

The Columbus Radiation Environment & Effects Package (CREEP)

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CREEP Objectives

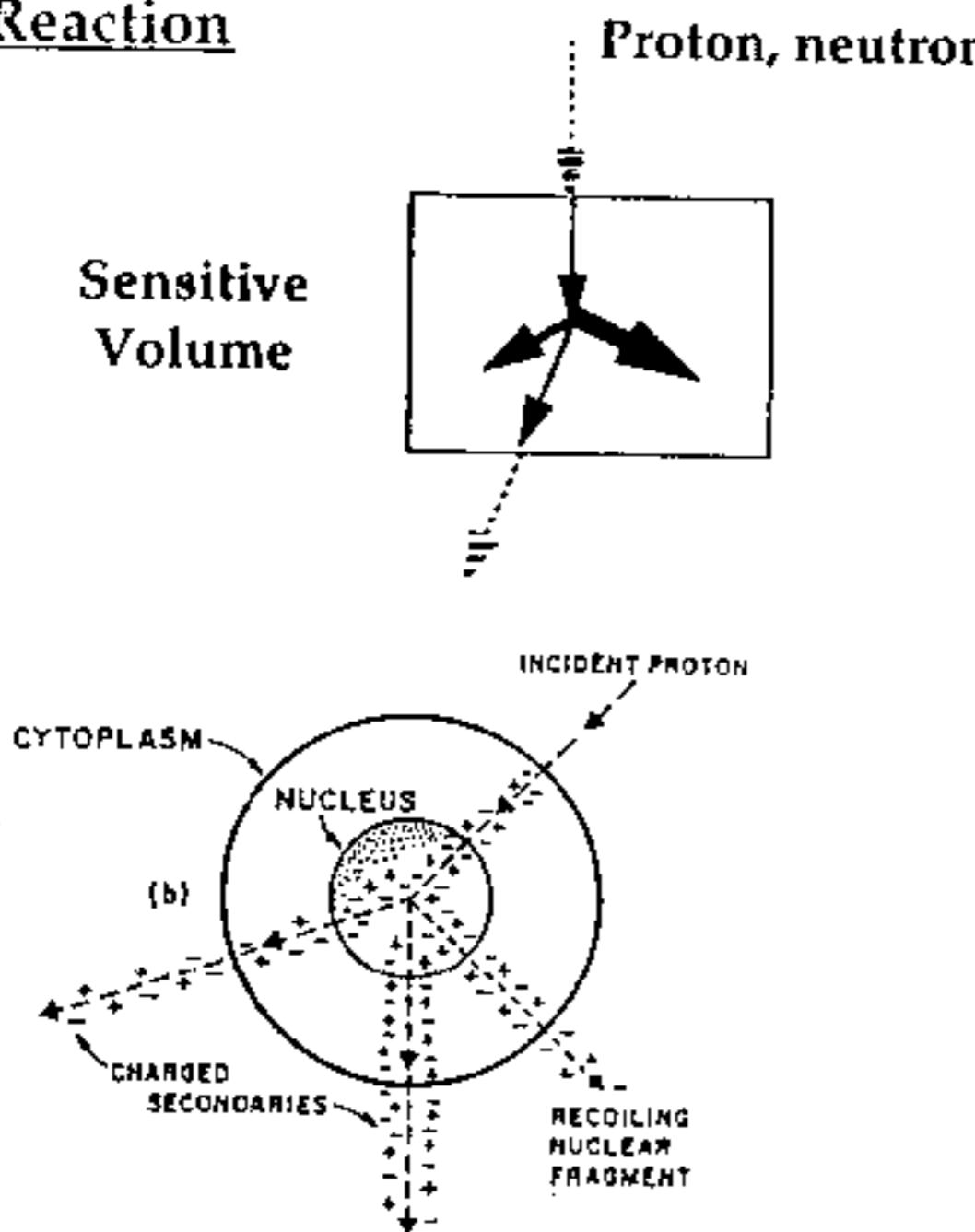
- characterize environments and effects of concern for:
 - other **payloads** (e.g. background, sensor degradation)
 - **systems** (e.g. SEU, latch-up, damage)
 - **long-term utilisation** (well characterized environment with technology effects)
 - **crew** (long term database of *primary* quantities) space radiation problems will become increasingly important
- ISSA early-utilization will be at **solar maximum** (e.g. solar events)
- provide data for **modelling** (i) environment (ii) shielding
- provide comprehensive **effects data**, with ground **correlations**
- flight test **important space technologies** for wide application

Key problem areas

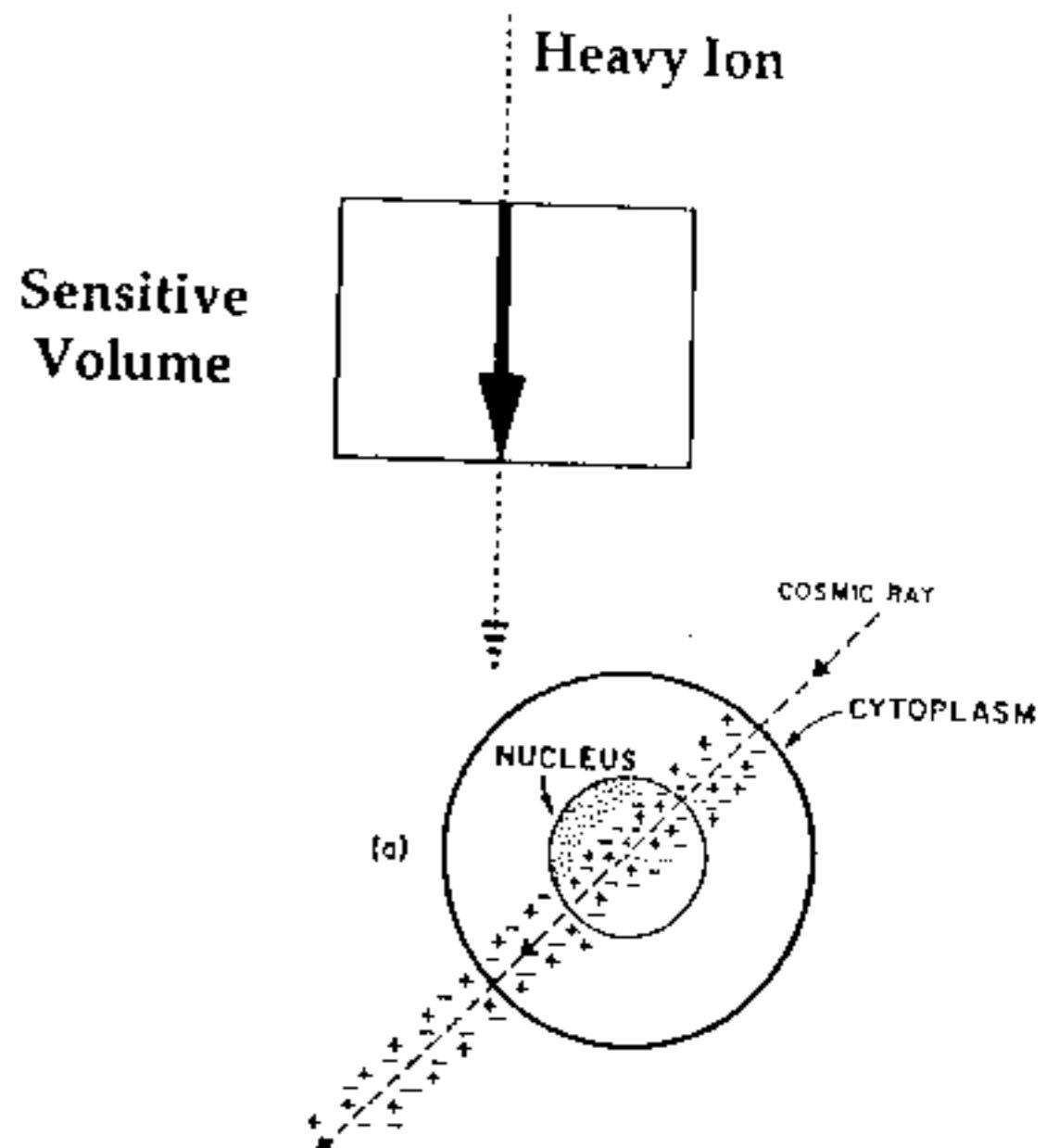
- ISSA radiation environment interacts with B-field and atmosphere:
time-dependent modulation; asymmetries;
- models currently poor;
- 52° orbit exposed to cosmic rays, “solar flares”;
- energetic proton effects: SEU; damage; background;
- radiobiological effects to crew;
- radio-activation
- shielding effects and secondary radiations
- “space weather” systems in context of station use and crew protection

SINGLE EVENT EFFECTS & RADIobiological EFFECTS

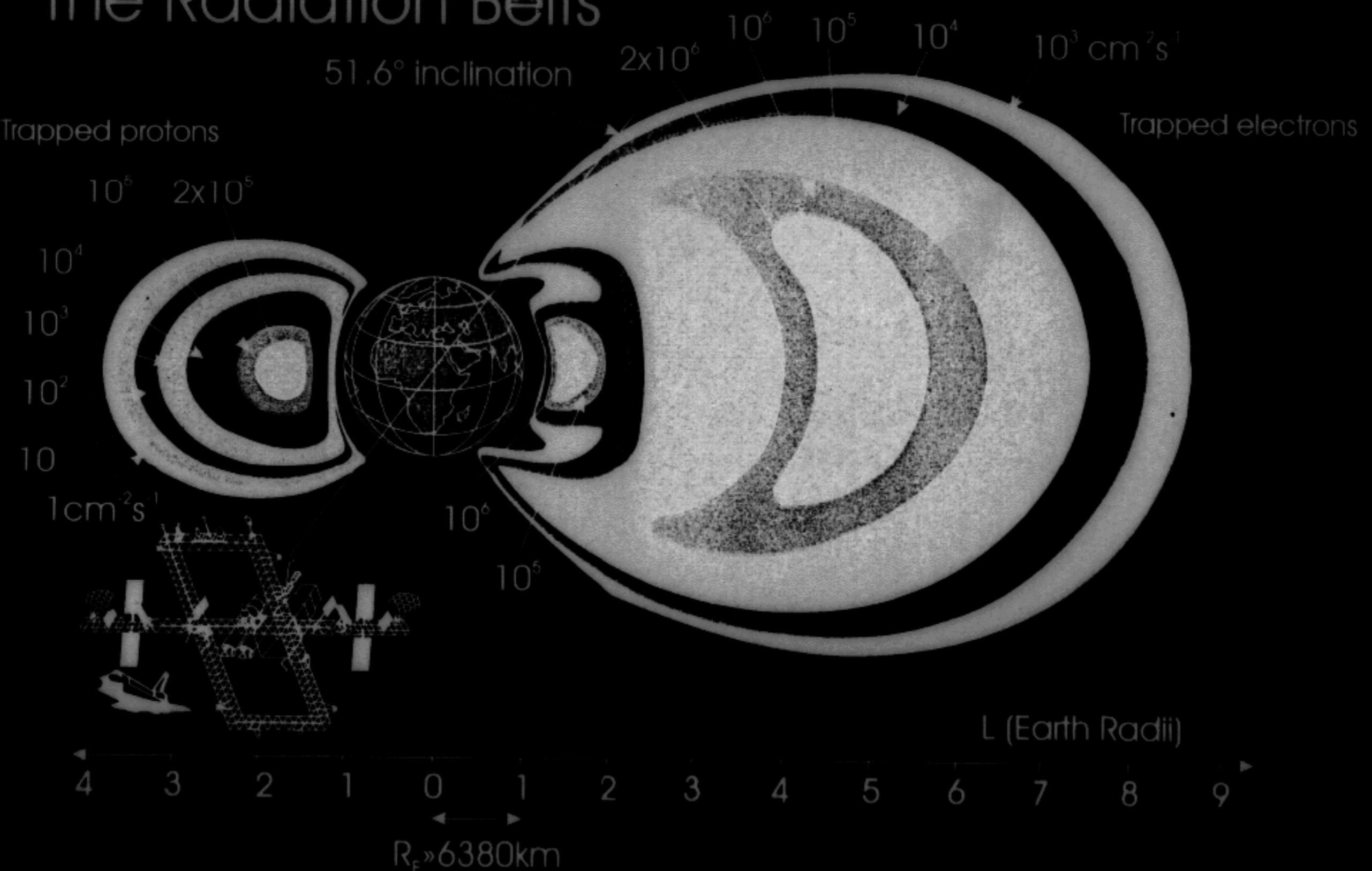
Nuclear Reaction



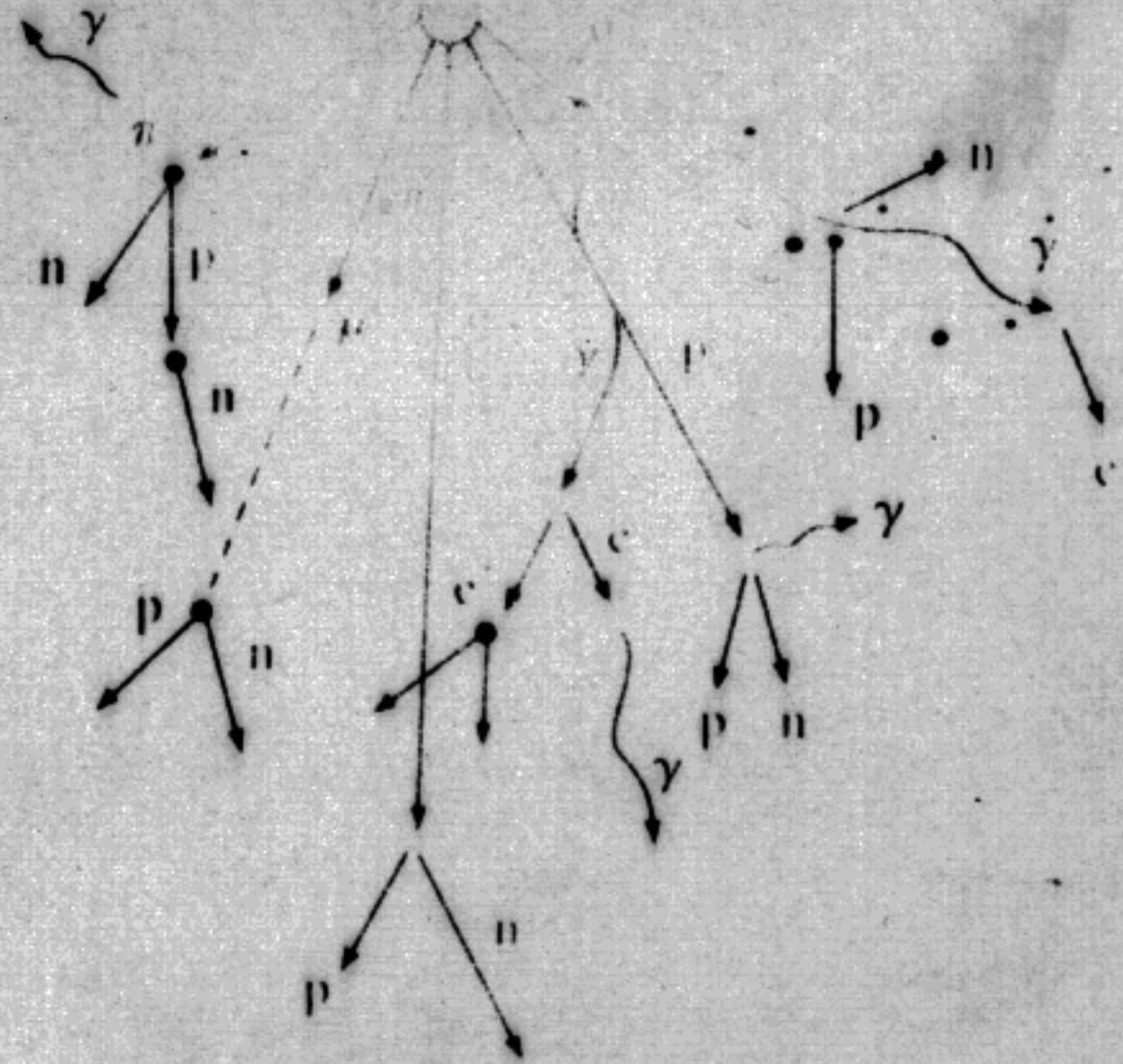
Direct Ionization



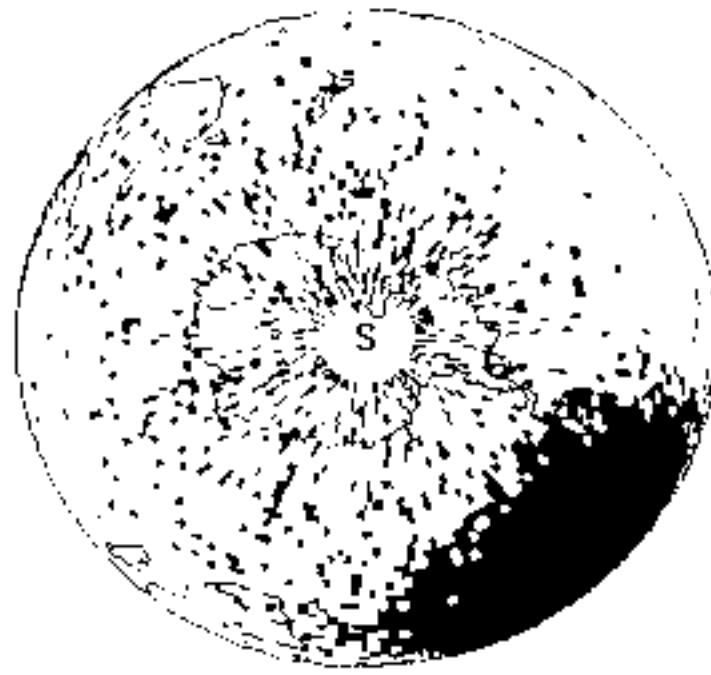
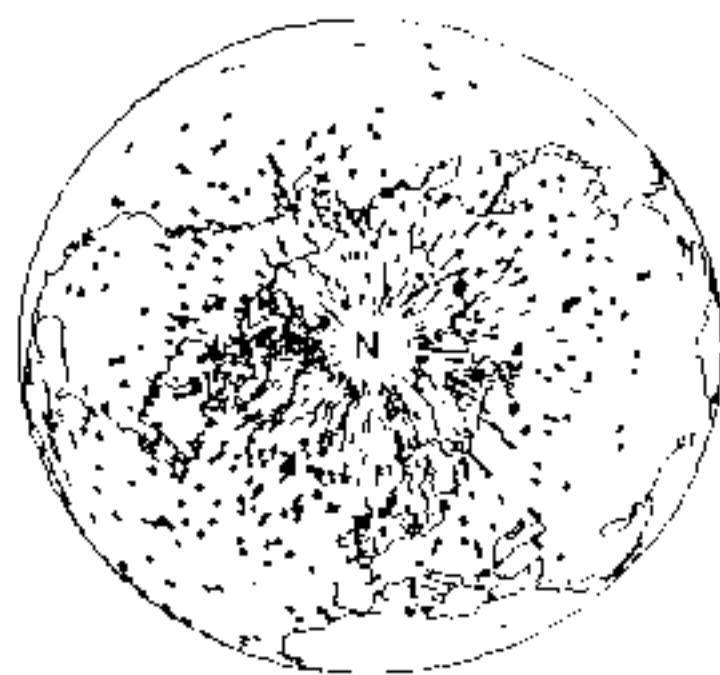
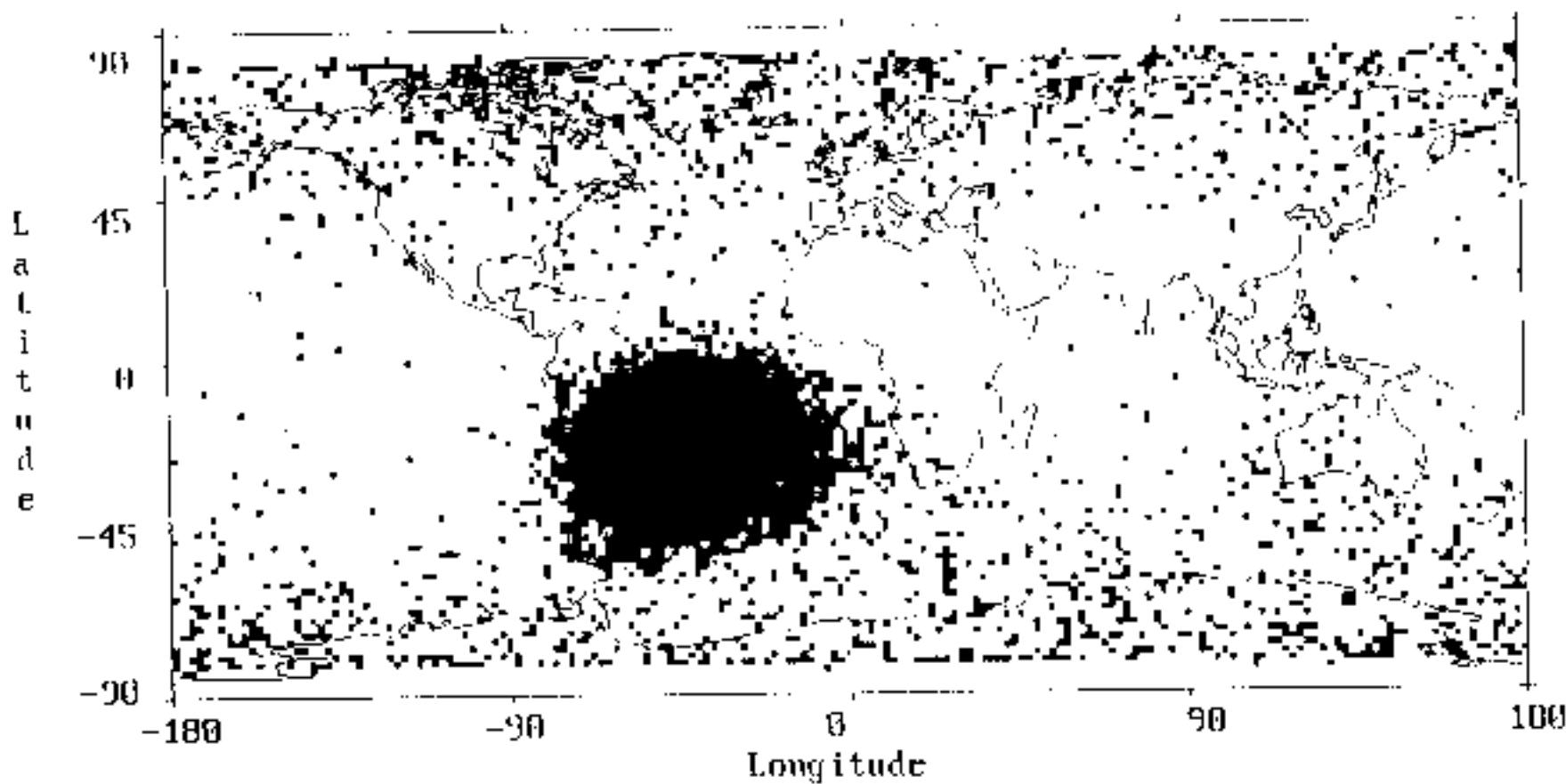
The Radiation Belts



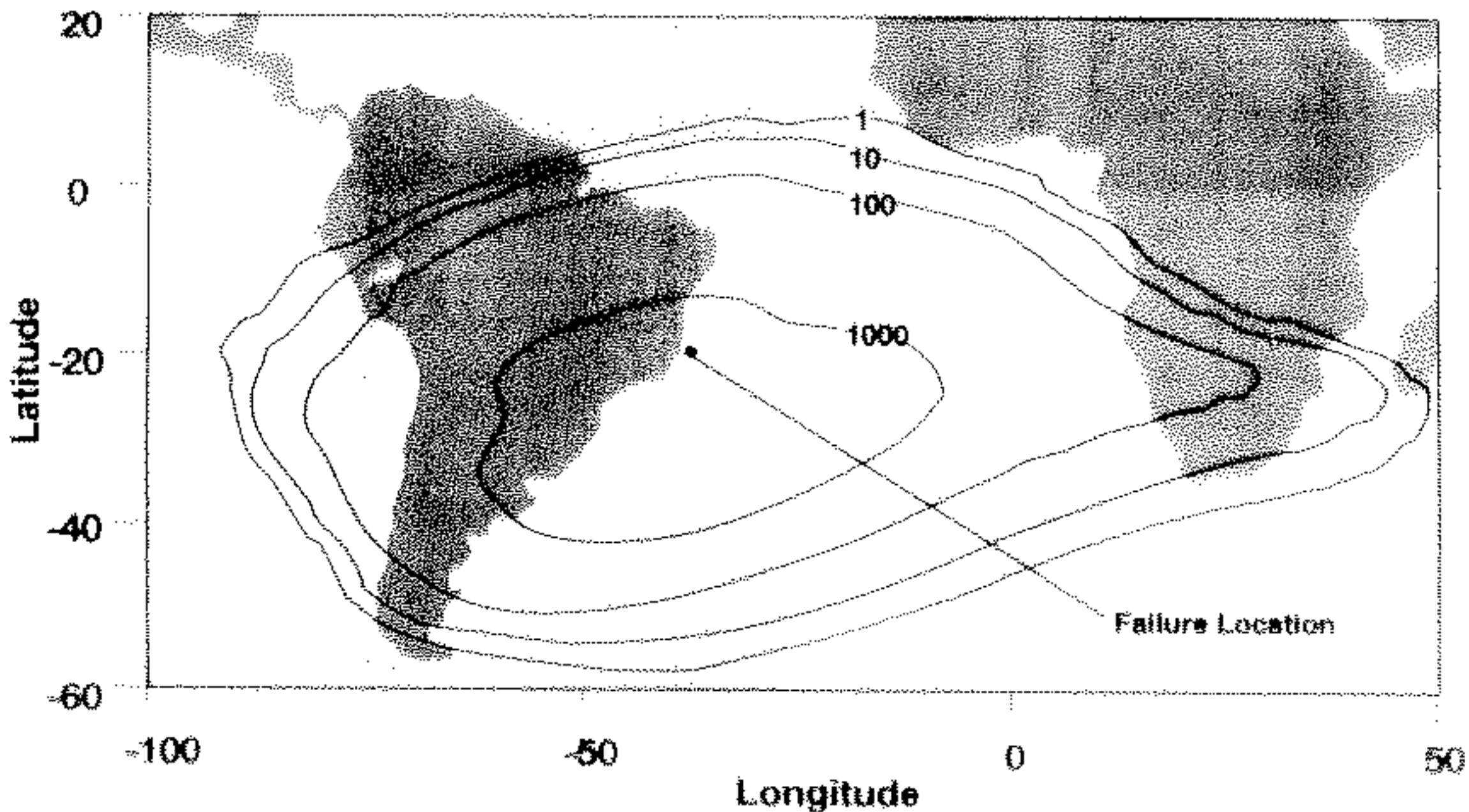
High inclination LEO missions intersect radiation in the South Atlantic Anomaly and the 'horns' of the outer belt



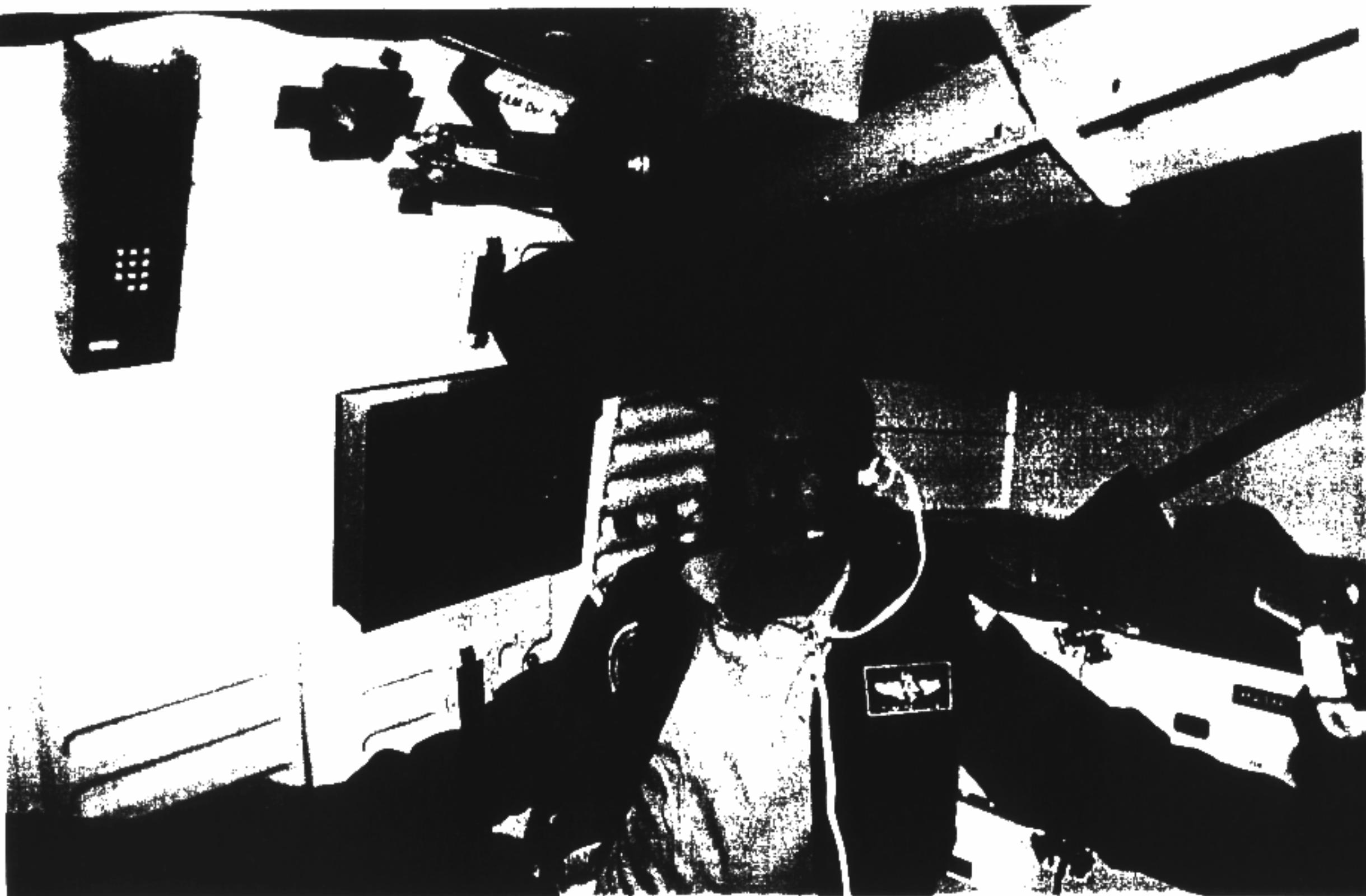
UoSAT 2 OBC MEMORY UPSETS : TEXAS TMS4416 16k x 4 BIT NMOS DRAM



7995 Events in 1364 Days
September '88 - May '92

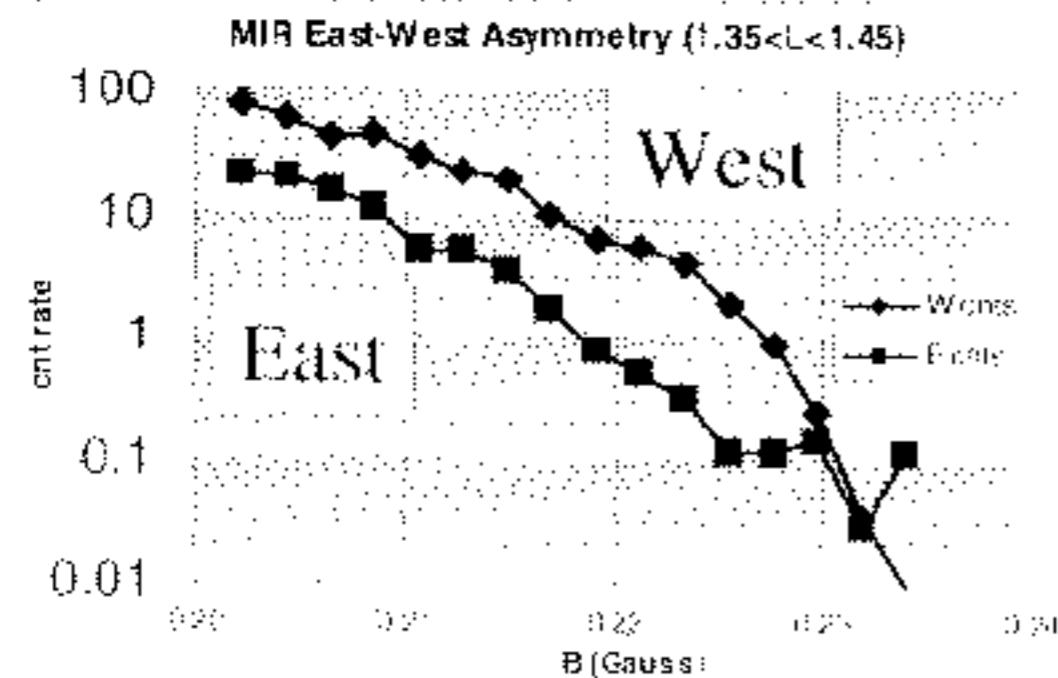
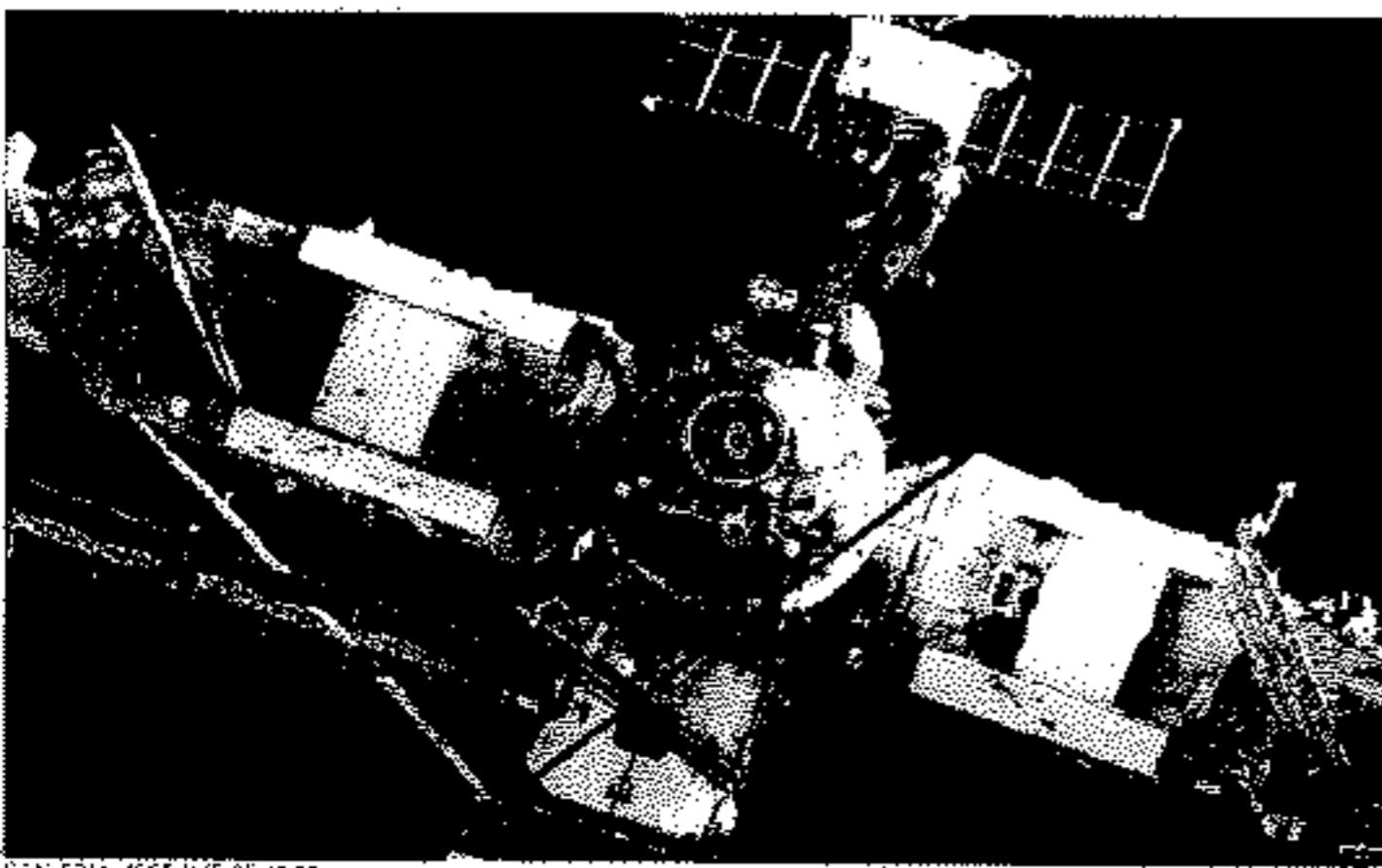


Failure Location in South Atlantic Anomaly showing
Proton contours (p/cm²/sec)



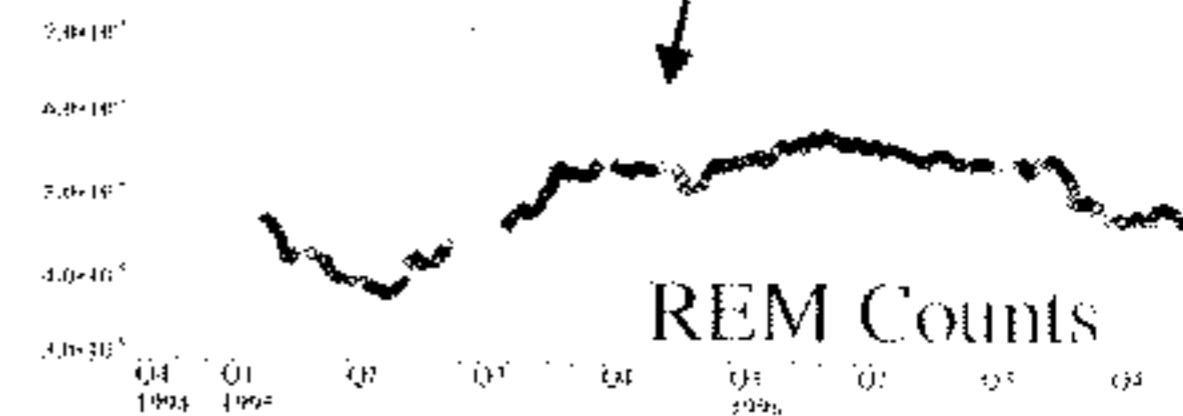
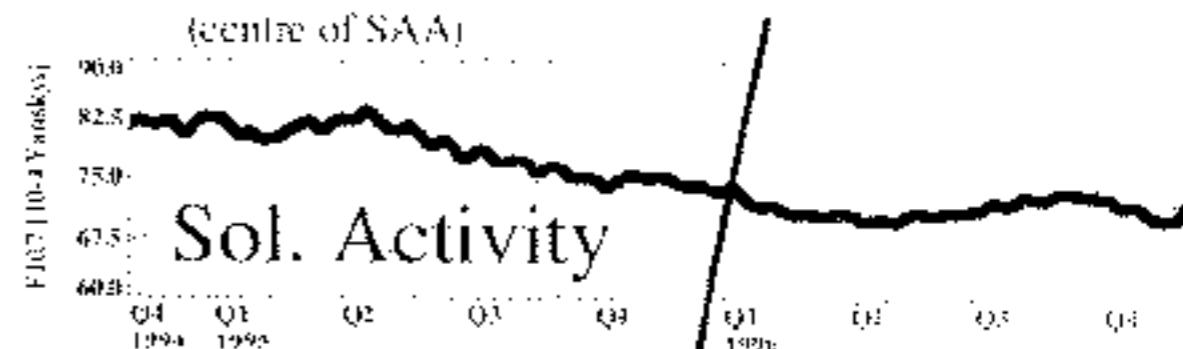


DERA



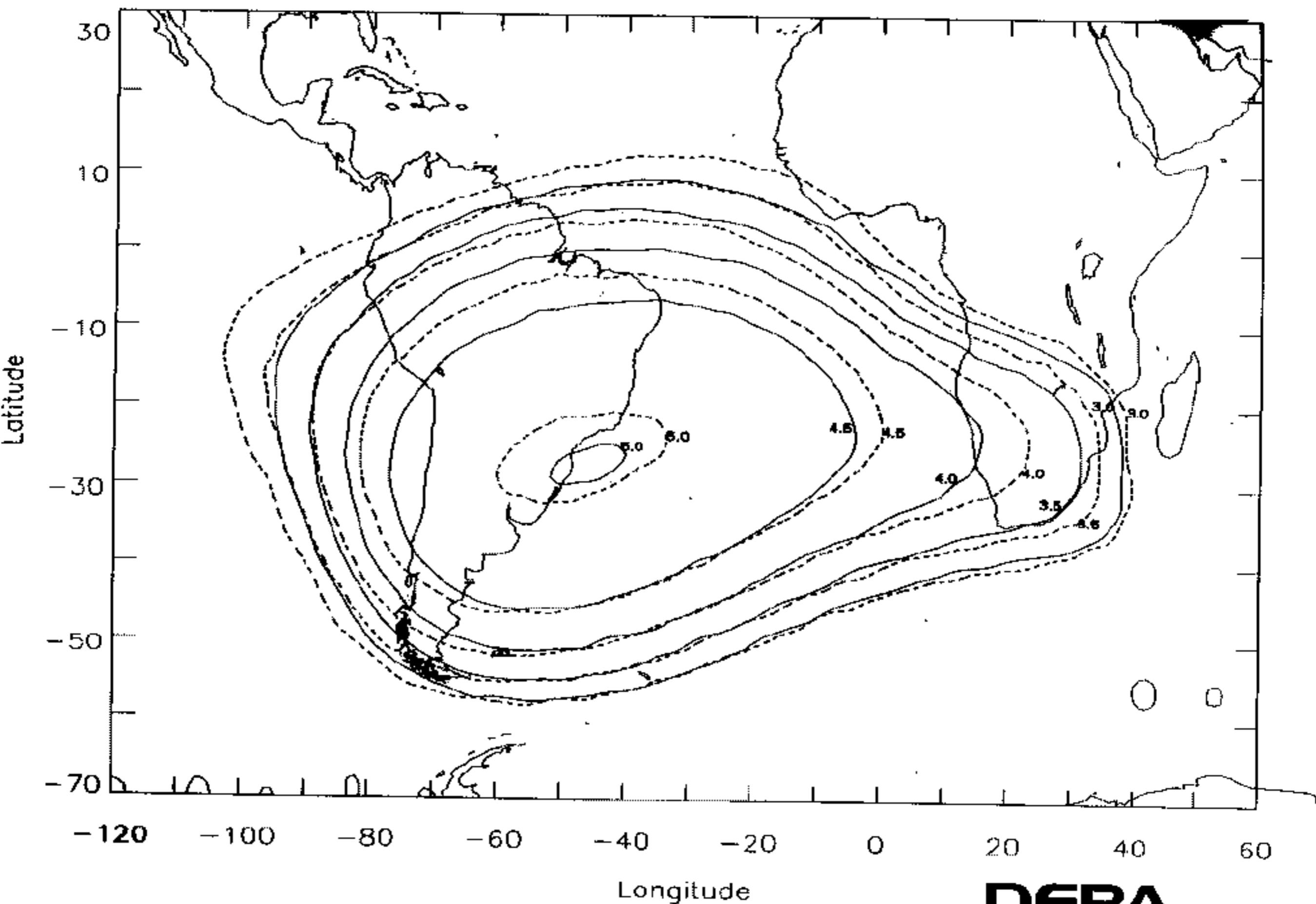
**REM on Outside of MIR: LEO,
400 km, 52° Nov. 1994 - Nov.
1996 (PSI/ESTEC)**

Response to atmosphere

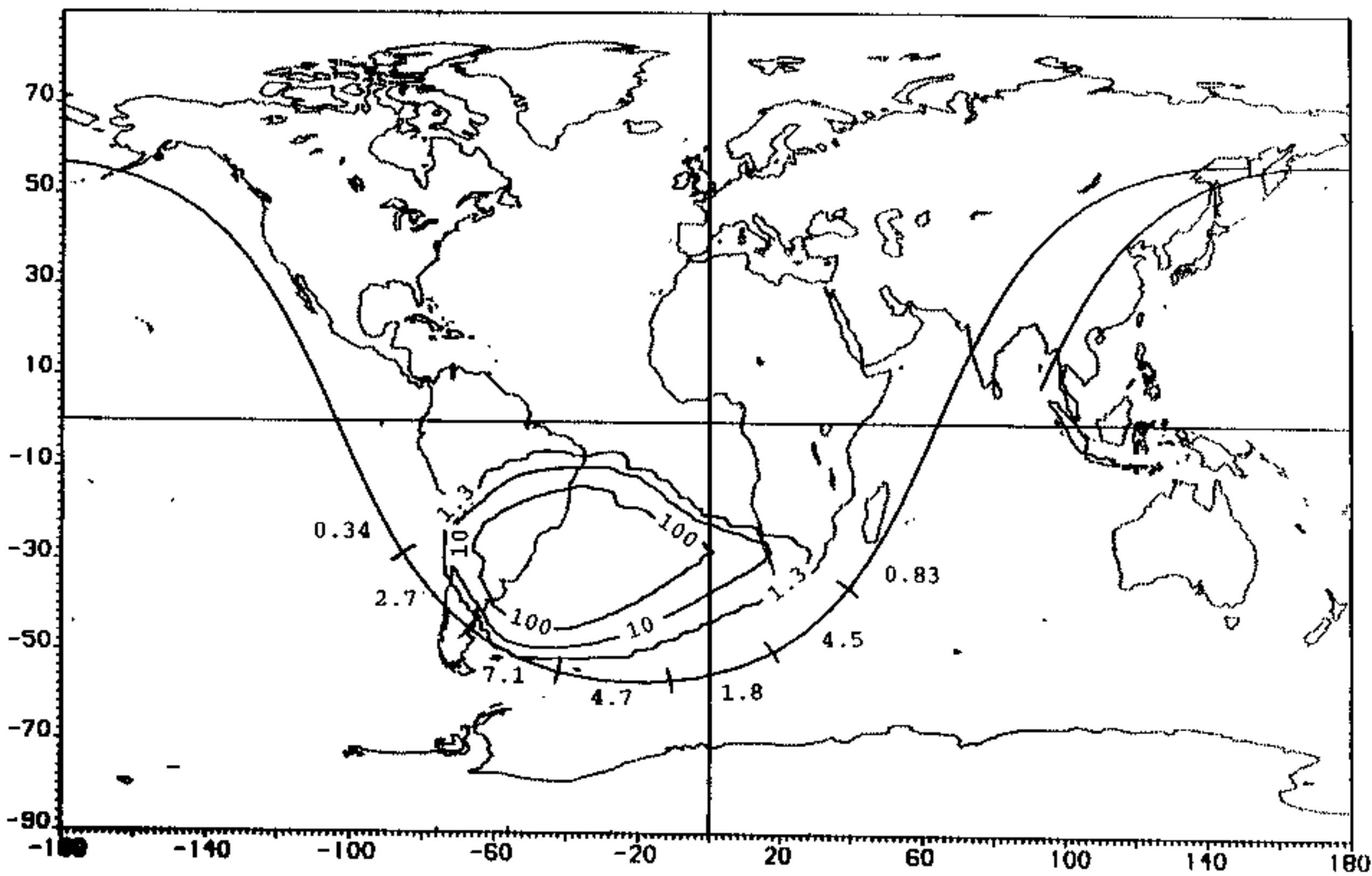


Anisotropy in environment

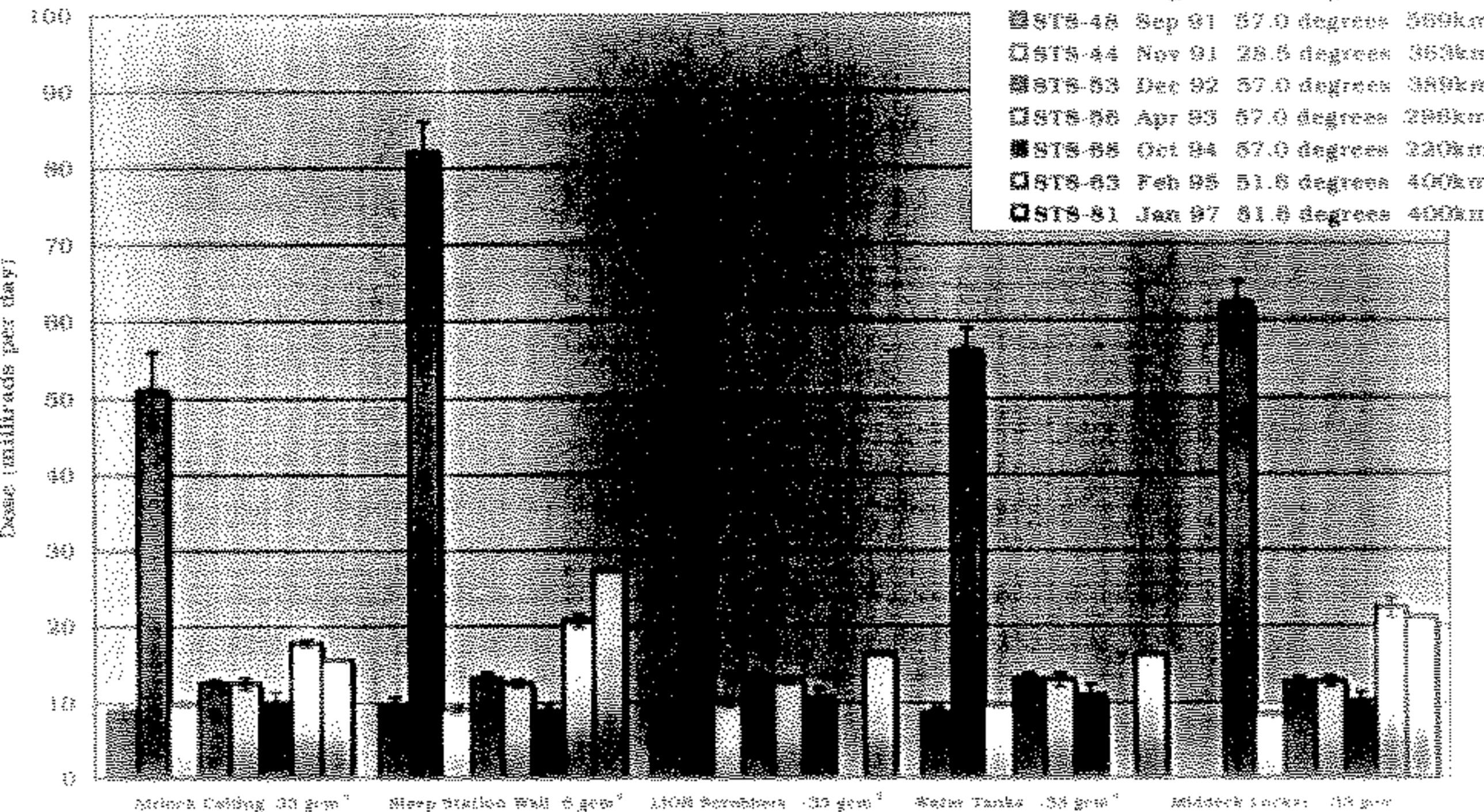
SAA for all 1992 & 1995 UOSAT-3 Ch1 data



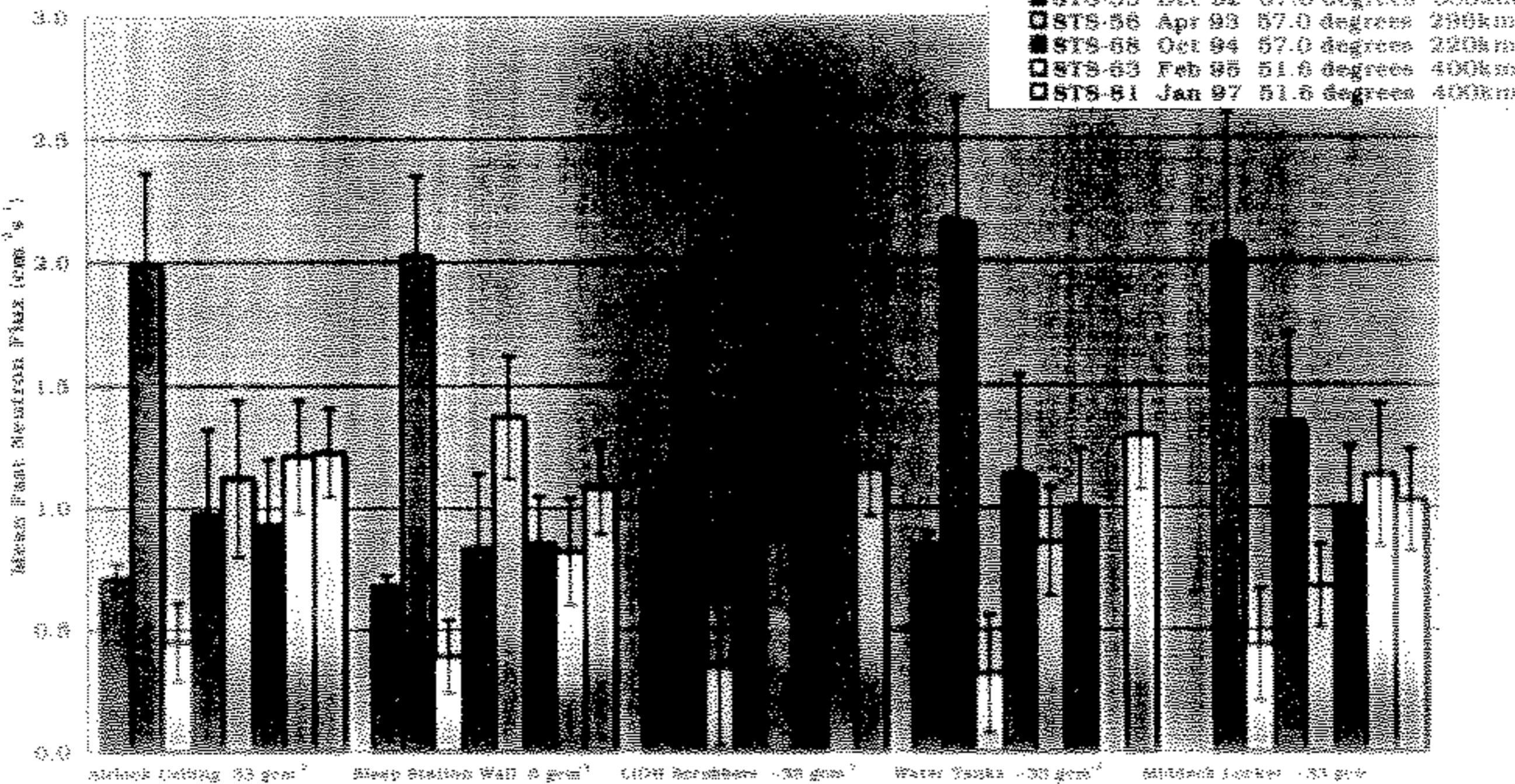
Flux map at 541.0 km for protons > 100.0 MeV at solar maximum
Flux model : AP8 Units : $\text{cm}^{-2}\text{s}^{-1}$
STS48 MET 1/11:00 to MET 1/13:00 1970 FIELD



Thermoluminescent Dosimeter Results



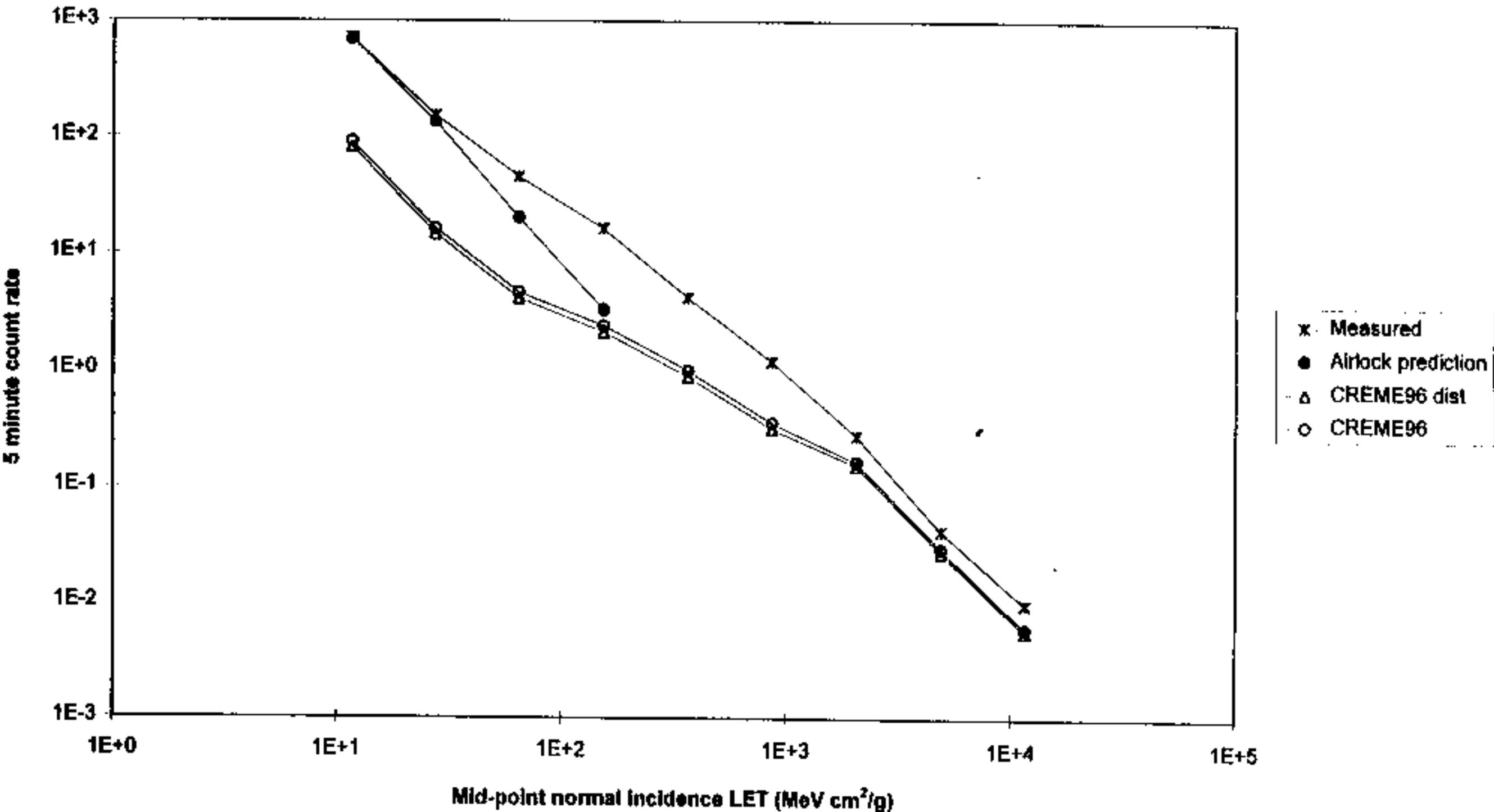
Nickel Activation Soil Results - Fast Neutrons



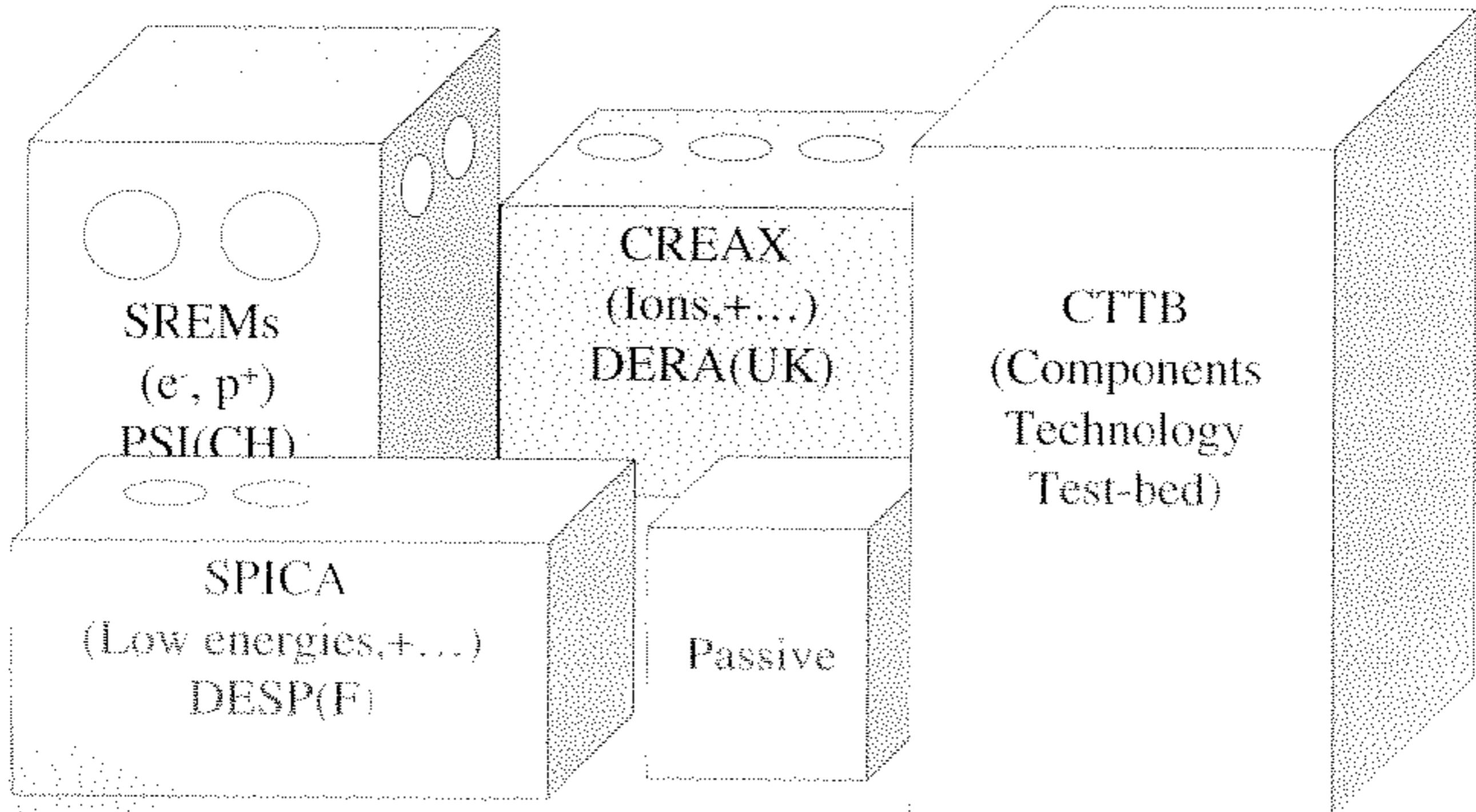
CREAM passive detector results from on board MIR

Location	STS-89 Long Jan - Jun 98 51.6° 380km			STS-89 Short Jan-98 51.6° 380km
	Average dose rate from DERA TLD's [mrad(Si)/day]	Thermal neutron flux from gold foil [cm ⁻² s ⁻¹]	Fast neutron flux from nickel foil [cm ⁻² s ⁻¹]	Mean neutron dose from bubble detectors [mrem/day]
Loon#1 (Sleep Station Core Module)	21.8±1.0	0.148 ± 0.056	1.138 ± 0.061	10.5 ± 2.8
Loon#2 (Panel 410)	18.0±2.8	0.176 ± 0.049	1.363 ± 0.074	10.8 ± 3.3
Loon#3 (KVANT-II Water Tanks)	21.2±0.4	0.163 ± 0.055	1.152 ± 0.075	10.5 ± 3.2
Loon#4 (KVANT-II)	26.4±1.2	0.274 ± 0.053	1.151 ± 0.073	9.8 ± 2.2
Loon#5 (Panel 323)	16.6±0.4	0.203 ± 0.049	1.295 ± 0.074	10.5 ± 2.5

CREAM flight on STS-63 (Feb 1995)
Middeck airlock location - orbit averaged



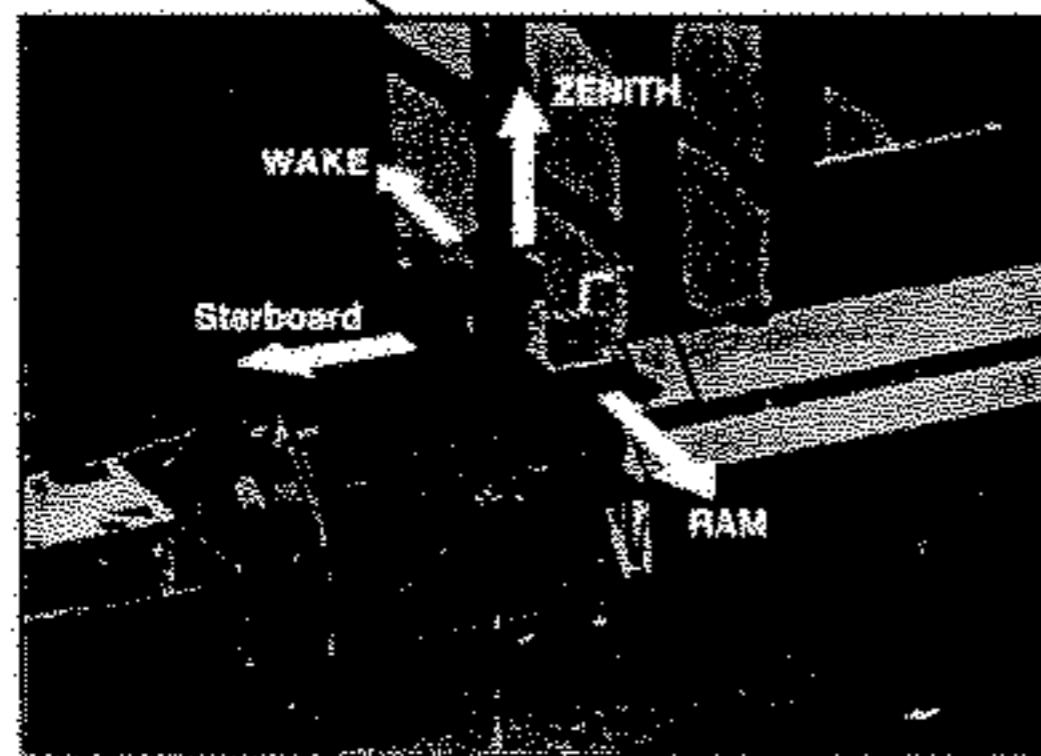
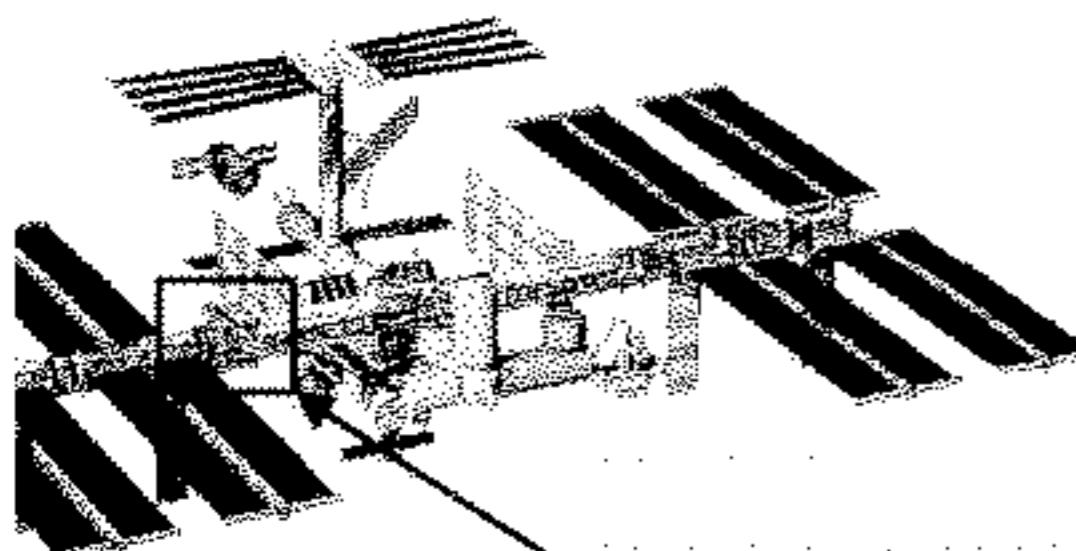
Columbus Radiation Environments and Effects Package



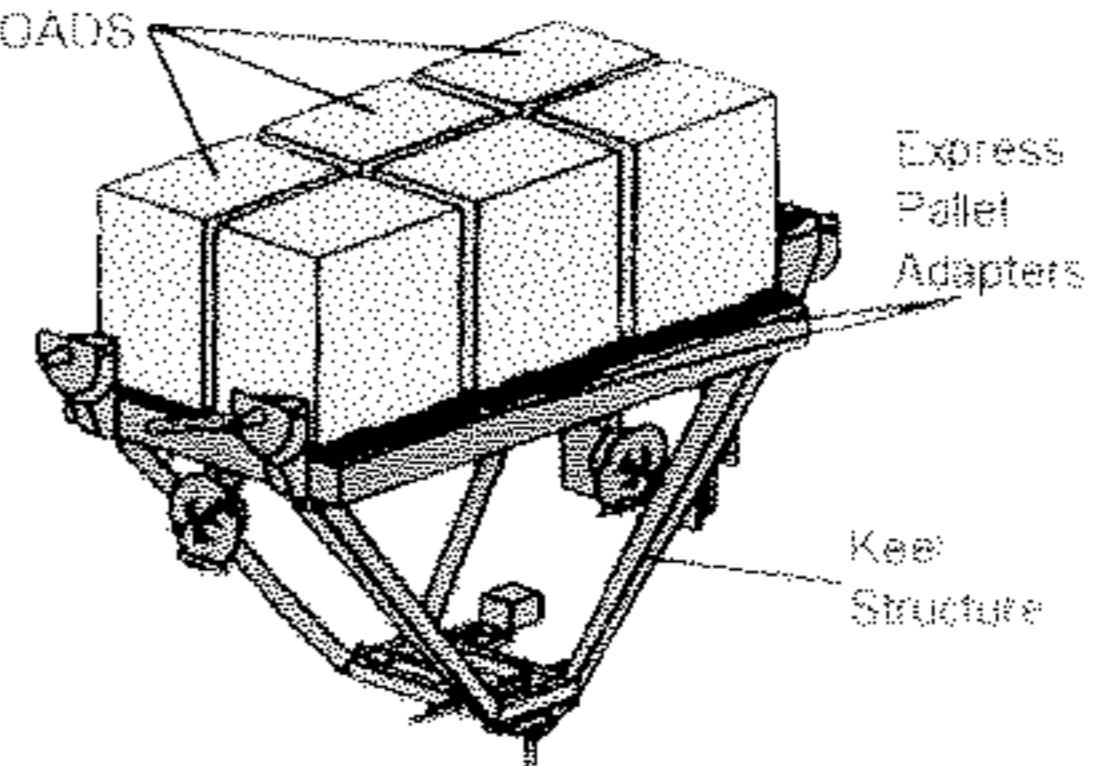
CREEP

SRM (x2)*	Directional electrons & protons (radiation belt, solar energetic particles)	0.5 MeV < E _e < 10 MeV 10 MeV < E _p < 300 MeV
CREAX*	Ions, protons, nuclear interactions, activation studies	1.5 < LET < 10 ⁵ MeV.cm ² /g
SPICA*	Low energy electrons and protons for materials studies	0.05 MeV < E _e 3 MeV < E _p
CTTB	Sub-experiments for technology research / demonstration / test	Digital, analog, photonic, ...
Ground	- Comprehensive calibration & simulation - CTTB technology pre- & post-flight evaluation;	- European proton (PSI), ion (various) & electron (various) beams;
	Data analysis & dissemination	- open data policy (except commercial) - real-time & near real-time data

* overlapping capabilities exploited for validation



PAYLOADS

Express
Pallet
AdaptersKee
Structure

CTTB**(Technology Test-Bed)**

- A unit containing technologies for exposure to radiation:
 - Digital Electronics (microprocessors, memories, ASICs,...)
 - Analog electronics
 - Photonics (CCD's, fibre-optics opto-electronics,...)
 - others (novel detectors, etc.)
- Coordinated with ground activities
- benefits from comprehensive local sensing
- Opportunities will be co-ordinated by CREEP prime (TBS)
- Supplier pays
- Already interest

Benefits of CREEP

- all key issues addressed
- support to station & payloads, including industry
- further extending & complementing good track record of radiation monitoring in Europe (STRV, MIR, STS, ...)
- unique, comprehensive ground-based part
- possible internal experimental part
- wide international collaboration, including US and Japan
- “space weather” will be important and “high profile” next solar max.
- public visibility: promotion of results, education & awareness