

Calculation of Radiation Exposure Levels in Low Earth Orbit and Beyond

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Knowledge for Tomorrow

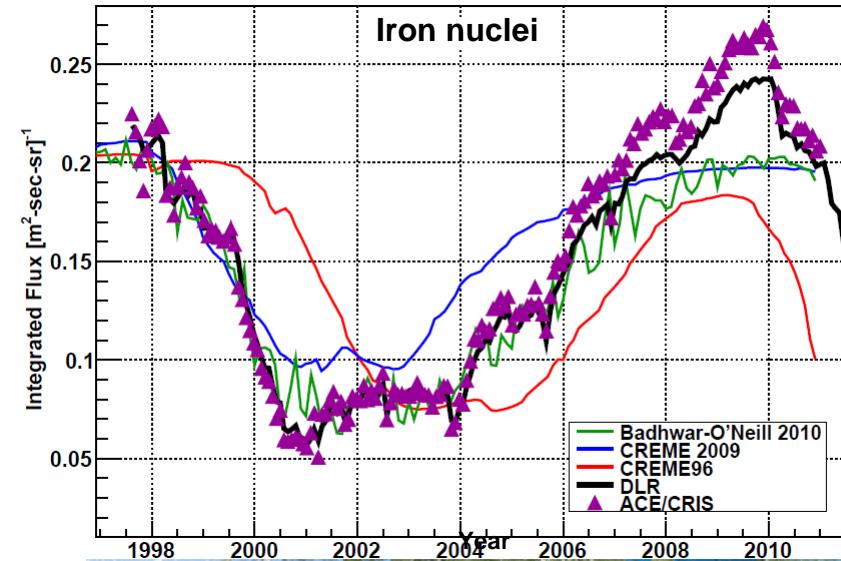
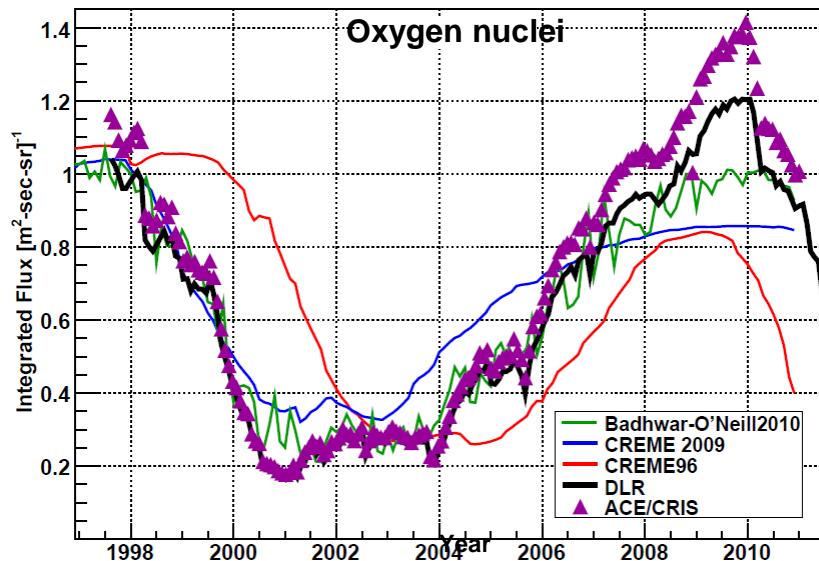
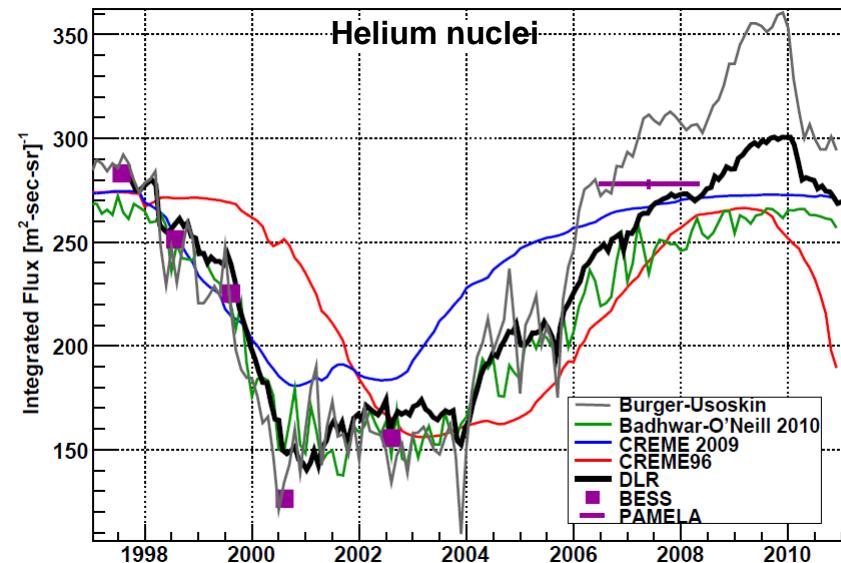
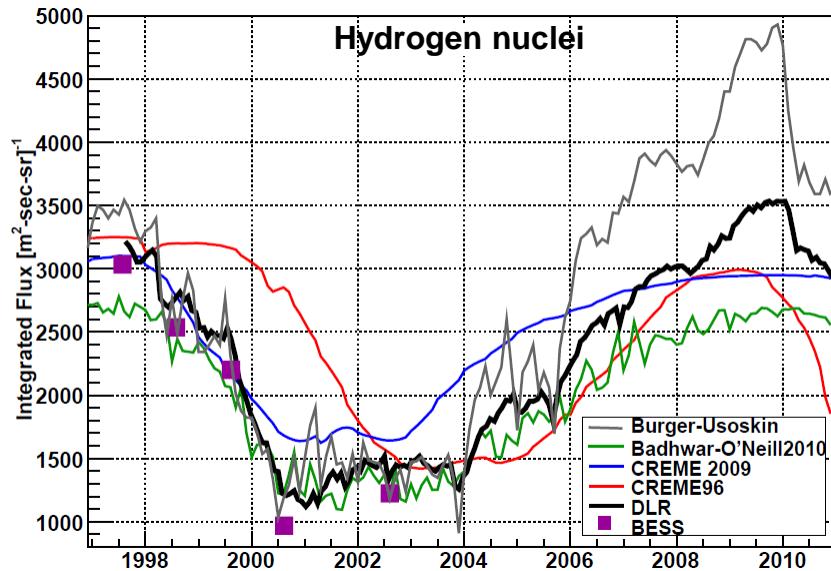


Outline

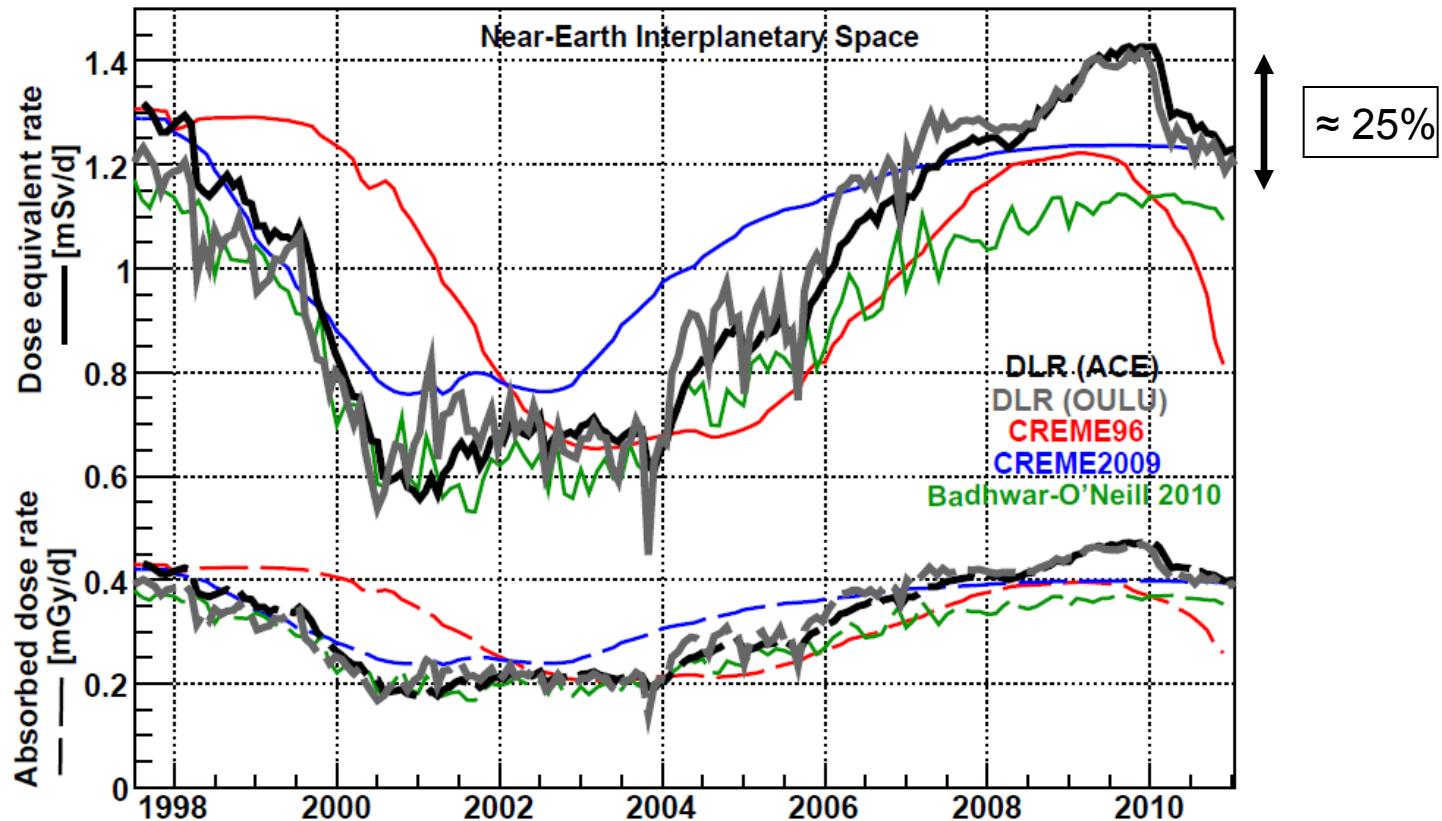
- GCR model comparison, particle fluxes / dose rate estimates
- Badhwar-O'Neill 2011 model
- Estimates of the radiation exposure
 - Interplanetary space
 - ISS orbit
 - Solar minimum / DOSIS / DOSIS 3D
 - Implementation of COLUMBUS shielding distribution
 - GCR / trapped protons, organ dose rates
- Shielding dependence of organ doses for spherical shielding geometry



Temporal variation in GCR flux (17th WRMISS)



GCR exposure in interplanetary space (17th WRMISS)



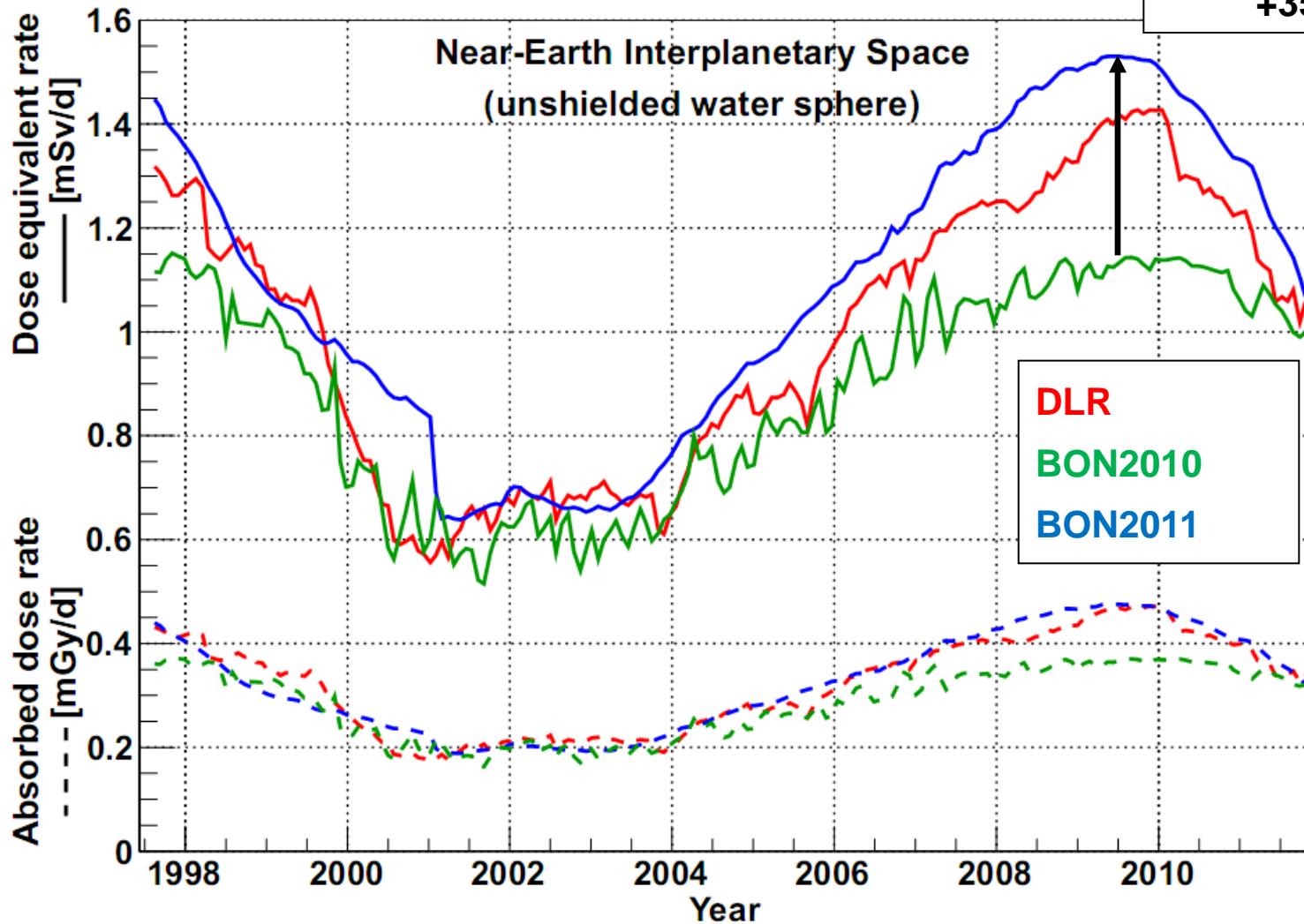
dD/dt ($\mu\text{Gy}/\text{d}$)	173 – 472
dH/dt ($\mu\text{Sv}/\text{d}$)	556 – 1427

GCR exposure in interplanetary space

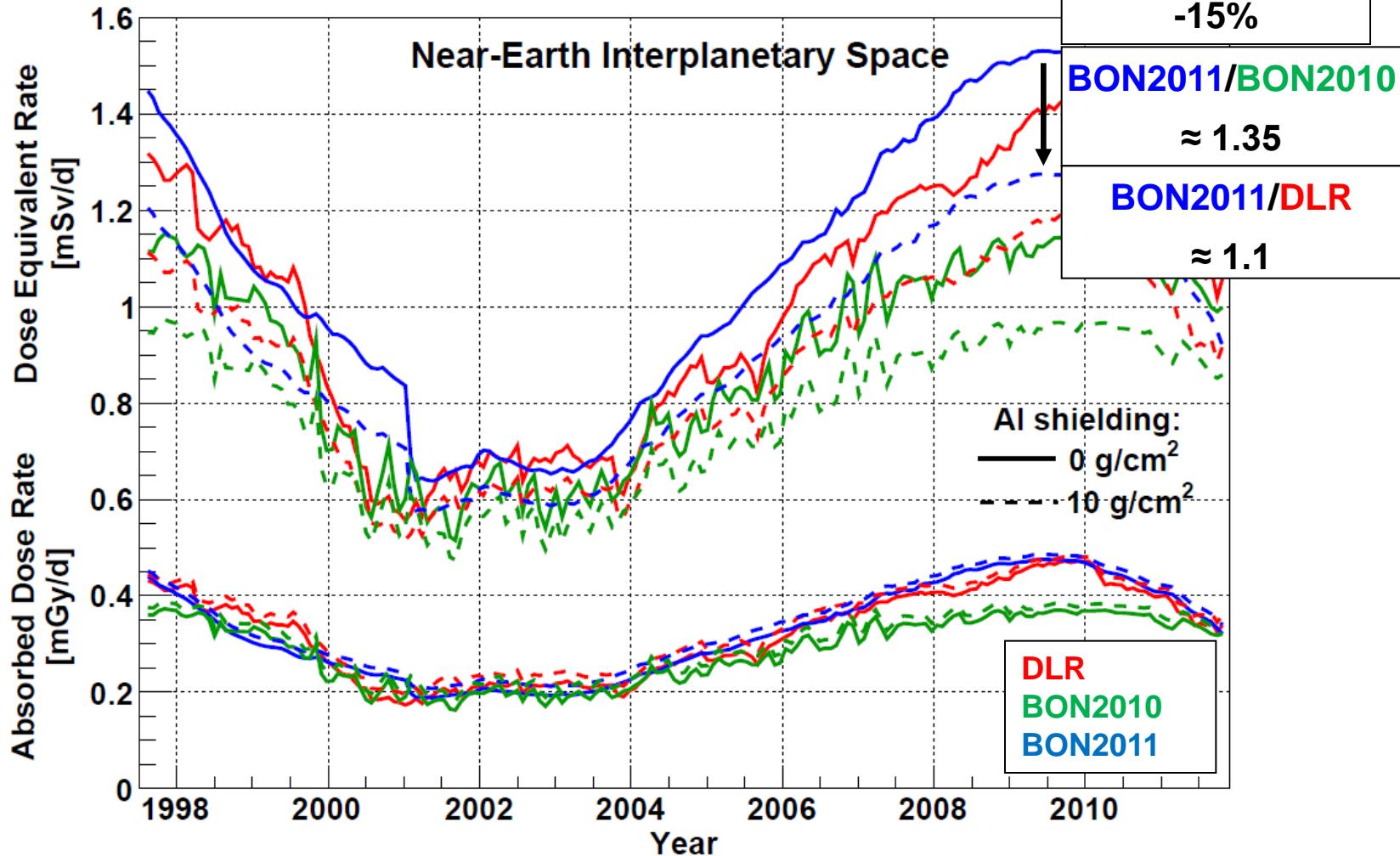
BON2010 → 2011

+35%

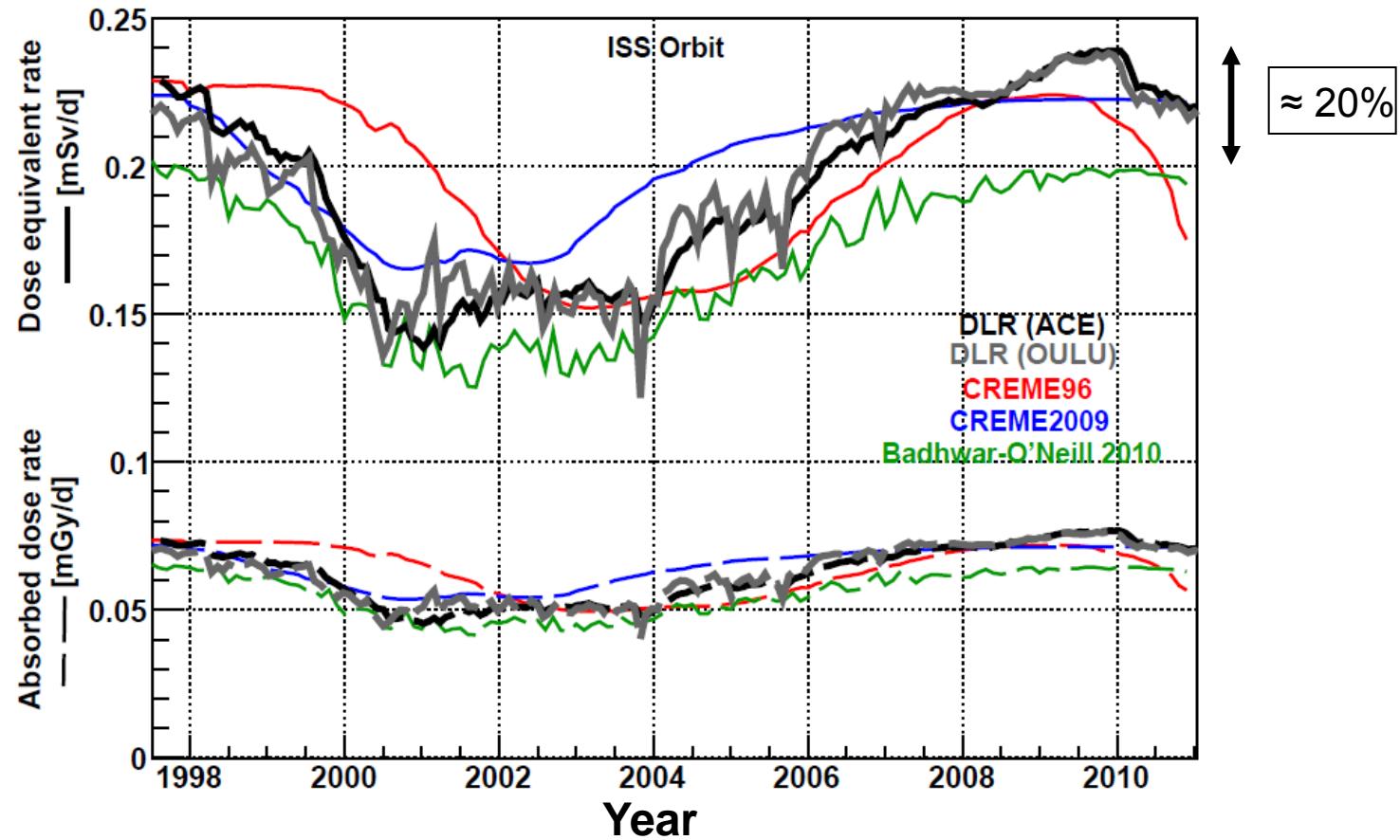
2012:
< 10%



GCR exposure in interplanetary space



GCR exposure in low Earth orbit (17th WRMISI)



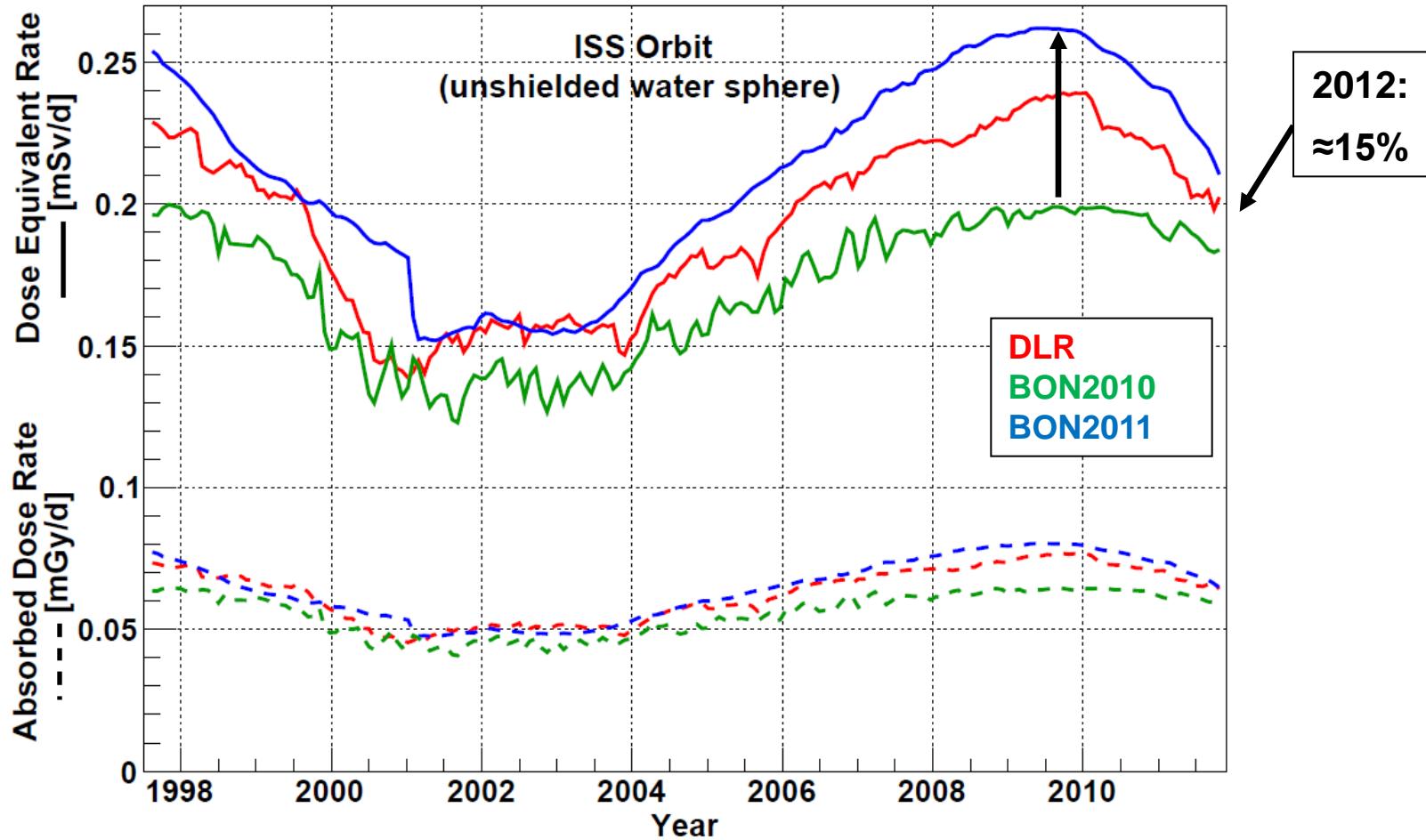
dD/dt ($\mu\text{Gy}/\text{d}$)	45 – 77
dH/dt ($\mu\text{Sv}/\text{d}$)	139 – 239

GCR exposure in LEO

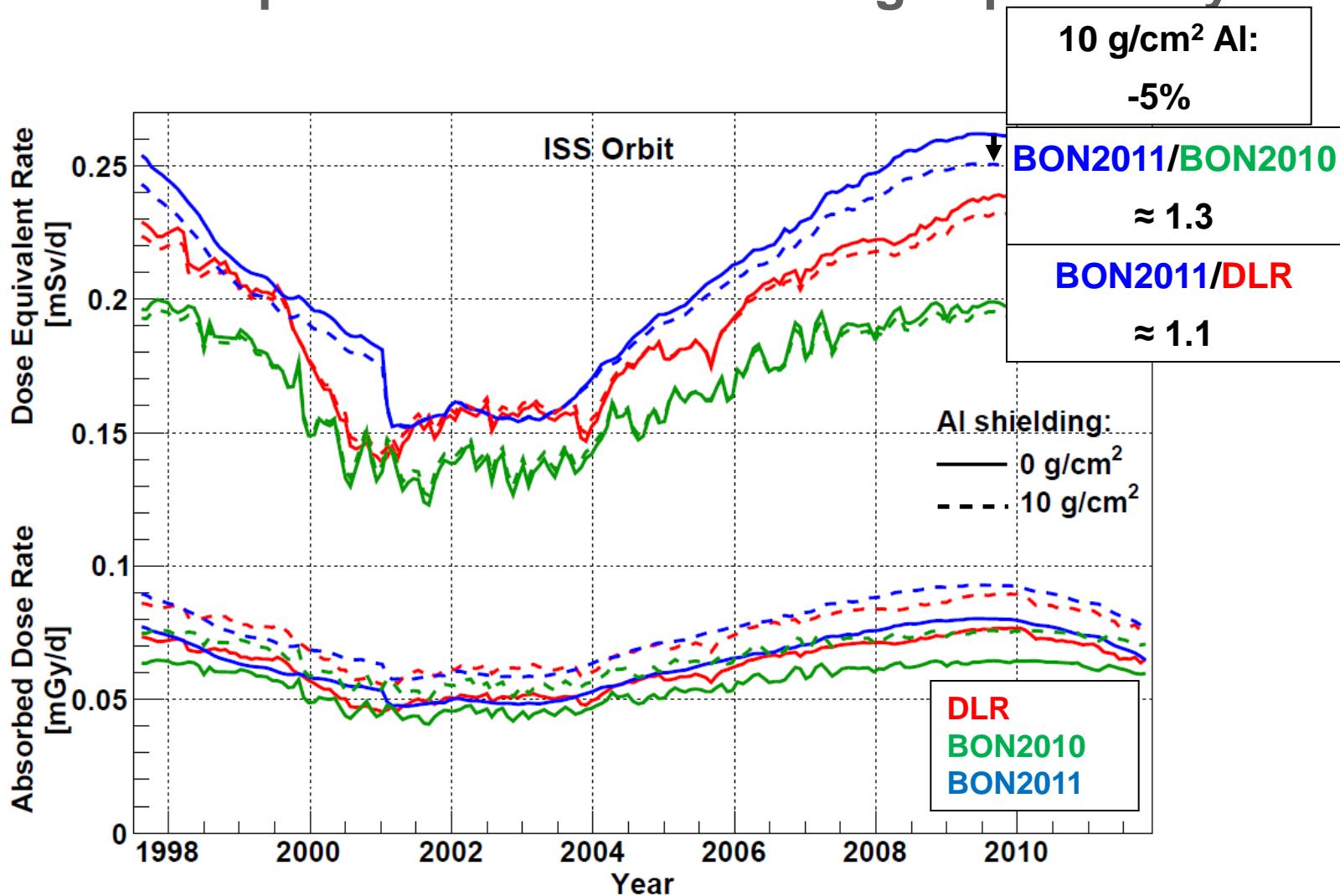
BON2010 → 2011

+30%

2012:
≈15%



GCR exposure in LEO - shielding dependency

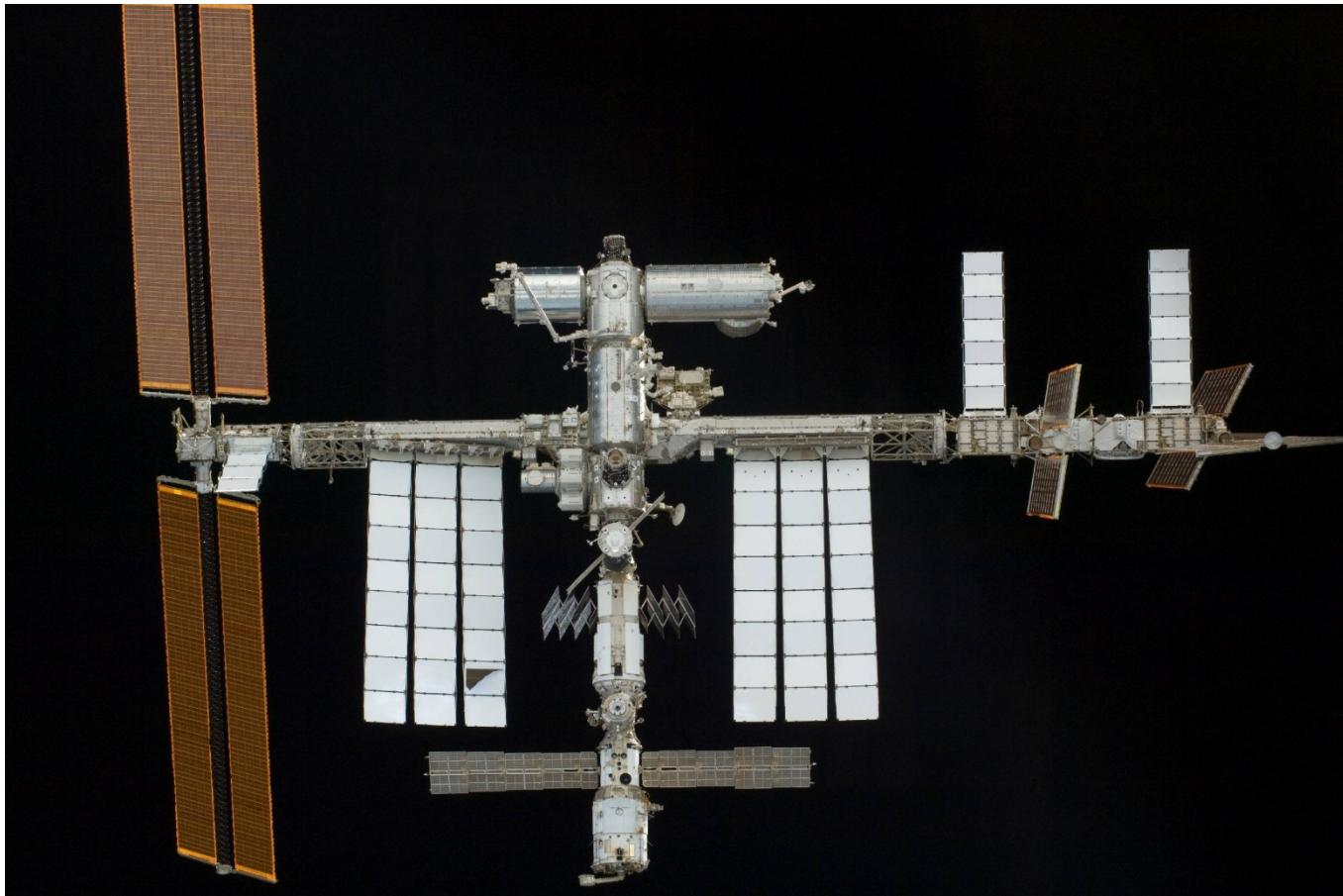


Summary of model comparison

- During solar minimum/GCR maximum 2009:
 - Interplanetary space:
differences \approx 25% (without BON2011), 1.15 – 1.42 mSv/d
 \rightarrow BON2011/BON2010 \approx 1.35
 \rightarrow BON2011/DLR \approx 1.1
10 g/cm² Al shielding \rightarrow -15% in dose equivalent rate
 - Low Earth Orbit
differences \approx 20% (without BON2011), 0.2 – 0.24 mSv/d
 \rightarrow BON2011/BON2010 \approx 1.3
 \rightarrow BON2011/DLR \approx 1.1
10 g/cm² Al shielding \rightarrow -5% in dose equivalent rate
- 2012: Differences reduced: <10% (interplanetary space); \approx 15% (LEO)



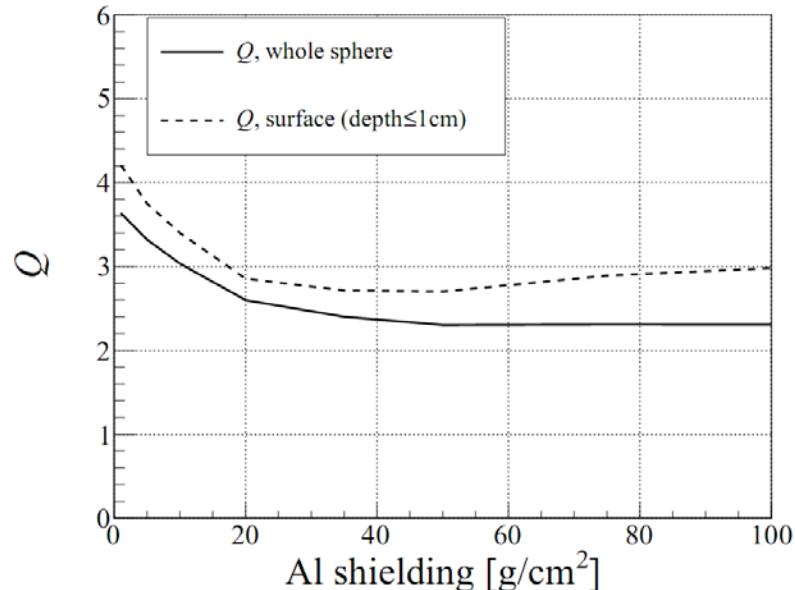
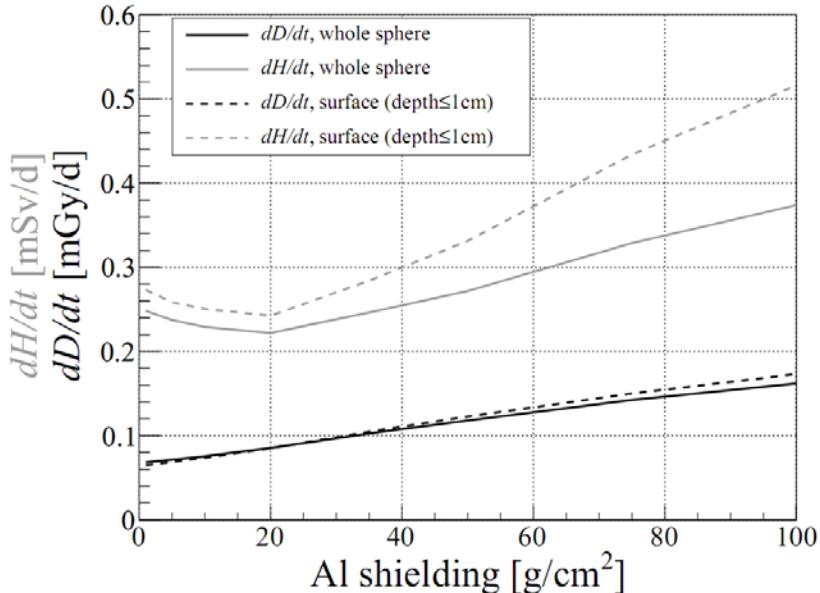
Radiation exposure at the ISS orbit



GCR radiation exposure on ISS Dec. 2009 (max. GCR intensity)

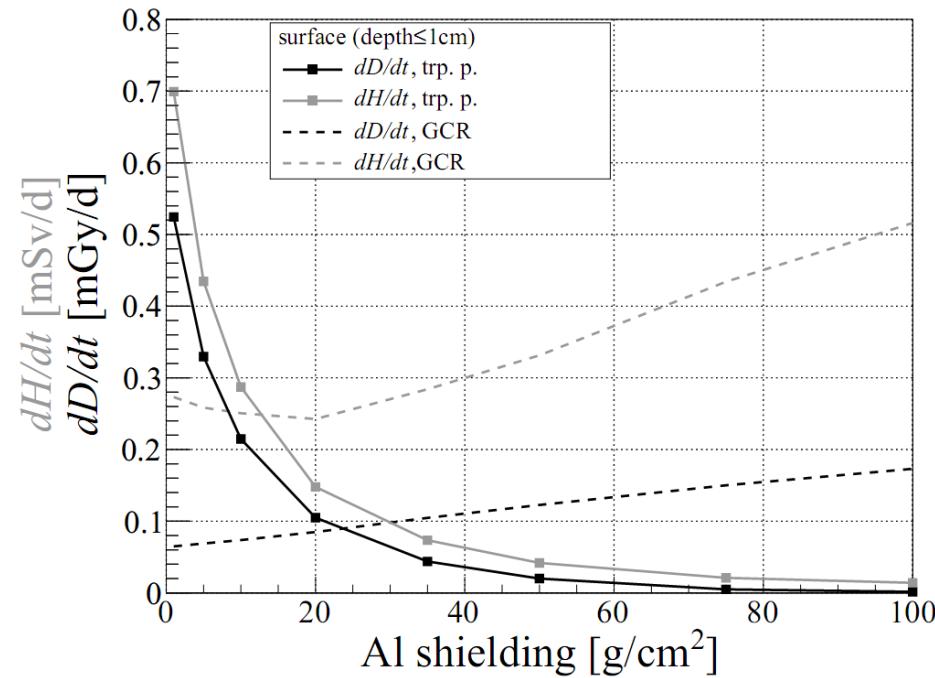
- Dose rates in a water sphere with radius 20 cm
 - whole sphere
 - surface (depth < 1cm)
- Spherical aluminium shielding
- GCR: $1 \leq Z \leq 26$

Matthiä et al., ASR 52 (2013)



GCR and trapped protons radiation exposure Dec. 2009 (max. GCR intensity), AP8min, 350 km

- Dose rates in a water sphere with radius 20 cm
 - surface (depth < 1cm)
- Spherical aluminium shielding
- GCR: $1 \leq Z \leq 26$
- Trapped protons: AP8min, 350km, ISS orbit



Solar minimum, Dez. 2009 (17th WRMISI)

Ap8-min 400 km, 25 g/cm² Aluminum

Absorbed Dose

GEANT4

Trapped protons:

$$100 \mu\text{Gy/d} < dD/dt < 160 \mu\text{Gy/d}$$

GCR:

$$85 \mu\text{Gy/d} < dD/dt < 105 \mu\text{Gy/d}$$

DOSTEL1/2, DOSIS 3D

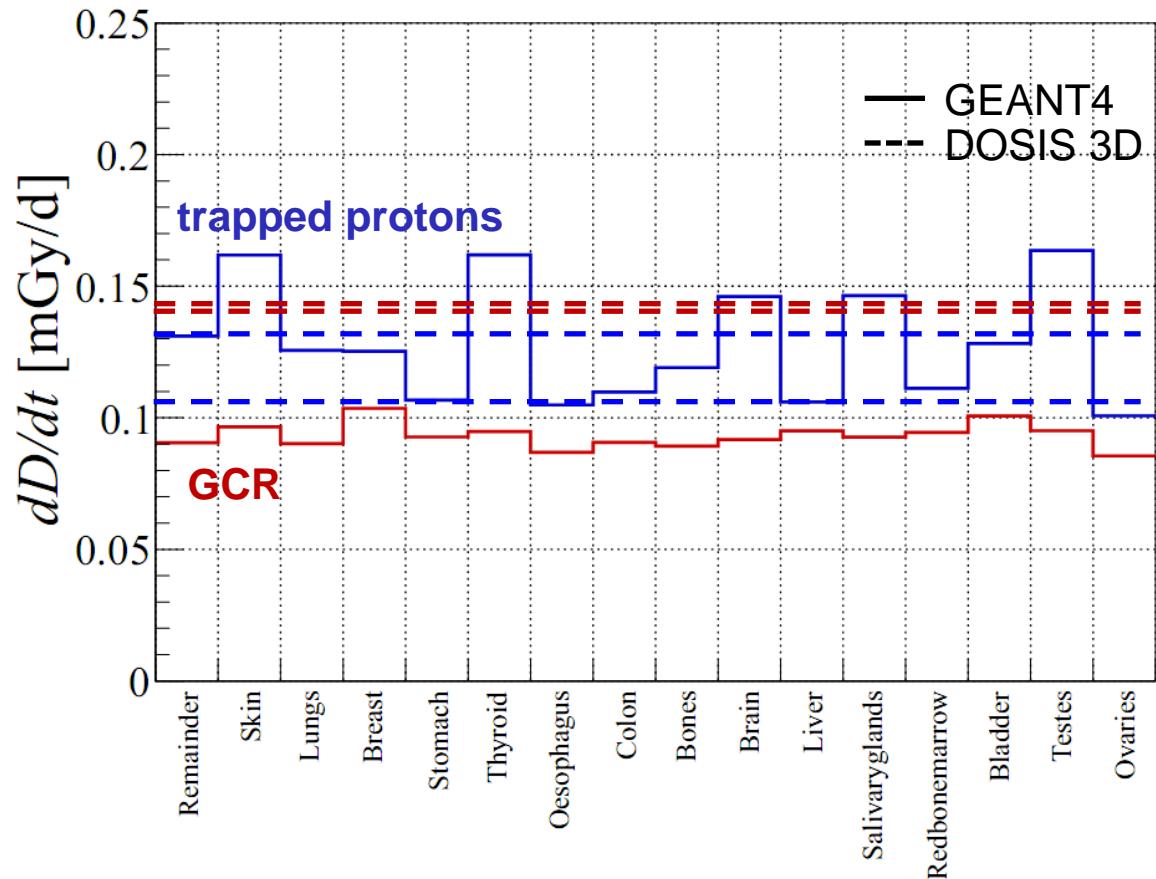
(S. Burmeister):

SAA:

$$dD/dt = 131 / 106 \mu\text{Gy/d}$$

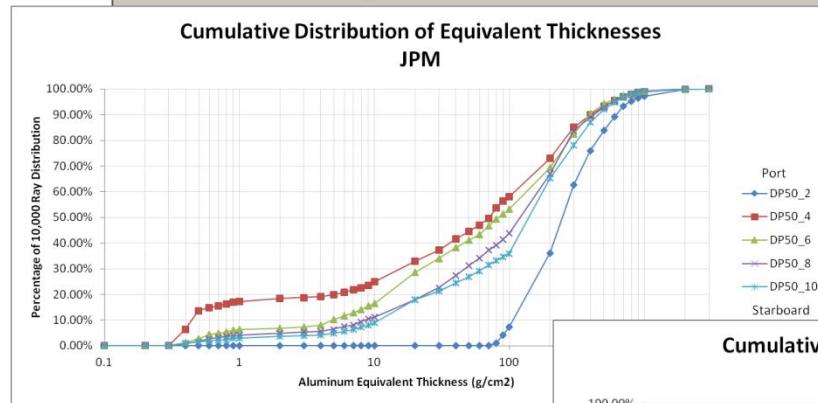
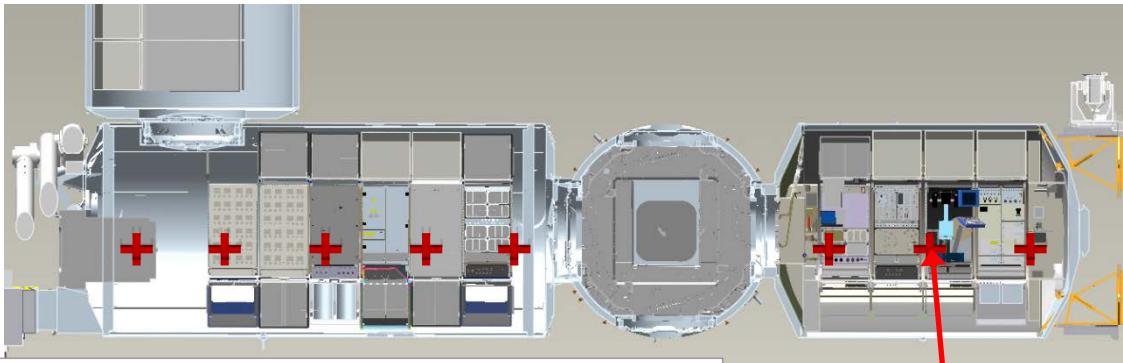
GCR:

$$dD/dt = 146 / 143 \mu\text{Gy/d}$$

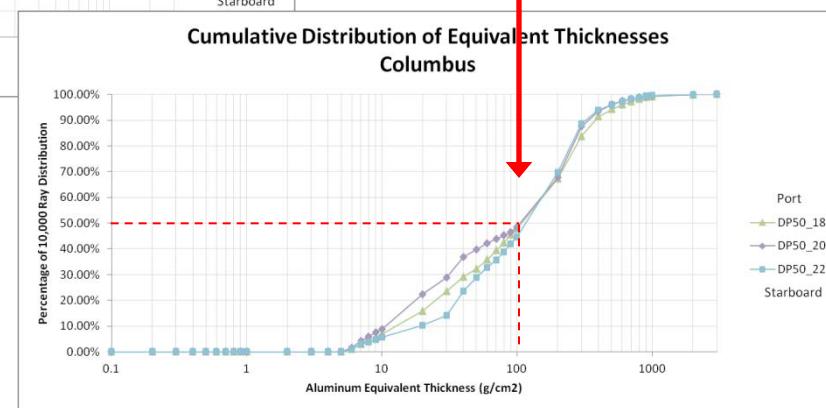


Columbus shielding distribution

N. Stoffle et al., 17th WRMIS (2012)

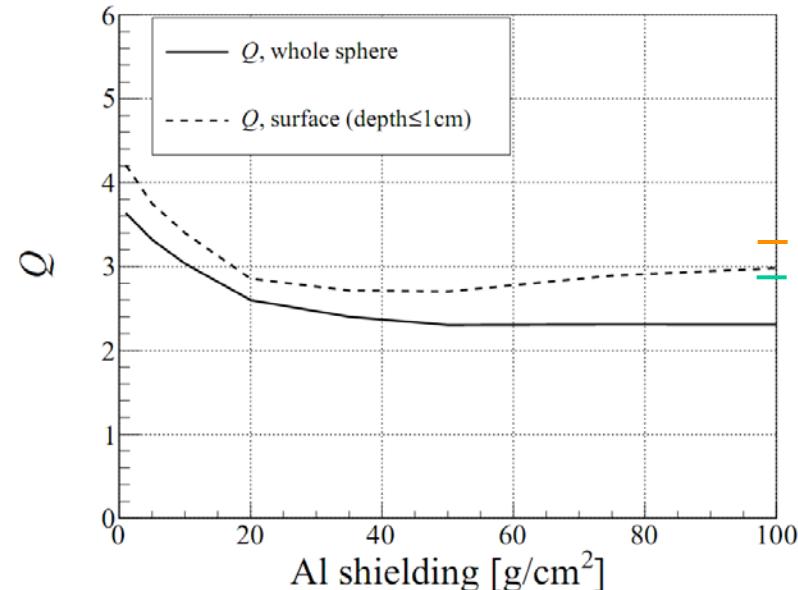
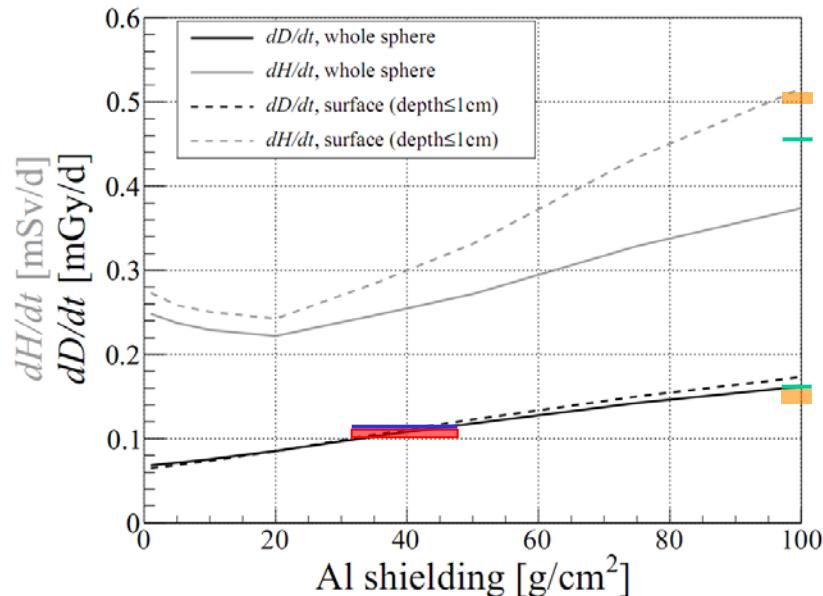


COLUMBUS
median shielding $\approx 100 \text{ g/cm}^2$



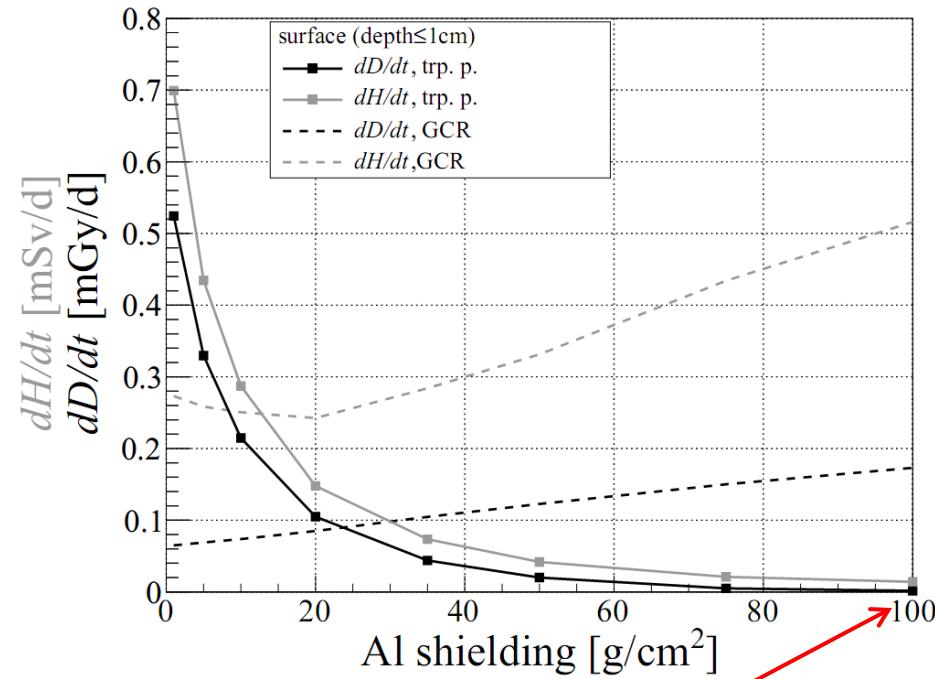
GCR radiation exposure on ISS Dec. 2009 (max. GCR intensity)

- Dose rates in a water sphere with radius 20 cm
 - surface (depth < 1cm)
 - whole sphere
- Spherical aluminium shielding
- Columbus Module:
 - $\approx 100 \text{ g/cm}^2$, Stoffle et al. (2012)
 - $158 \mu\text{Gy/d}$, $448 \mu\text{Sv/d}$, $Q = 2.84$, Semones et al. (2009)
 - $150 - 157 \mu\text{Gy/d}$, $496 - 517 \mu\text{Sv/d}$, $Q = 3.30$, Burmeister et al. (2012)
- Zvezda Service Module:
 - $32 - 47 \text{ g/cm}^2$, Jardnickova et al. (2009)
 - $100 - 110 \mu\text{Gy/d}$, Lishnevskii et al. (2012)
 - $125 \mu\text{Gy/d}$, Semones et al. (2009)



GCR and trapped protons radiation exposure Dec. 2009 (max. GCR intensity), AP8min, 350 km

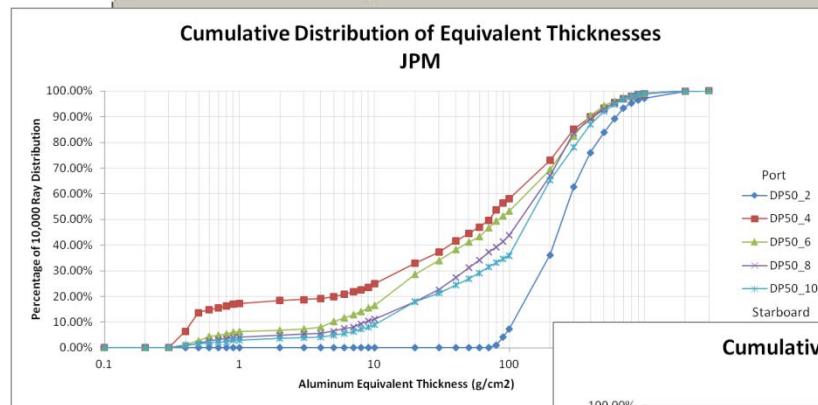
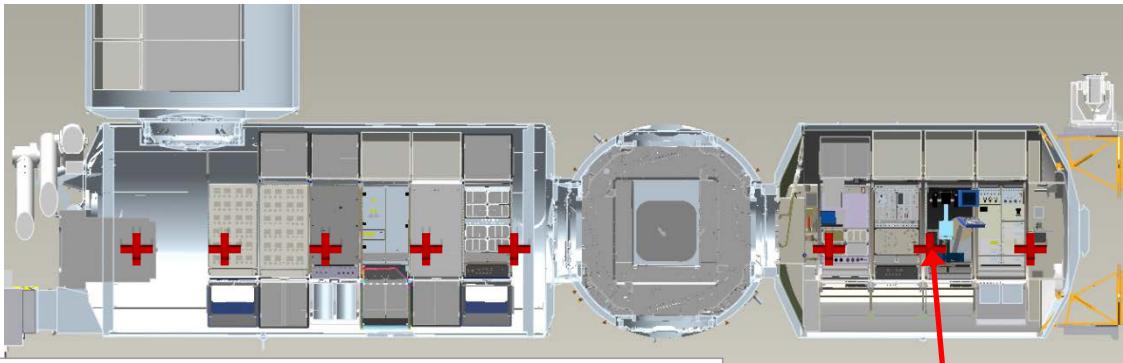
- Dose rates in a water sphere with radius 20 cm
 - surface (depth < 1cm)
- Spherical aluminium shielding
- GCR: $1 \leq Z \leq 26$
- Trapped protons: AP8min, 350km, ISS orbit



100 g/cm² Al: dose rate ≈ 0
but e.g. DOSIS: 70 – 80 µGy/d, 100 – 110 µSv/d !

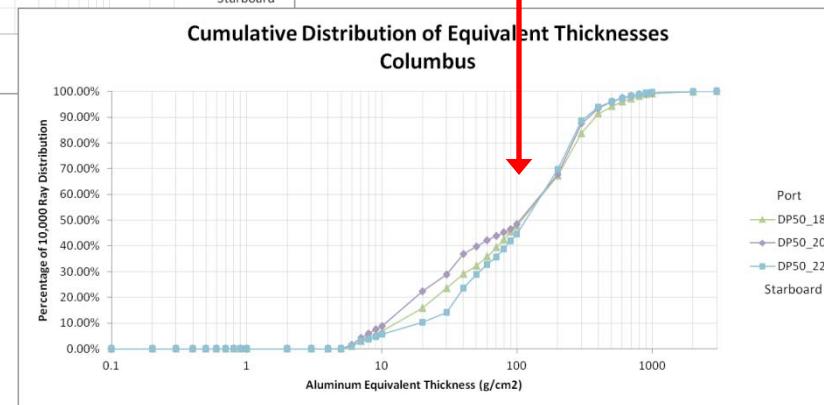
Columbus shielding distribution

N. Stoffle et al., 17th WRMIS (2012)



COLUMBUS

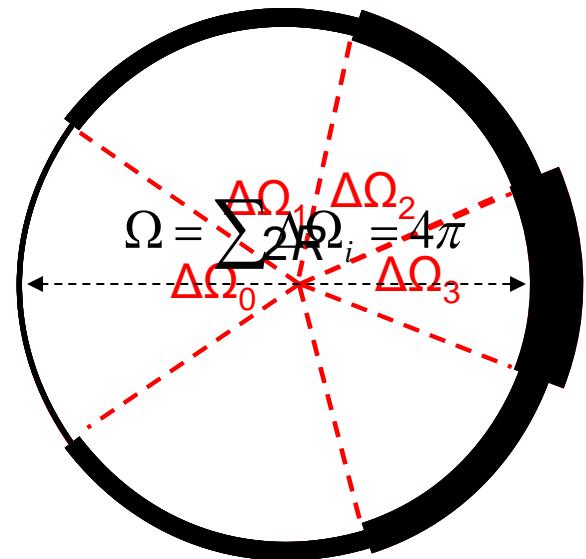
~~median shielding $\approx 100 \text{ g}/\text{cm}^2$~~
→ cumulative distribution function



Construction of shielding geometry from cumulative distribution function CDF

S [g/cm ²]	CDF [%]	$\Delta\Omega$
1	20	$0.2 \cdot 4\pi$
3	60	$0.4 \cdot 4\pi$
5	90	$0.3 \cdot 4\pi$
10	100	$0.1 \cdot 4\pi$

sphere with radius R



Radiation exposure on ISS, **trapped protons** + **GCR**

Dec. 2009 (max. GCR intensity)

target sphere: 5 cm radius

Target	trapped protons, AP8 min			GCR		
	0 g/cm ² – 0.5 g/cm ²			0 g/cm ² – 0.5 g/cm ²		
	dD/dt [μGy/d]	dH/dt [μSv/d]	Q	dD/dt [μGy/d]	dH/dt [μSv/d]	Q
5 cm H ₂ O sphere	75	111	1.48	140	480	3.43
Experiment:						
DOSIS DOSTEL 1*	80	112	1.40	157	517	3.30
DOSIS DOSTEL 2*	68	96	1.43	150	496	3.30
NASA TEPC §	165	293	1.78	158	448	2.84

*Burmeister et al., 17th WRMISI (2012)

§Semones et al., 14th WRMISI (2009)



Radiation exposure on ISS, **trapped protons** + **GCR**

Dec. 2009 (max. GCR intensity)

for organ dose calculation, target sphere: 20 cm radius

	trapped protons, AP8 min			GCR		
	0 g/cm ² – 0.5 g/cm ²			0 g/cm ² – 0.5 g/cm ²		
Target	dD/dt [μGy/d]	dH/dt [μSv/d]	Q	dD/dt [μGy/d]	dH/dt [μSv/d]	Q
20 cm H ₂ O sphere	55	83	1.50	142	435	3.06
Experiment:						
DOSIS DOSTEL 1*	80	112	1.40	157	517	3.30
DOSIS DOSTEL 2*	68	96	1.43	150	496	3.30
NASA TEPC §	165	293	1.78	158	448	2.84

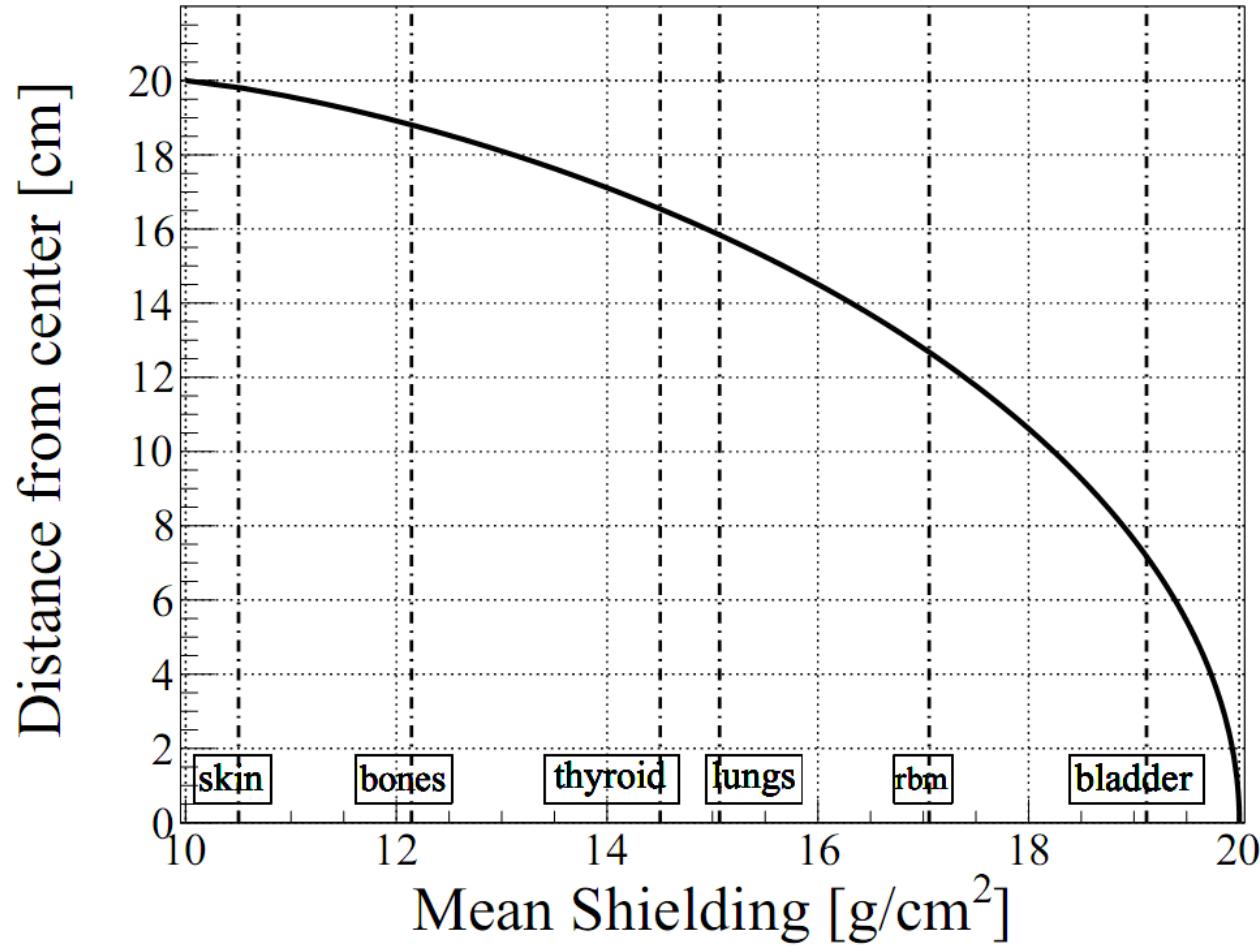
*Burmeister et al., 17th WRMISI (2012)

§Semones et al., 14th WRMISI (2009)

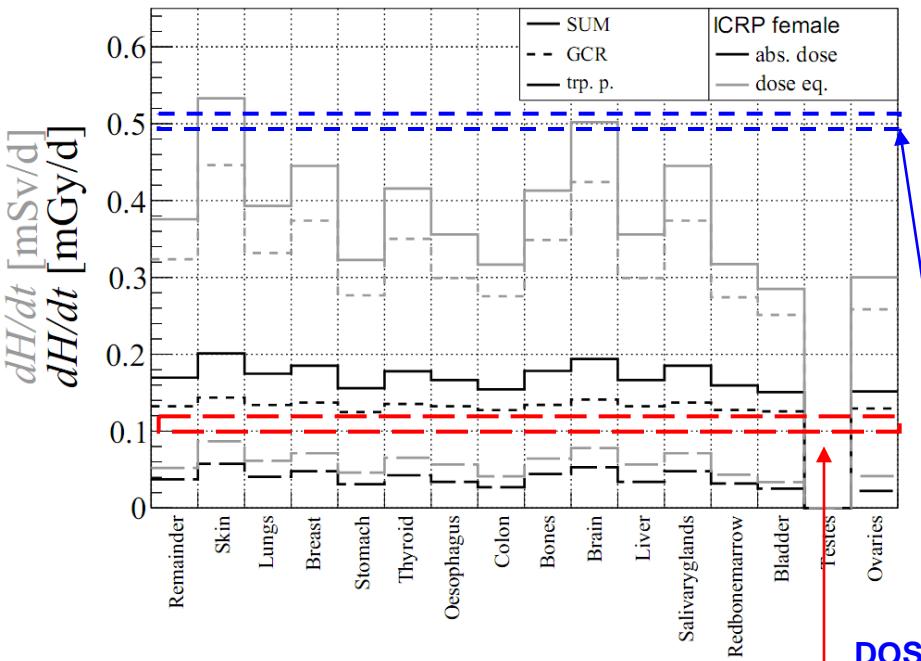


Mean shielding in a water sphere with radius 20 cm

Isotropic irradiation – Positions in the sphere corresponding to average shielding of selected organs of the male ICRP phantom



Radiation exposure on ISS, trapped protons + GCR Dec. 2009 (max. GCR intensity)

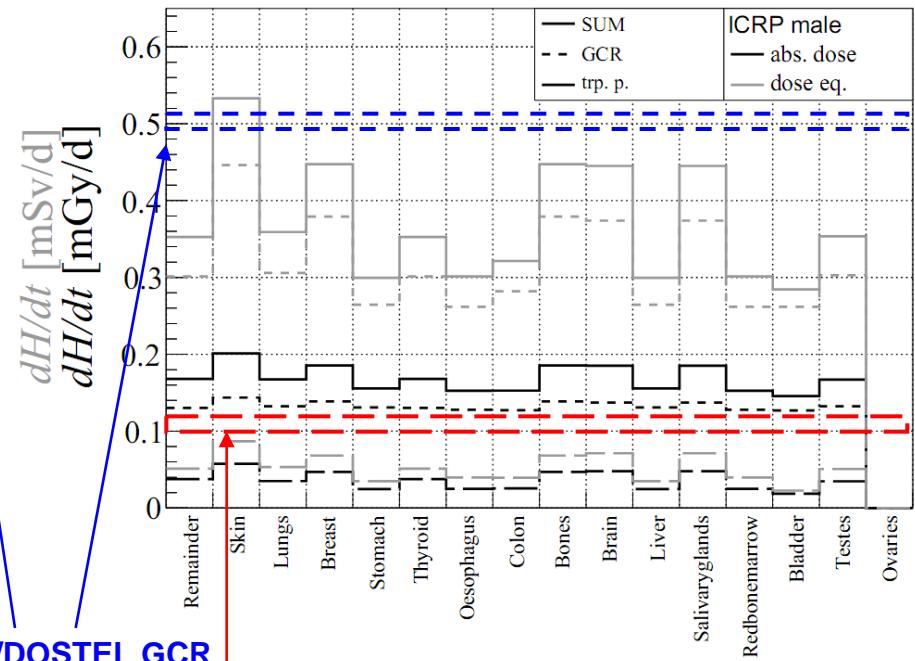


DOSIS/DOSTEL GCR

DOSIS/DOSTEL SAA

Male/Female:

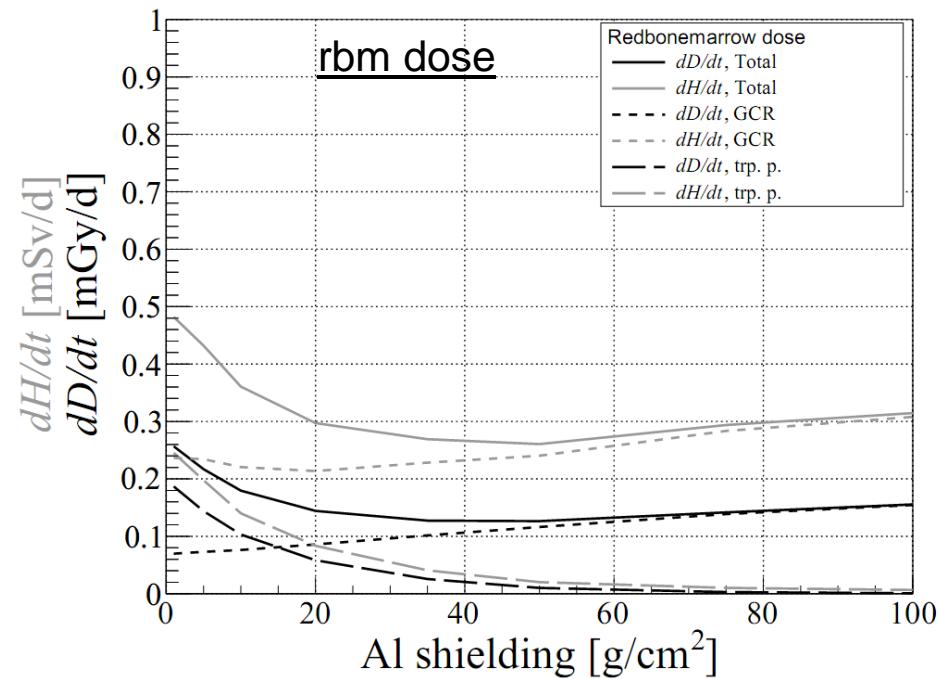
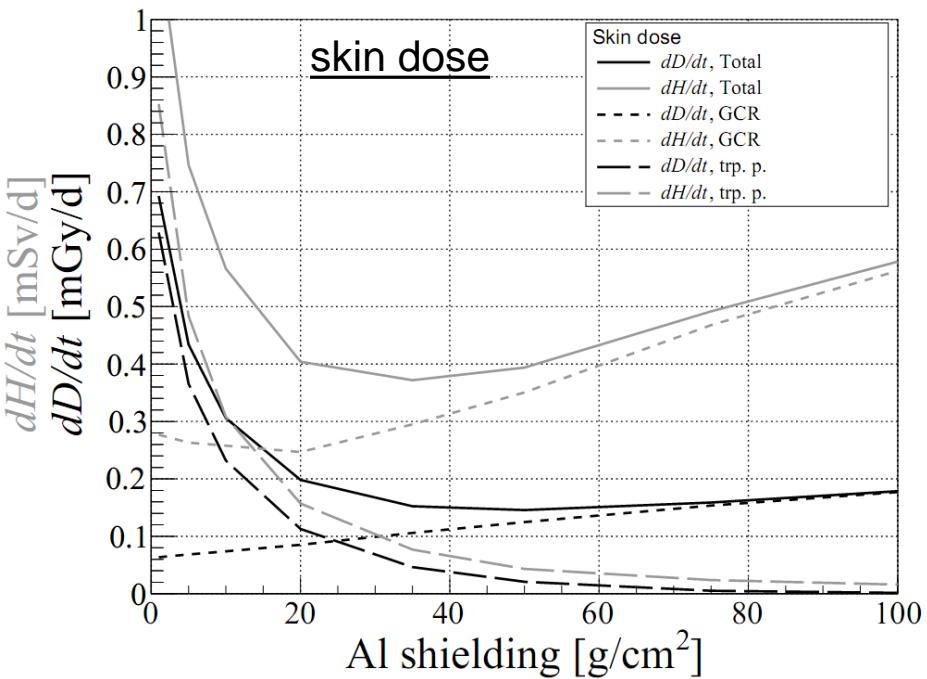
0.28 mSv/d (bladder) – 0.54 mSv/d (skin)



Radiation exposure on ISS, trapped protons + GCR

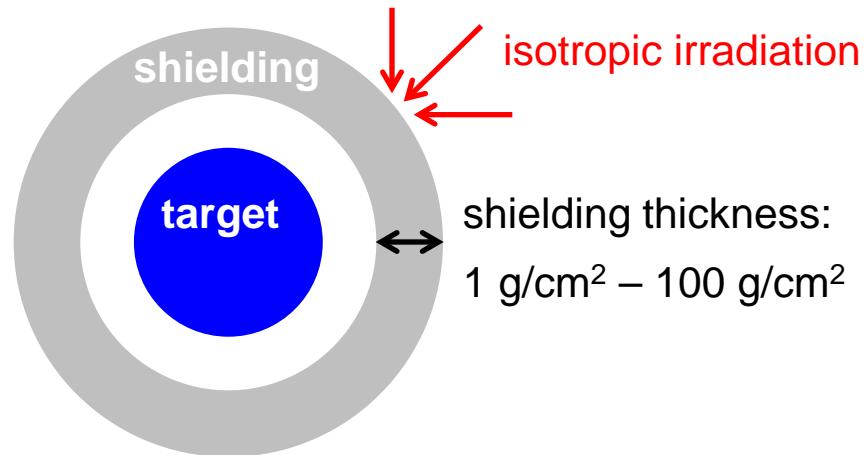
Dec. 2009 (max. GCR intensity)

Organ dose vs. shielding



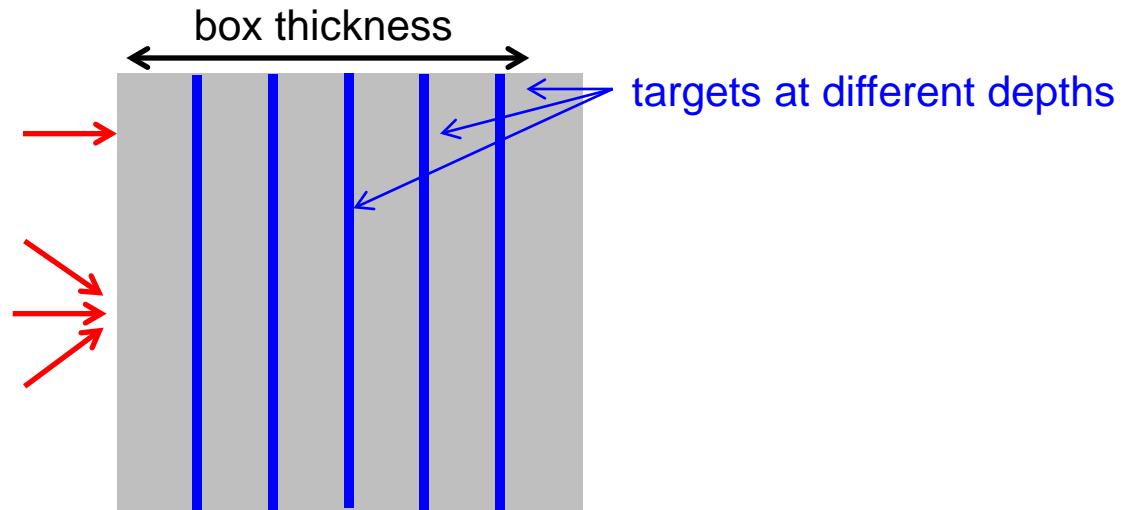
Comparison of different shielding geometries

- Spherical shielding



- Planar shielding:

a) pencil beam



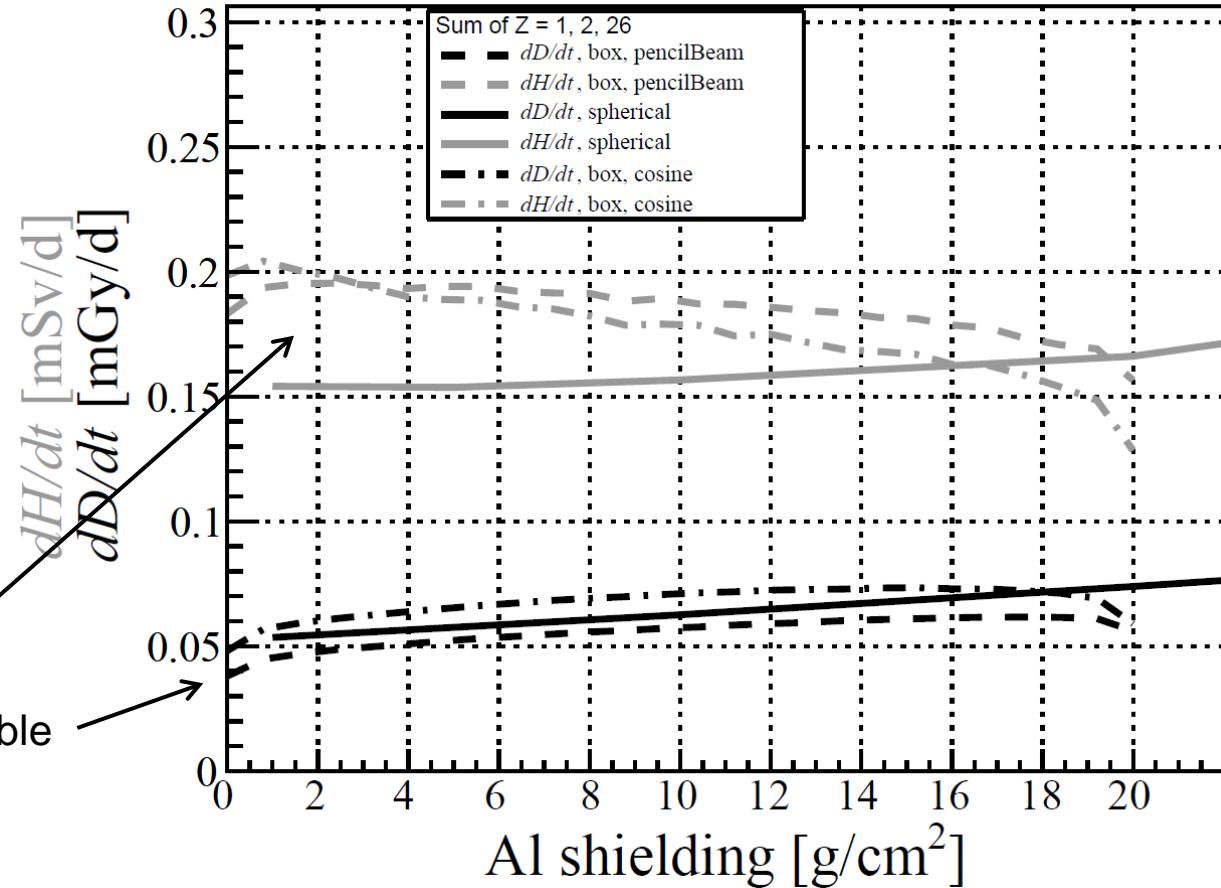
b) cosine

Comparison of different shielding geometries

Box of 20 g/cm² Aluminium

- Solar minimum
- Z=1, 2, 26
- Box, 20 g/cm² Al, 7.4 cm
- Different spherical shielding

quality factor differs
absorbed dose comparable

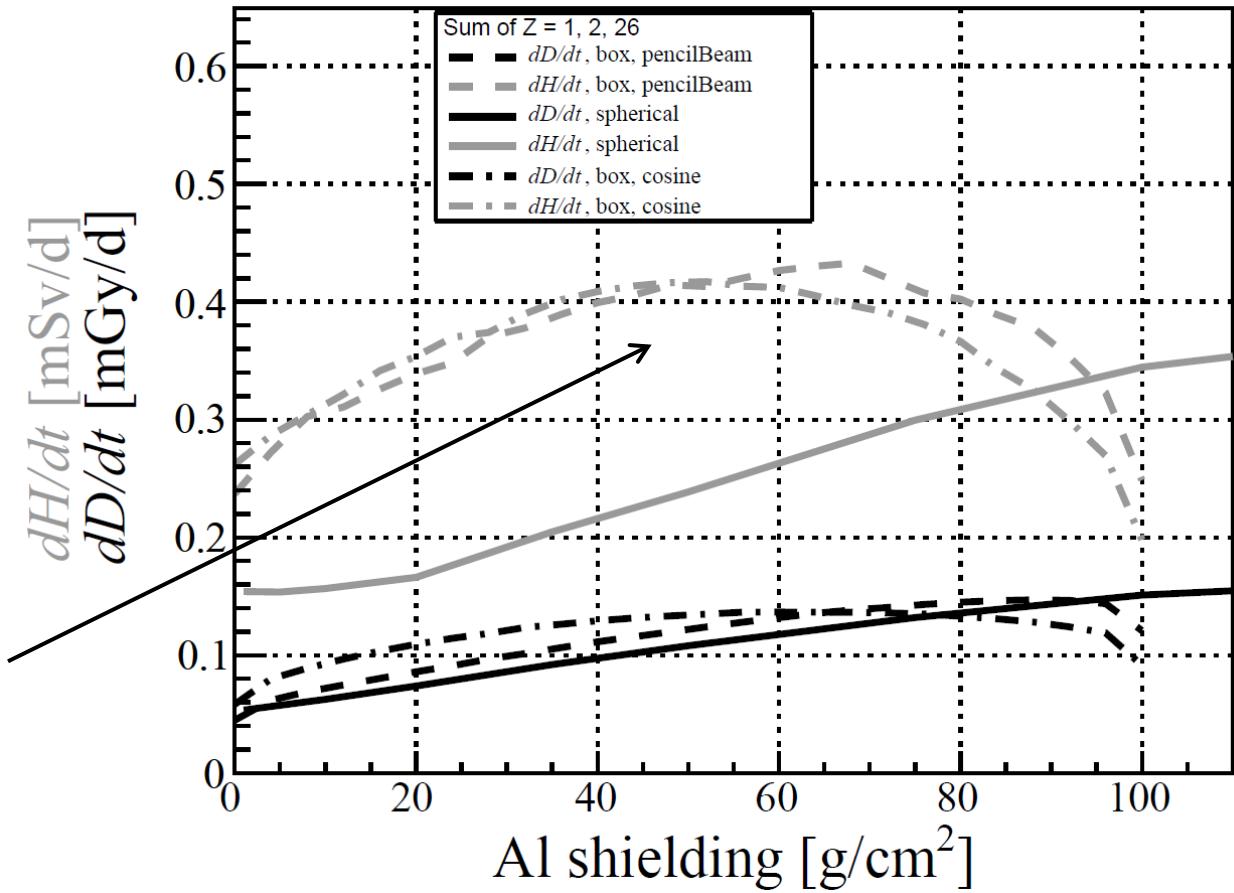


Comparison of different shielding geometries

Box of 100 g/cm² Aluminium

- Solar minimum
- Z=1, 2, 26
- Box, 100 g/cm² Al, 37,2 cm
- Different spherical shielding

secondary particle build-up
backscattering

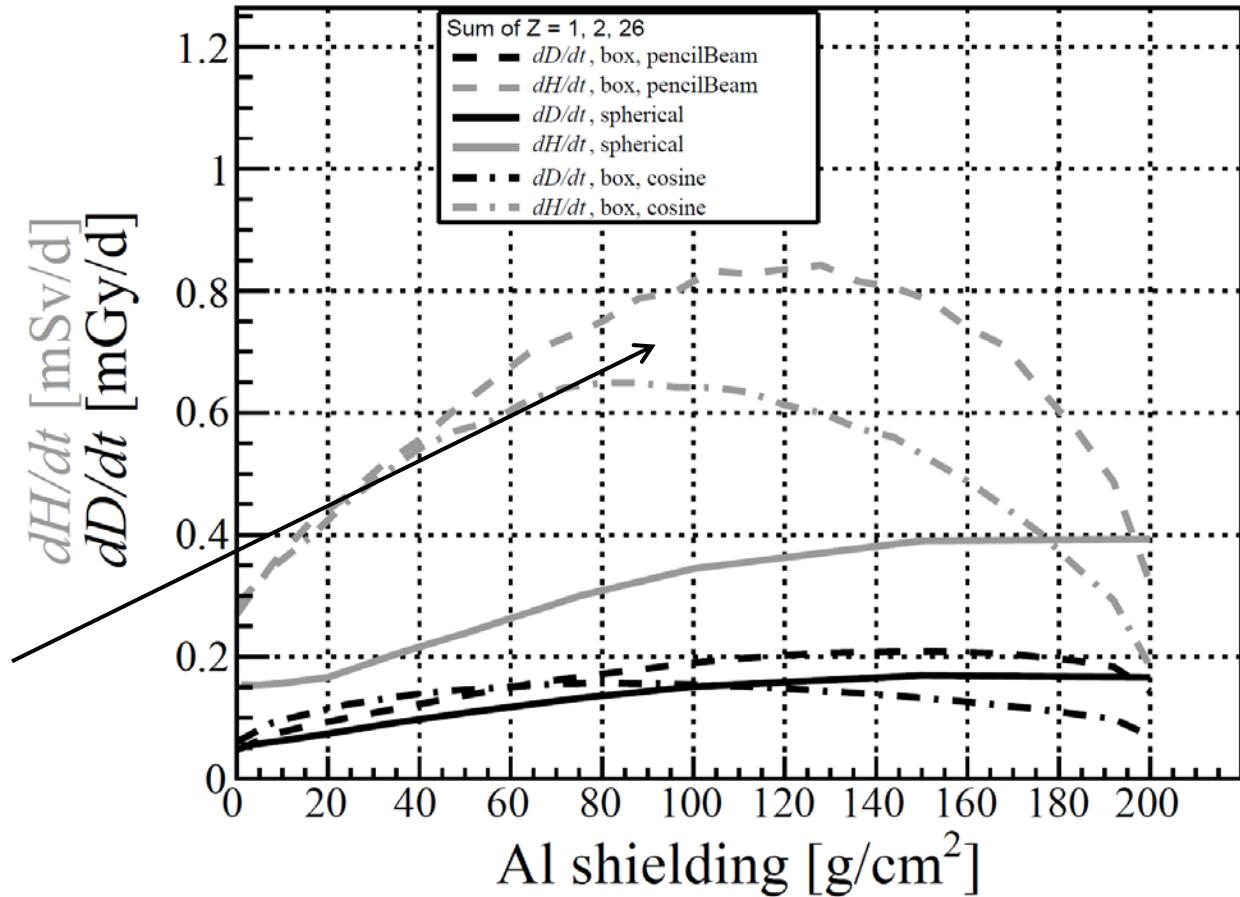


Comparison of different shielding geometries

Box of 200 g/cm² Aluminium

- Solar minimum
- Z=1, 2, 26
- Box, 200 g/cm² Al, 74,3 cm
- Different spherical shielding

secondary particle build-up
backscattering



Summary

- Dose estimates are strongly dependent on GCR model:
 - Solar minimum:
 - 35% differences (interplanetary space)
 - 30% differences (LEO)
- With DLR model: good agreement with measurements for 2009
- Organ dose equivalent rates in Columbus (ICRP male/female phantoms):
0.28 mSv/d (bladder) – 0.54 mSv/d (skin) (*meas.* $0.45^{\$}$ – 0.52^{*} mSv/d)
- Spherical shielding geometry (LEO, GCR + trp. p.):
 - Calculated organ dose rates decrease up to 30 – 50 g/cm², increase above
- Different shielding geometries not comparable

Thank you for your attention!



*Burmeister et al., 17th WRMISST (2012)
\\$Semones et al., 14th WRMISST (2009)

