of radiation with LET lower than few keV/ μm ;
2. The contribution of primary long range cosmic heavier charged particles represents, as estimated by O'Sullivan, about 22% of total LET spectrometer signal in dose, about 34% in dose equivalent (ICRP 60 QF).

Microdosimetry distribution as a function of sensitive volume dimension - calculated by TRIOL MC code for TE sensitive volume



Acknowledgements

- Many of results presented in this contribution were obtained as part of the ICCHIBAN research project using heavy ions at HIMAC-NIRS, NSRL–BNL, nuclotron JINR Dubna, and protons at Loma Linda UMC. We are much obliged to the staff of all laboratories and, particularly, to organizers of the run, Y. Uchihori, J. Miller, E.R. Benton, A.G. Molokanov, and V.P. Bamblevski for their help. We are also much obliged to our colleagues from the IMBP of the Russian Academy of Sciences, V.A. Shurshakov, and Y.A. Akatov Moscow for their help in the studies performed onboard of ISS. Also we thank to P.Bilski and P. Olko (UFJ Krakow) for TLD cooperation
- Studies were also partially supported through the grant No. 202/04/0795 of the GA CR and the IRP AV0Z10480505.

Thank you for your attention !