



Status of the Trapped Model AE9/AP9/SPM (IRENE) for the ISS Environment



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WRMISS19, 9-11 September 2014, Krakow, Poland



Outline



- Why a new trapped model?
- A quick review of Budapest presentation (WRMISS18, Sep. 2013)
- Availability of AE9/AP9/SPM on SPENVIS
 - I. Implementation on the current version SPENVIS-4 (IRENE)
- Missing databases in AE8/AP8 and different run modes of AE9/AP9
- AE9/AP9 spectral (energy) gridding for a GTO orbit
- AP9 ISS environment simulation (4 April 2014)
 - I. Perturbation vs. Monte Carlo mode (advantages/disadvantages)
- AP9 model verification with CRRES Model (4 July 2011)
- AP9 validation using POES and TACSAT4 satellites measurements
 - I. January 1999 (after cycle 22 solar minimum)
 - II. January 2005 (after cycle 23 solar maximum)
 - III. October 2011
- Summary and current/future works



Why a New Trapped Model



- Currently there are ~ 1100 satellites in orbit:
 - I. ~ 500 in low Earth orbit (LEO)
 - II. ~ 70 in medium Earth orbit (MEO), mostly US-GPS and RU-GLANOS
 - III. ~ 450 in geosynchronous/geostationary orbit (GEO)
 - IV. ~ 35 in highly elliptical orbit (HEO)
- ~ 80% of global financial markets rely on US-GPS satellites atomic clocks for synchronization (cesium or rubidium)*
 - *US GPS clock error ~ 50 nano sec./day
 - *RU GLANOS clock error: better than US-GPS ???
 - *EU-ESA Galileo (Munich/Fucino)
- Global space industry revenue ~ \$300 billions in 2012
 - I. Communication ~ \$180 billions
 - II. TV (transmission, etc...) ~ \$90 billions
 - III. Radio broadcast (XM, Sirius, etc...) ~ \$5 billions
 - IV. Others ~ \$25 billions



Review of WRMISS18 (part I)

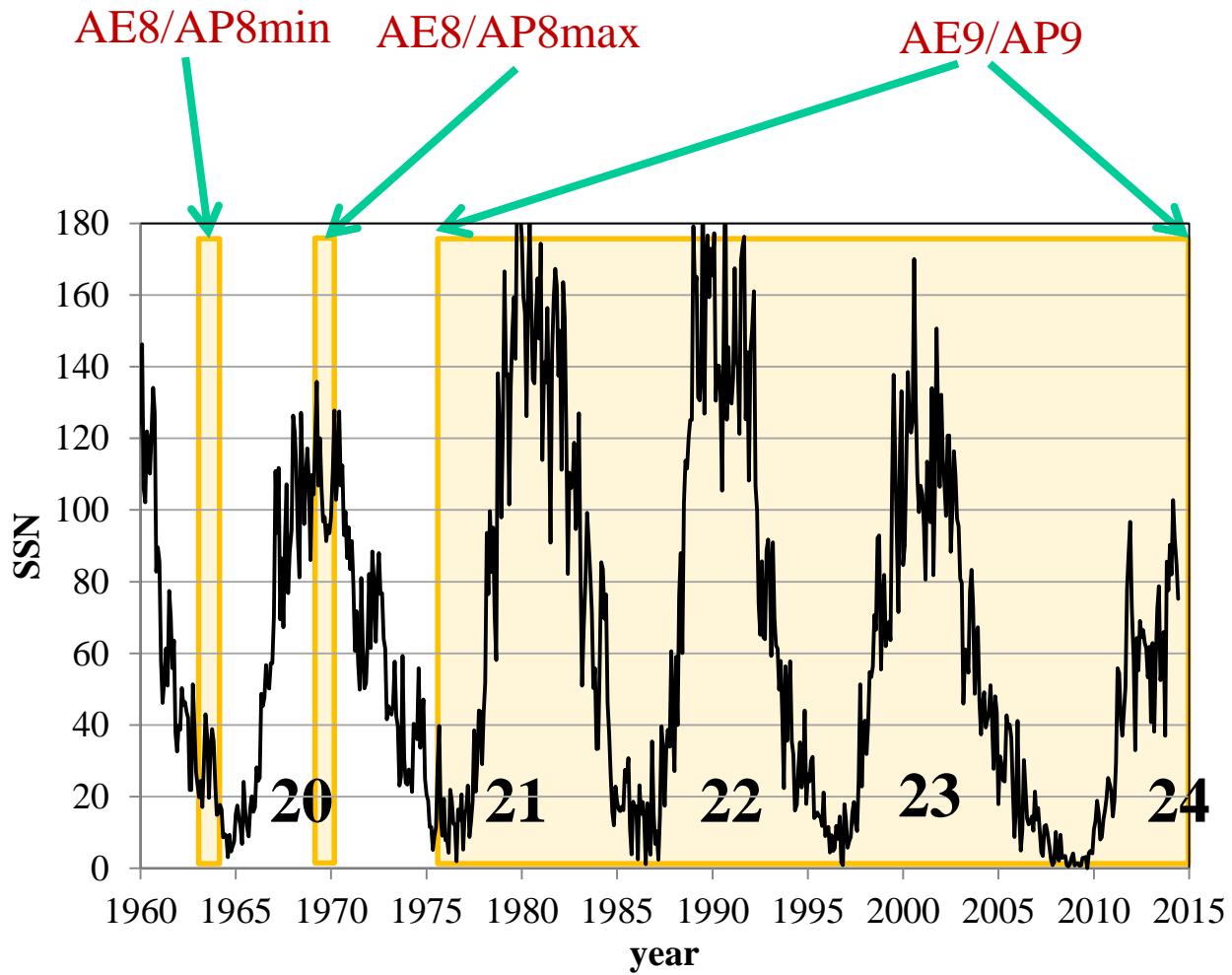


- AE8/AP8 are **static models** (1964/1972 maps) with following limitations
 - I. Lacks **probability distribution** or **error bars** (i.e. no statistics)
 - II. Lacks **sufficient spectral (energy) coverage** for plasmas and energetic electrons/protons
 - III. Are only **omni-directional**
 - IV. Lacks **sufficient spatial coverage** for most orbits (**MEO, GEO, GTO, HEO, ...**)
 - V. AE8/AP8 **under/over predict** most measurements
- In order to overcome the above limitations, US developed AE9/AP9/SPM
 - I. **SPM** is **Space Plasma Model** with directional (anisotropy) capabilities
 - II. In Budapest I mentioned that AE9/AP9 (no SPM) will be licensed to **SPENVIS**
- As of now AE9/AP9 model is not yet used for **ISS daily operations**
- No **ISS external measurements** were included in developing the AE9/AP9/SPM

For the rest of the talk: AE8/AP8 will be called “old model”

AE9/AP9 will be called “new model”

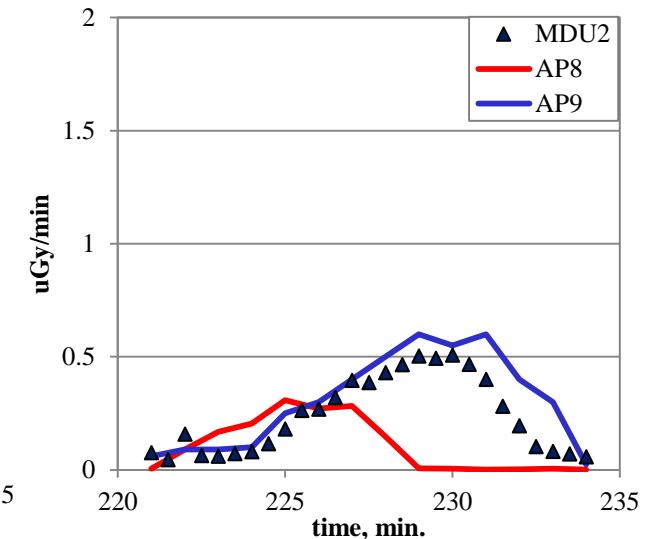
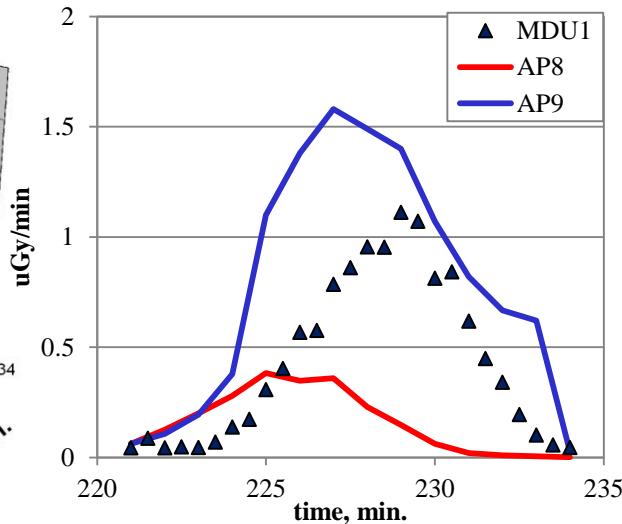
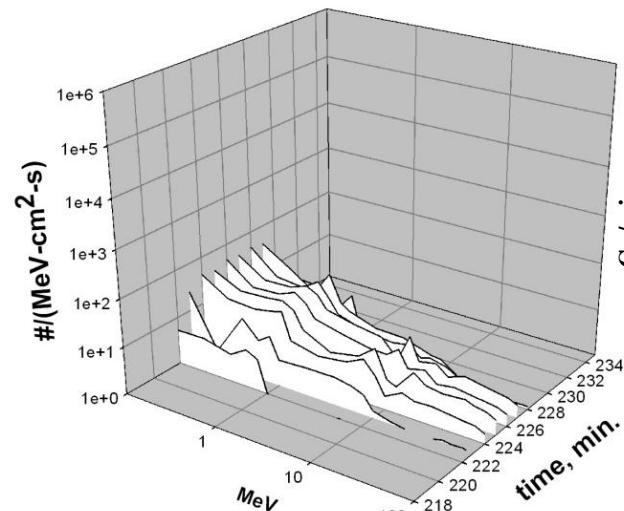
Review of WRMIS18 (part II)



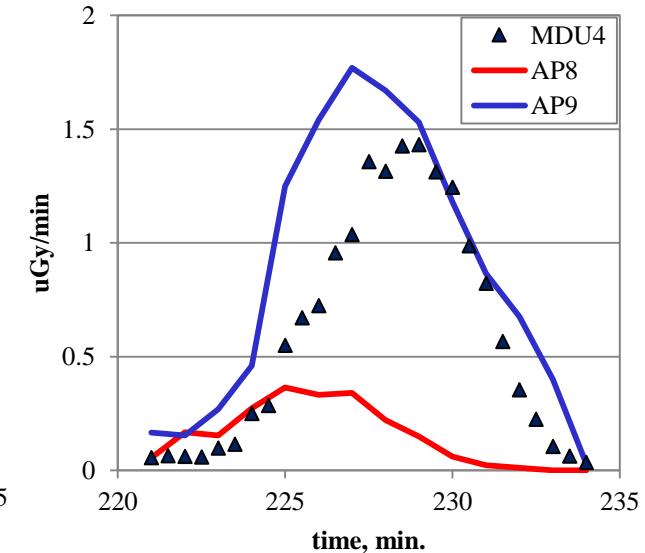
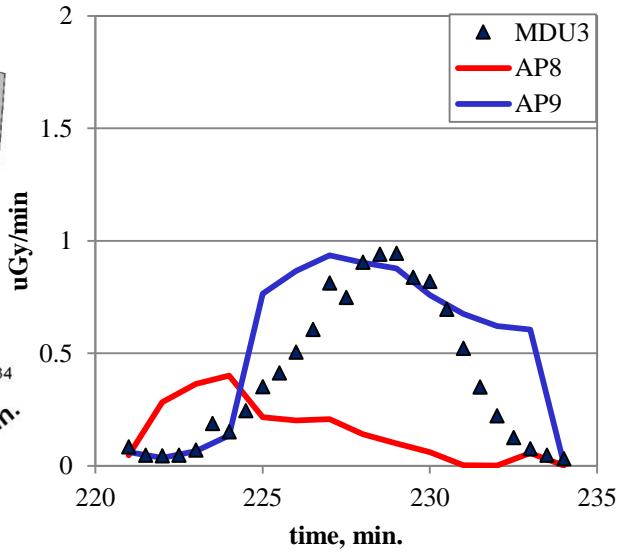
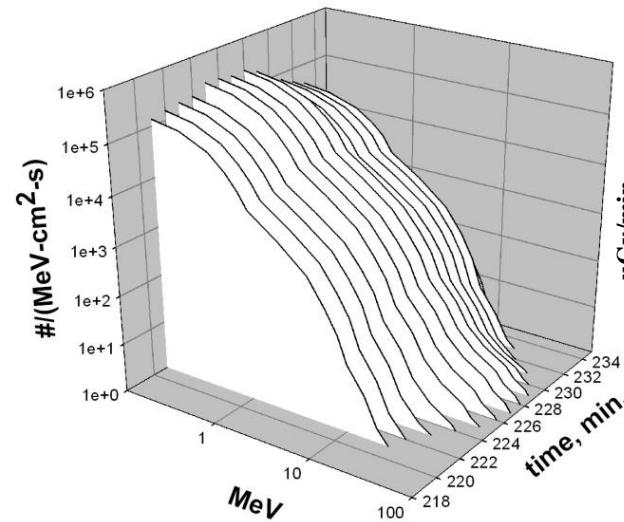
Review of WRMIS18 (part III)



AP8max



AP9





Availability of AE9/AP9 on ESA's SPENVIS-4



SPENVIS Project: TEST-FFB Radiation sources and effects Trapped radiation: IRENE AE9/AP9 parameters

The current version of the AP9/AE9 model is provided for evaluation purpose by its [development team](#)

AE9/AP9 model input

Model run mode: mean

Energies: mean percentile perturbed montecarlo

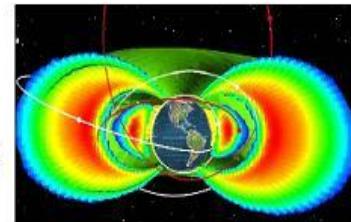
IS energies

Warning: due to the size of the trajectory (4142 points)
the execution time could exceed your allowed quota.

Reset Run << Back

???

Tool developed by



Los Alamos
NATIONAL LABORATORY
EST. 1943



Missing Databases in AE8/AP8 Model



- AE8/AP8 has no database to provide information on the uncertainty in the mean flux maps due to **measurement/gap-filling** errors
- AE8/AP8 has no database to provide information on the uncertainty in the mean flux maps due to **dynamic variations of space weather processes**
- AE9/AP9 provides databases for both **measurement/gap-filling** errors and **dynamic variations of space weather processes** errors
- These databases allow **extraction of statistical information** from AE9/AP9



AE9/AP9 Runs Modes



- The AE9/AP9 model offer 4 **run modes** corresponding to various types of flux data
 - **Mean:** mean behavior of the model with no uncertainty added
 - **Percentile:** statistical behavior of the model with no uncertainty added. Uncertainty allows dynamic estimation of design margin (e.g. 99% CL) which allows study of surface or internal charging, SEU and evaluation of satellite lifetime
 - **Perturbed Mean (PM):** adds the uncertainty in the mean flux maps due to **measurement/and gap-filling** errors
 - **Monte Carlo (MC):** contains all of the **PM** uncertainty plus an estimate of the **dynamic variations** due to **space weather processes**

***PM** and **MC** selections require ‘number of scenarios (runs)’. I will show results for 10 scenarios later

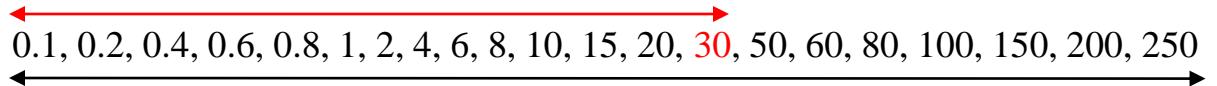


Energy Bins Differences (part I)

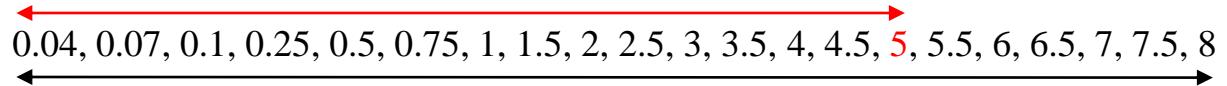


Default IRENE grid

Proton Energies
(MeV)



Electron Energies
(MeV)



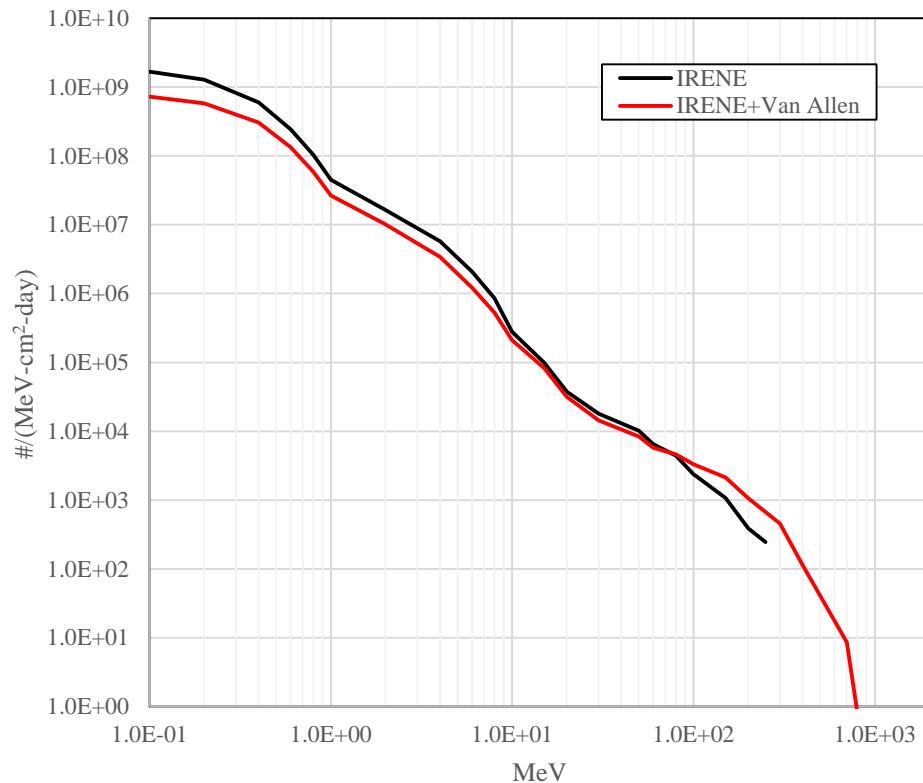
Latest version (1.2)
Proton Energies
(MeV)

0.1, 0.2, 0.4, 0.6, 0.8, 1, 2, 4, 6, 8, 10, 15, 20, 30, 50,
60, 80, 100, 150, 200, 300, 400, 700, 1200, 2000

Energy Bins Differences (part II)



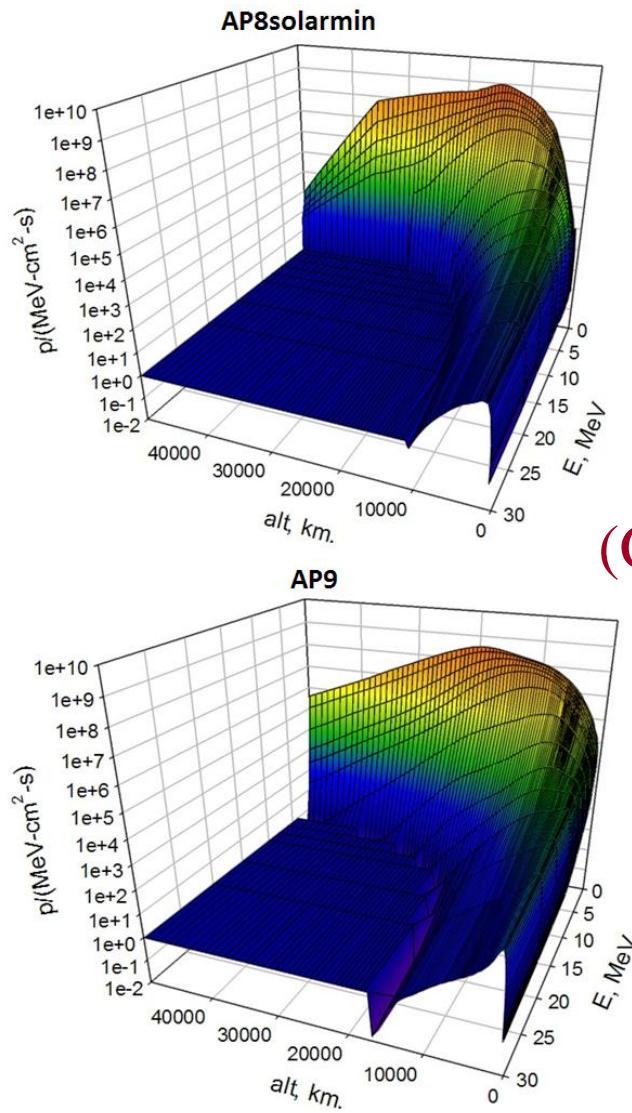
April 4, 2014



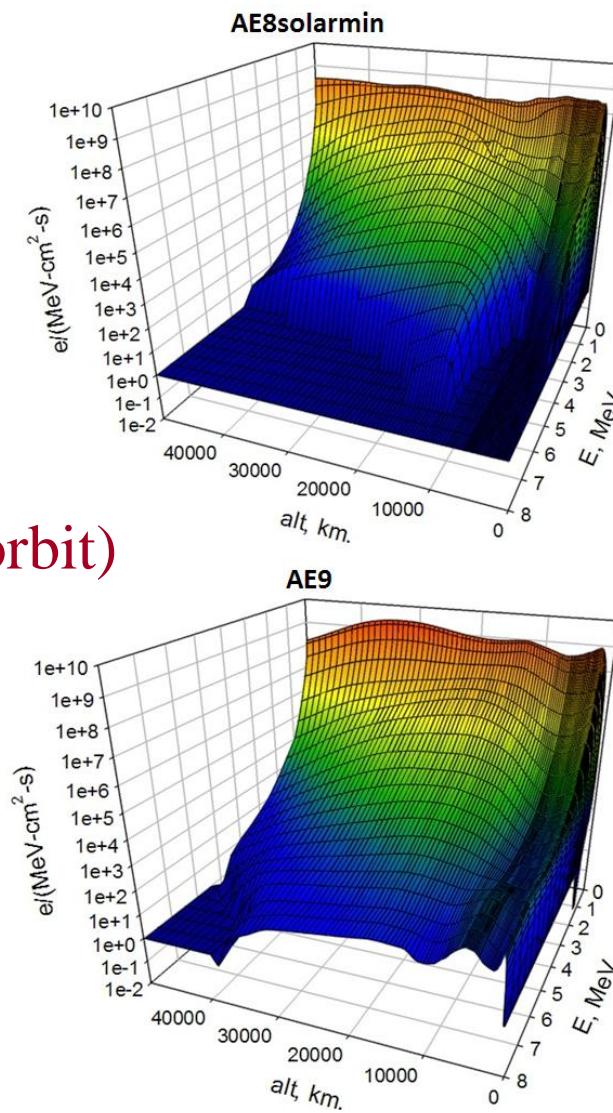
Aluminum 1016 g/cm²

Water 805.4 g/cm²

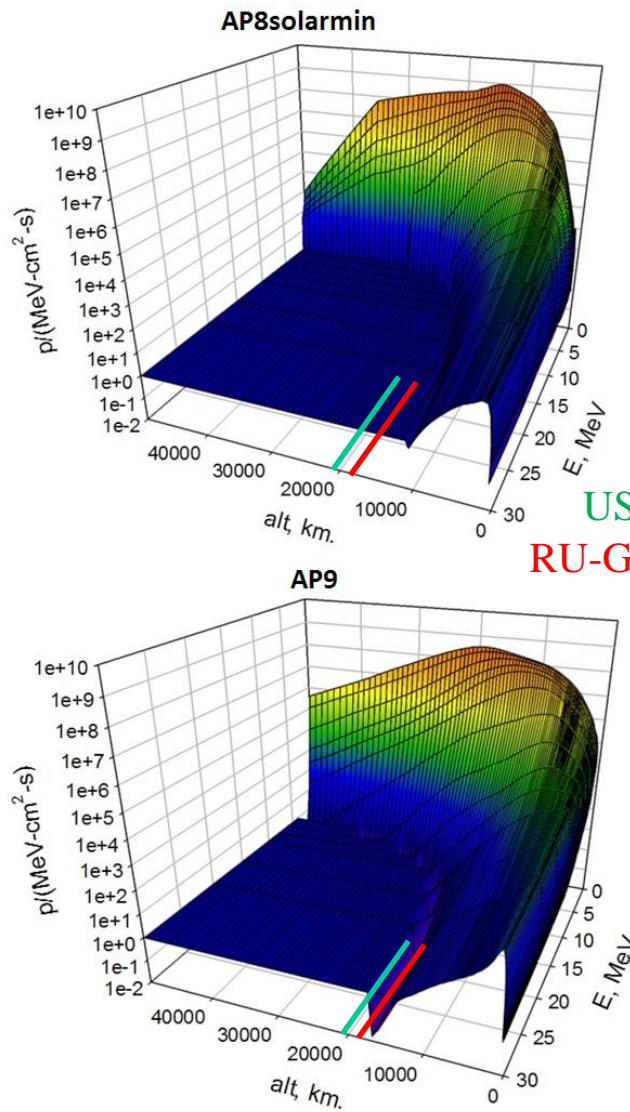
Energy Bins Differences (part II)



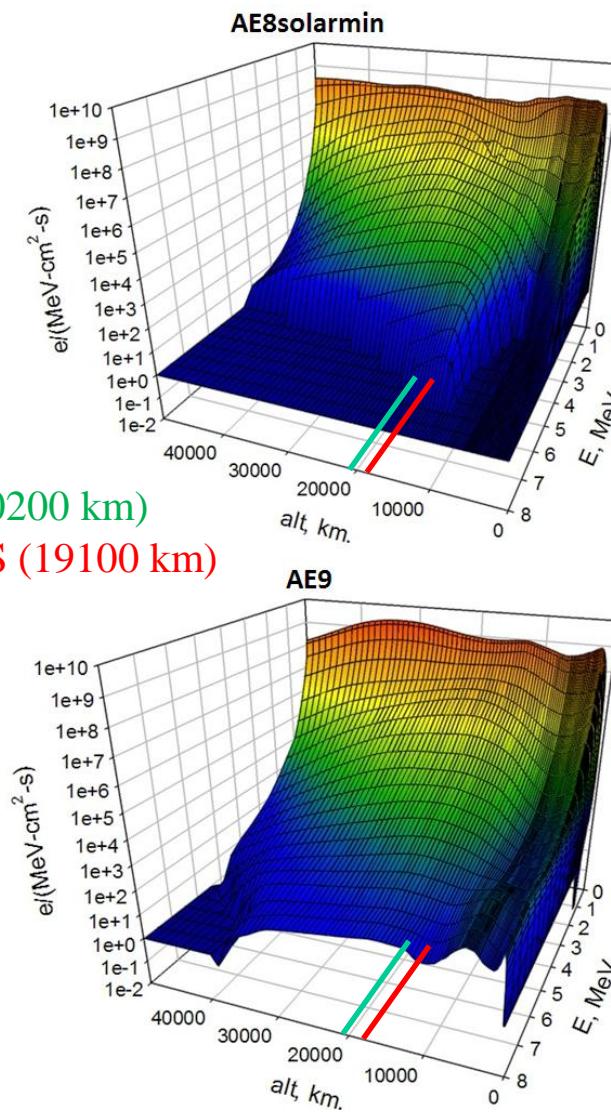
(GTO orbit)



Energy Bins Differences (part III)

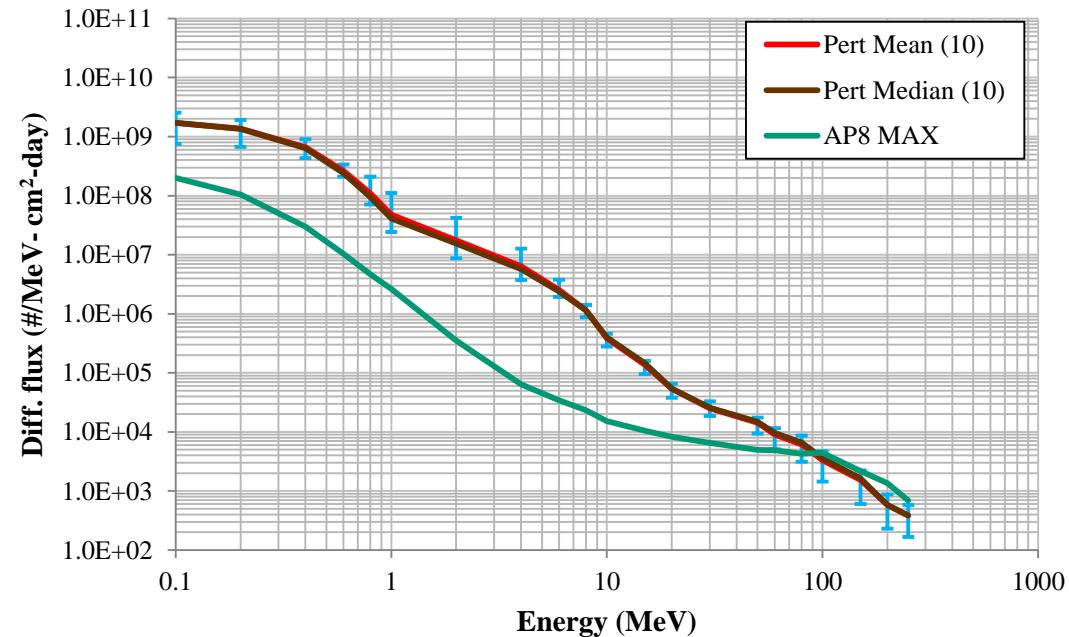
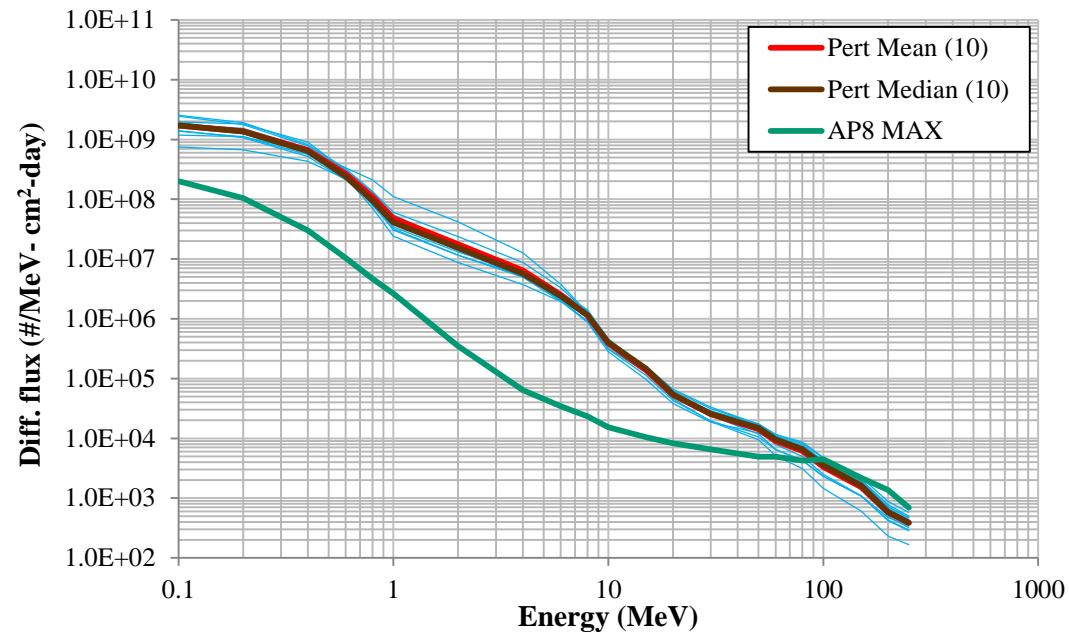


**US-GPS (20200 km)
RU-GLONASS (19100 km)**



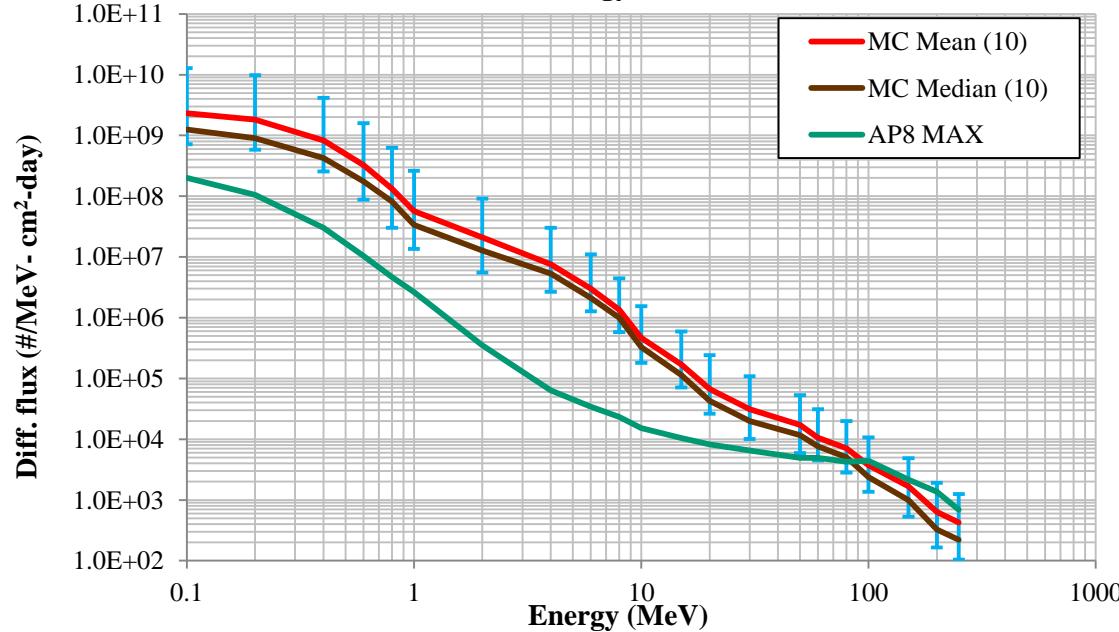
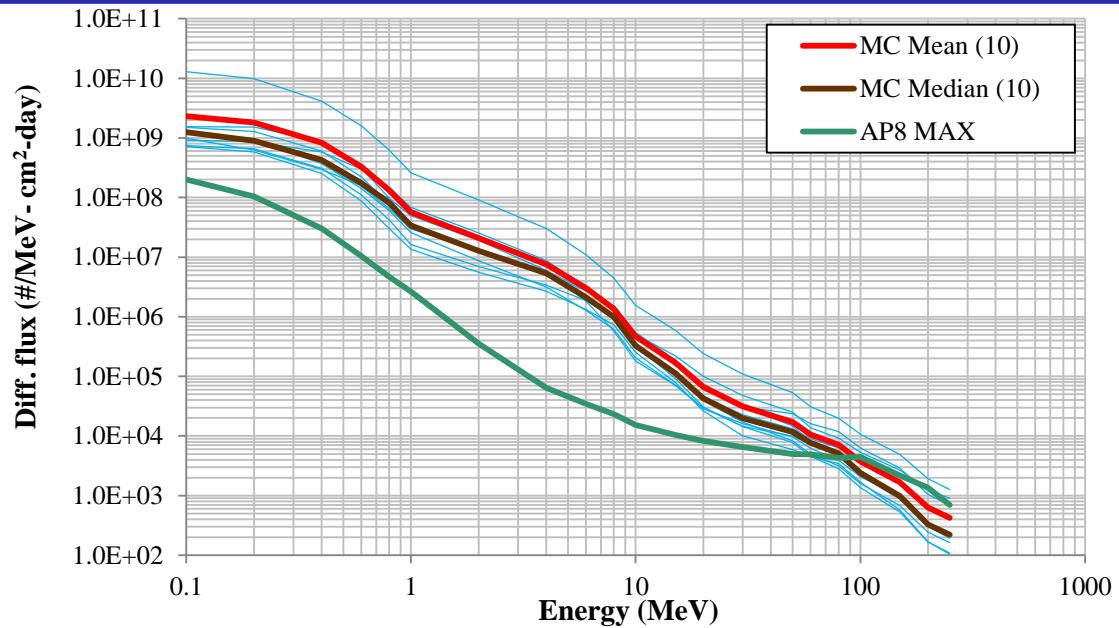


PM Mode (10) AP9 ISS Spectra (4 April 2014)



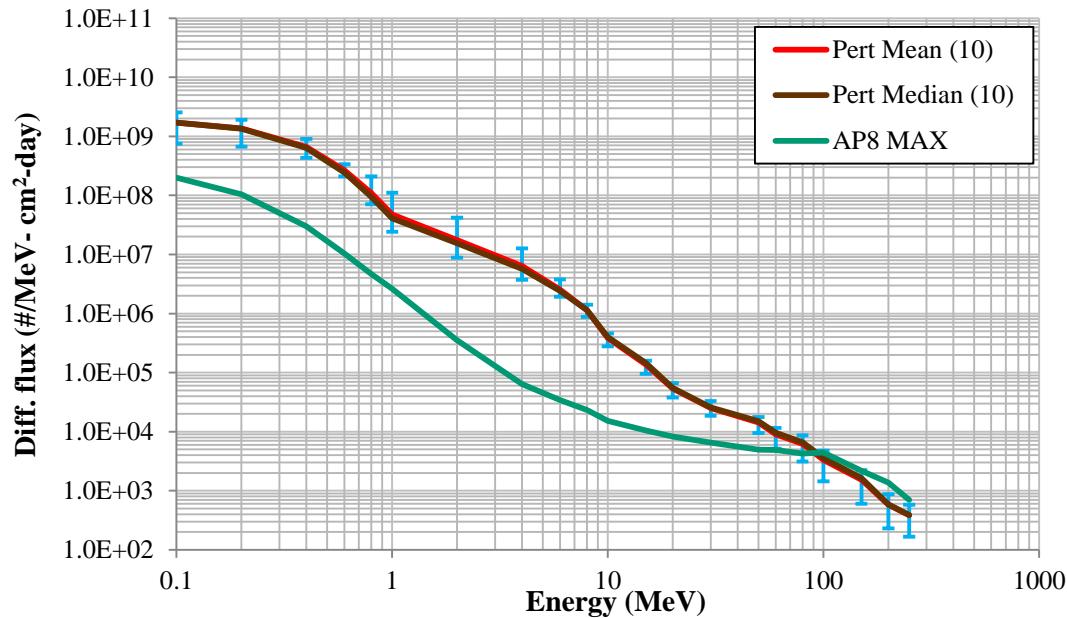


MC Mode (10) AP9 ISS Spectra (4 April 2014)

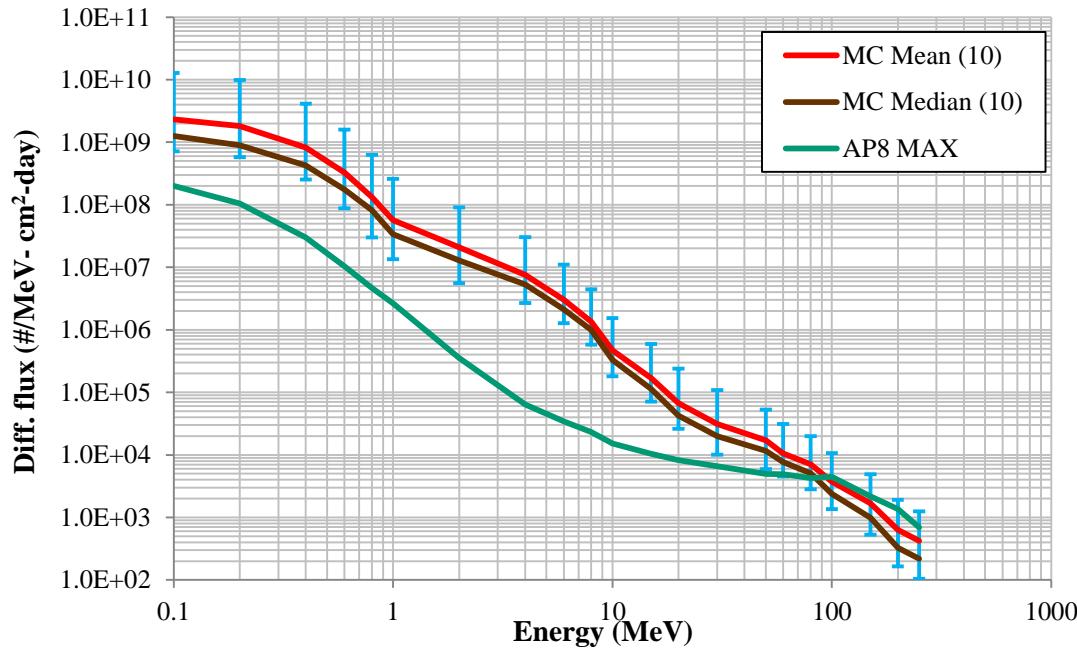




ISS, PM/MC Modes Comparison (4 April 2014)



PM mode (10)



MC mode (10)



Satellite/Sensor	Orbit	Energy range
Protons CRRES/PROTEL	50 km×33000 km, 18°	(MeV) 2.0 - 80

All AP9 scenarios ran for **1 week** mission time

All AP9 scenarios used **10 Monte Carlo** runs

For AP9 only aggregated mean, median and 95%CL values are included

For AP8 only **mean** values are available

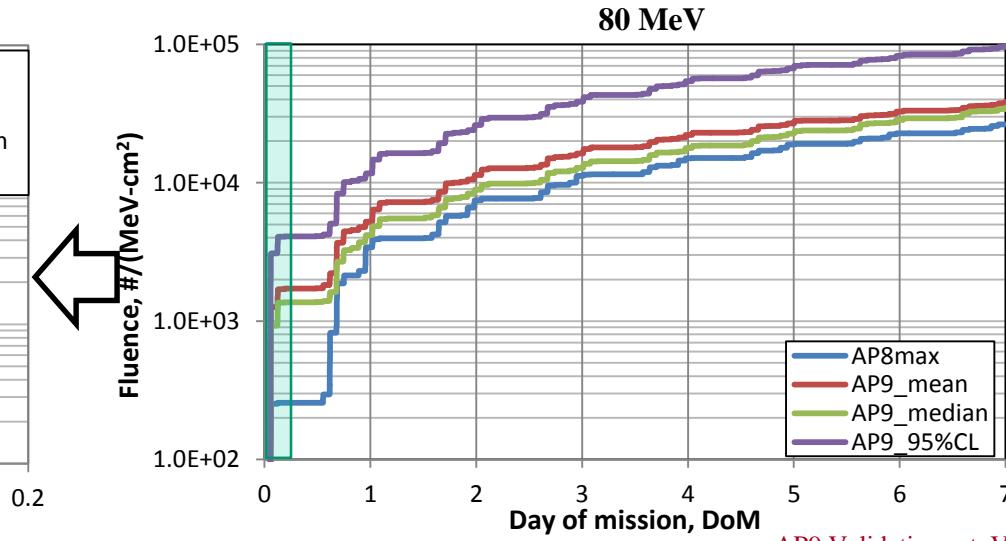
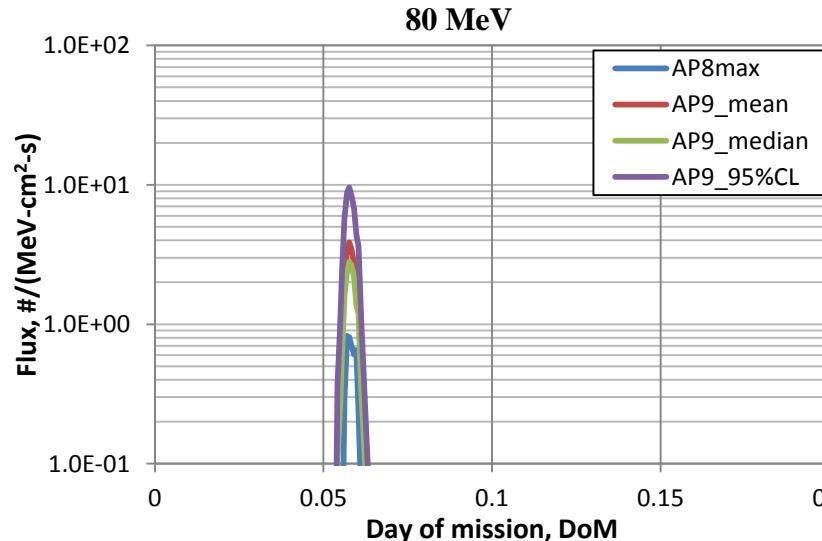
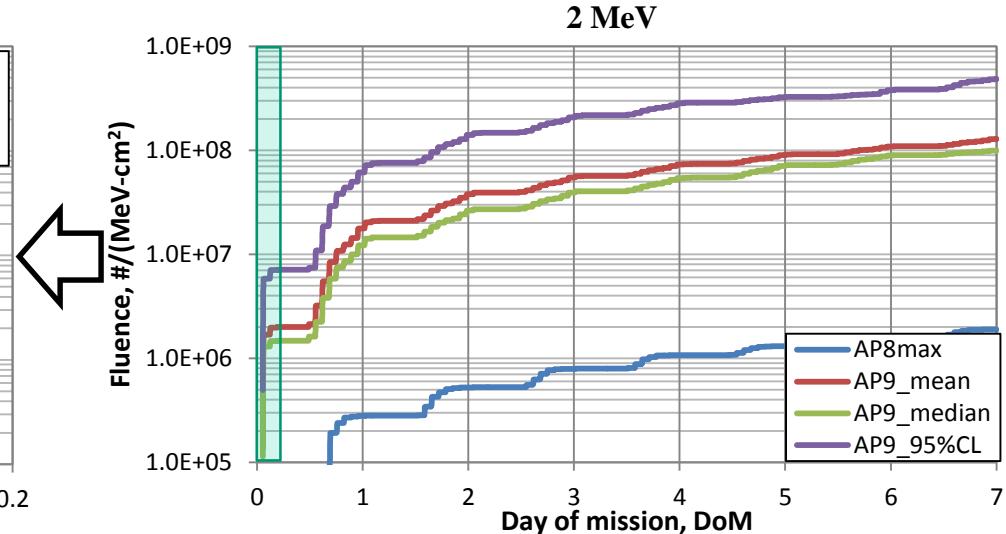
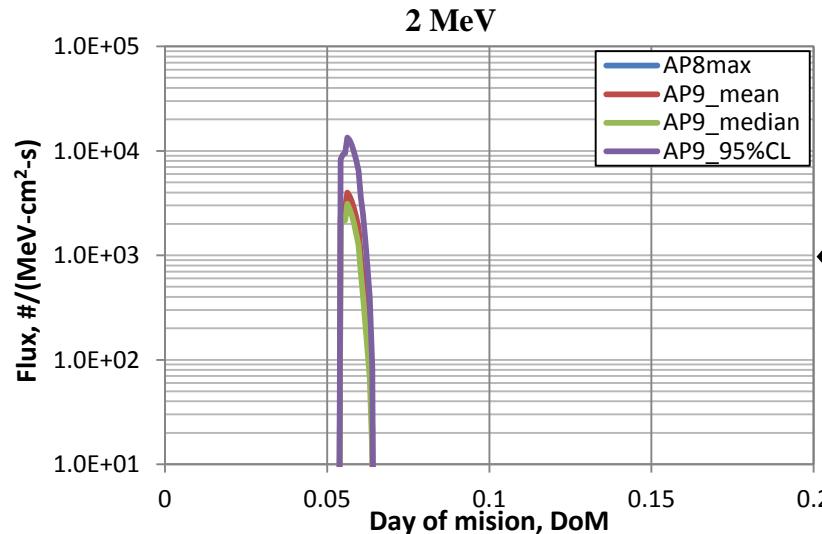
CRRES: Combined Radiation and Release Experiment Satellites



AP9 Model Verification with Other Models (part II)



AP8/AP9/CRRESPRO differential flux/fluence comparisons
LEO-ISS, 400 km X 51.6 deg., circular, epoch 04-Jul-2011

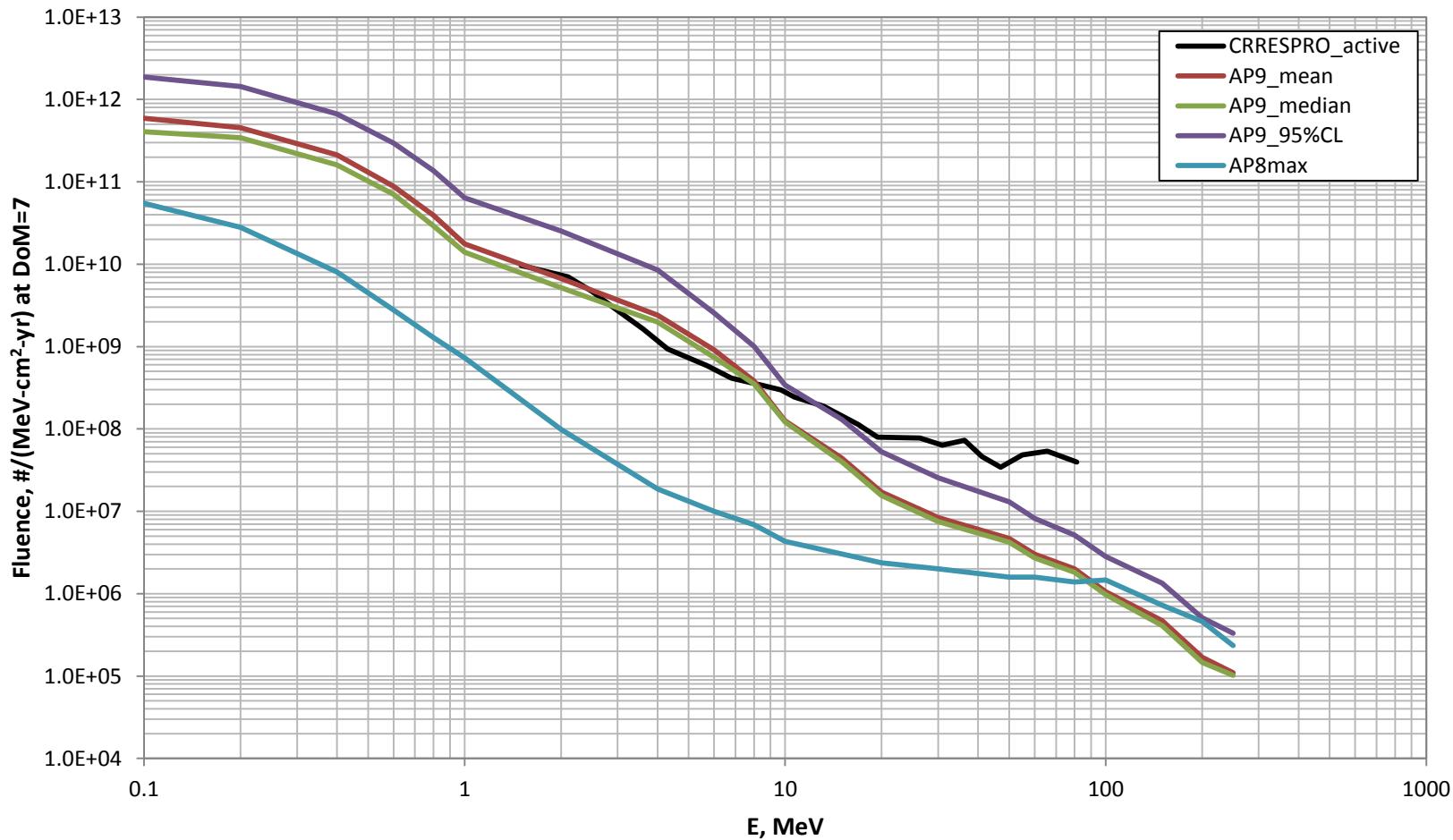




AP9 Model Verification with Other Models (part III)



AP8/AP9/CRRESPRO differential fluence comparison
LEO-ISS, 400 km X 51.6 deg., circular, epoch 04-Jul-2011





AP9 Model Validation Matrix (part I)



AP9 Validation

Satellite	Sensor	Orbit	Time Period	Energies (MeV)
POES	SEM2 MEPED	LEO 850 km, circular, 98.7°	Jul 1998 – Dec 2011	>16, >36, >70, >140
TACSAT4	CEASE	MEO 735 km x 12024 km, 63.5°	Oct 2011 Dec 2012	>16, >29, >39, >44, >72

POES Validation epochs:

Jan. 1999 (after cycle 22 solarmin)

Jan. 2005 (after cycle 23 solarmax)

TACSAT4 Validation epoch:

Oct. 2011

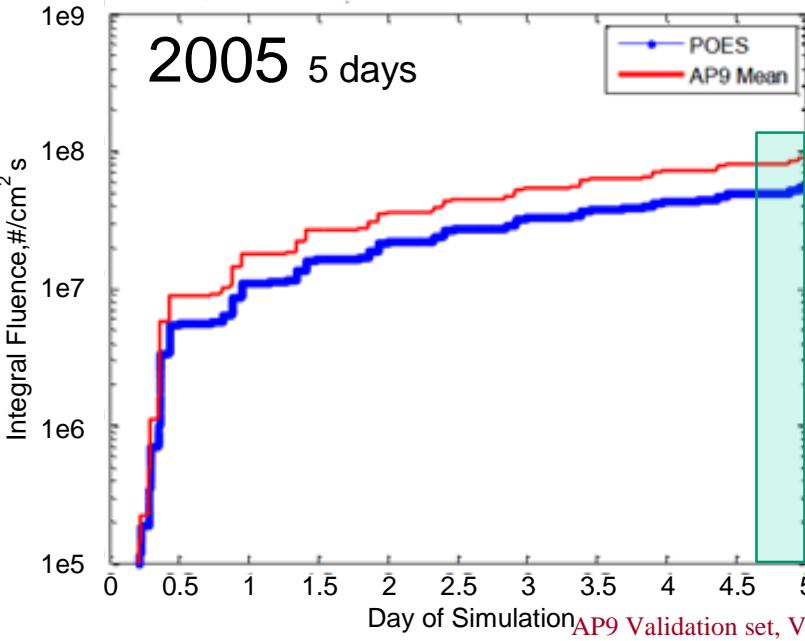
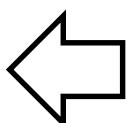
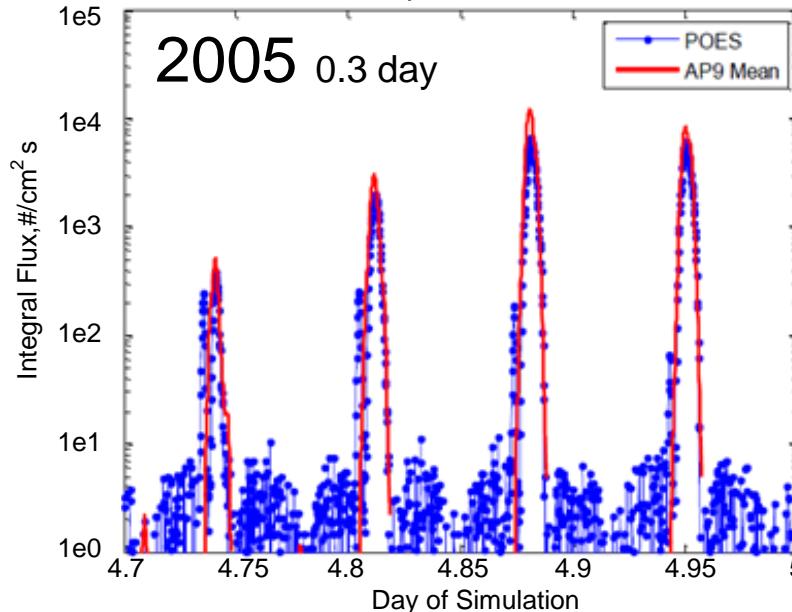
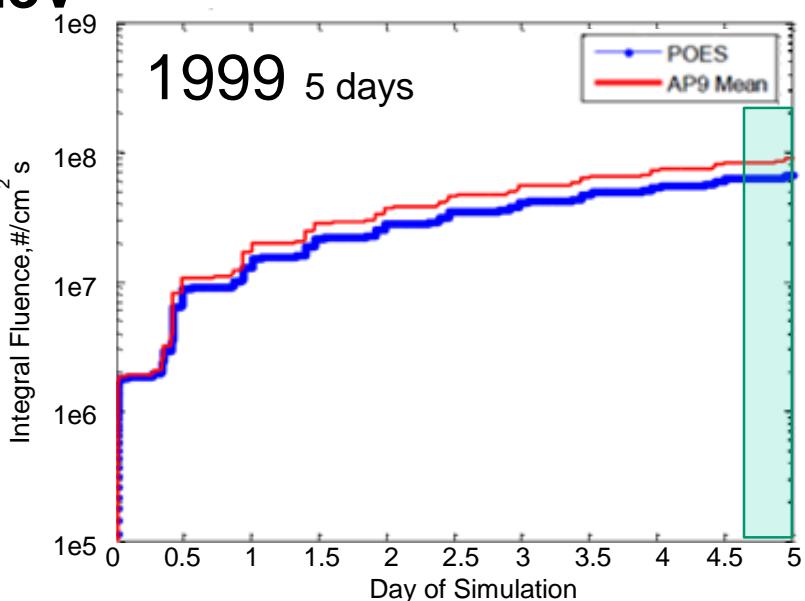
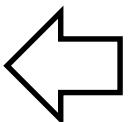
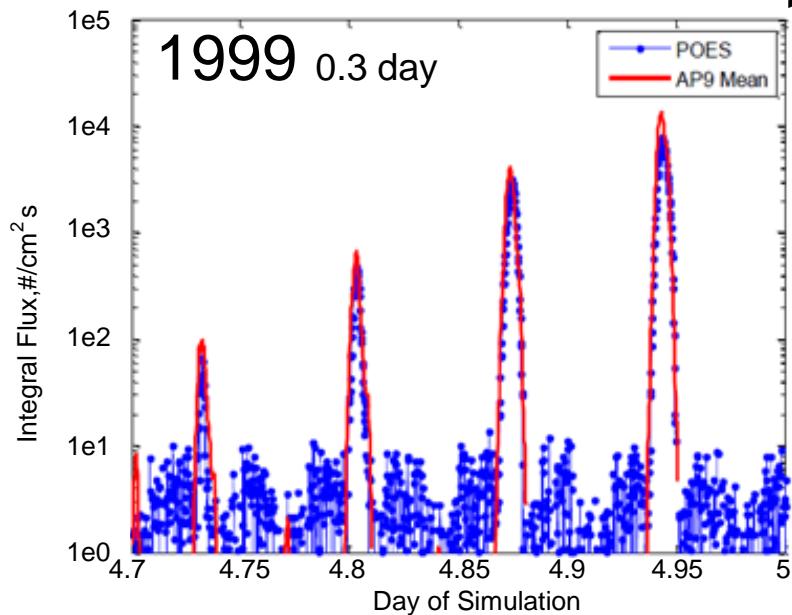
✖ : Excluded due to electron contamination



January 1999/2005 AP9 Validation, POES (part II)

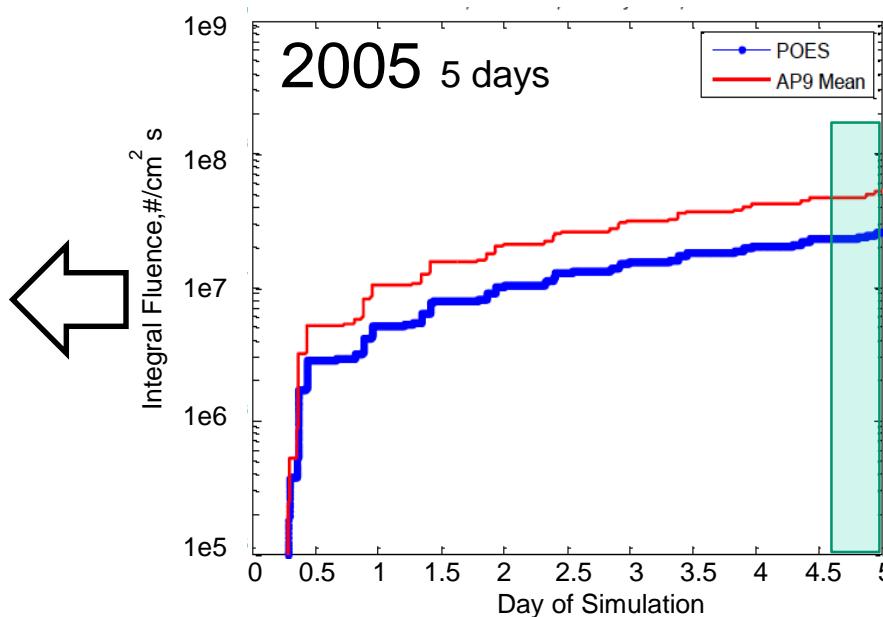
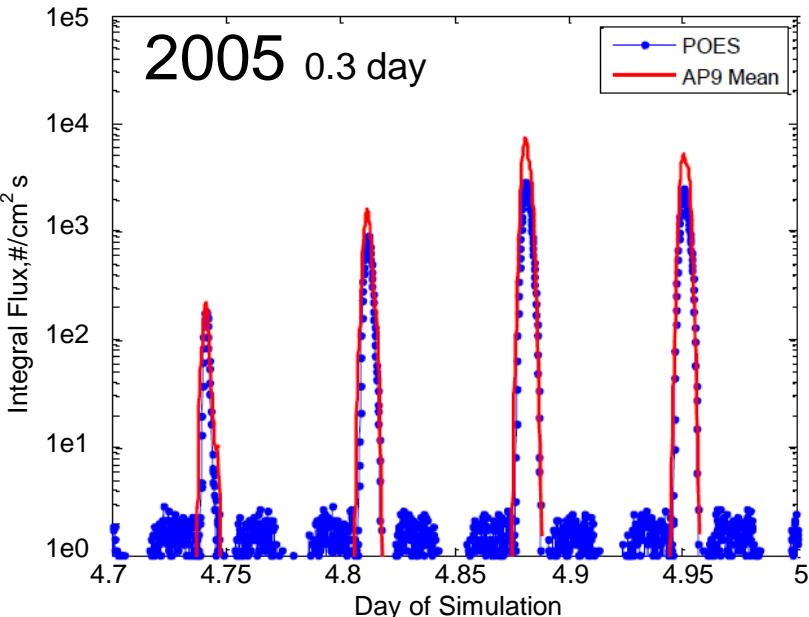
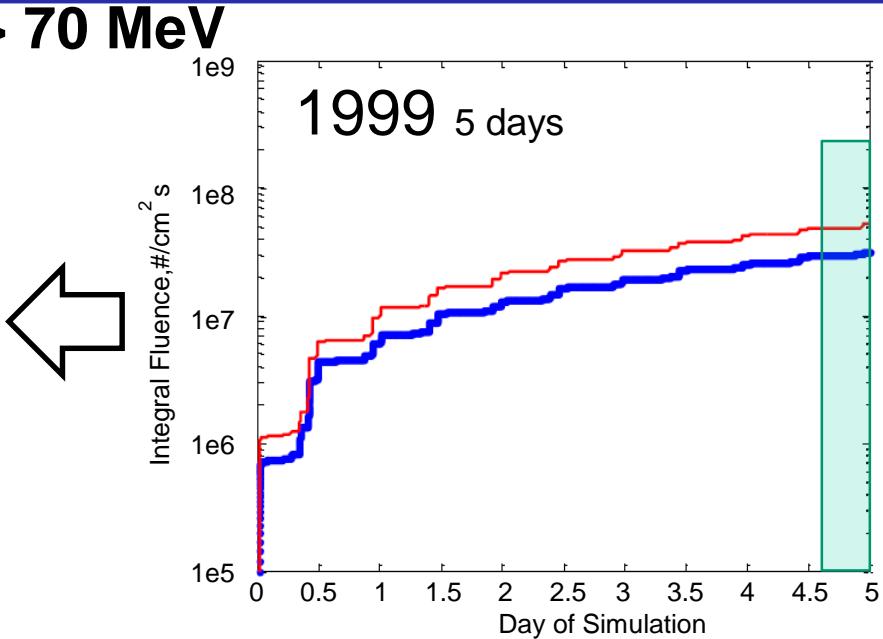
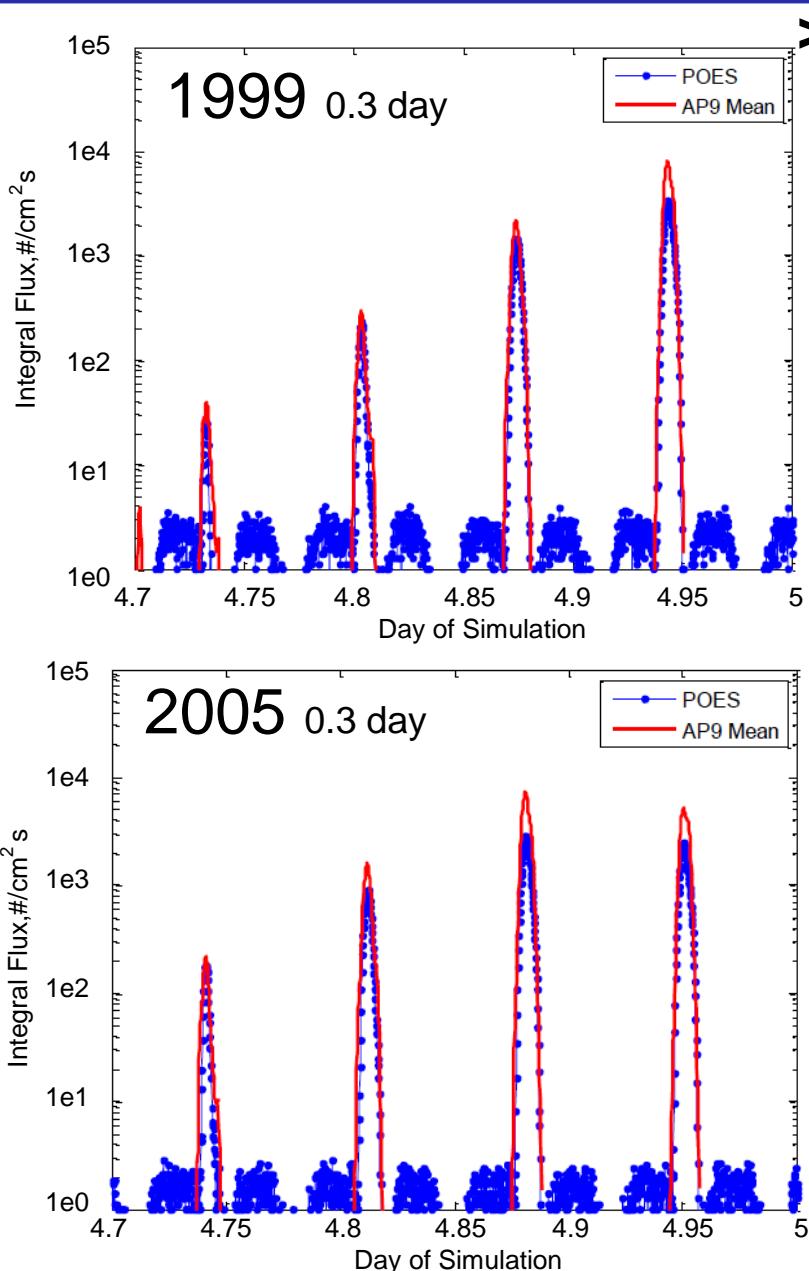


> 36 MeV





January 1999/2005 AP9 Validation, POES (part III)

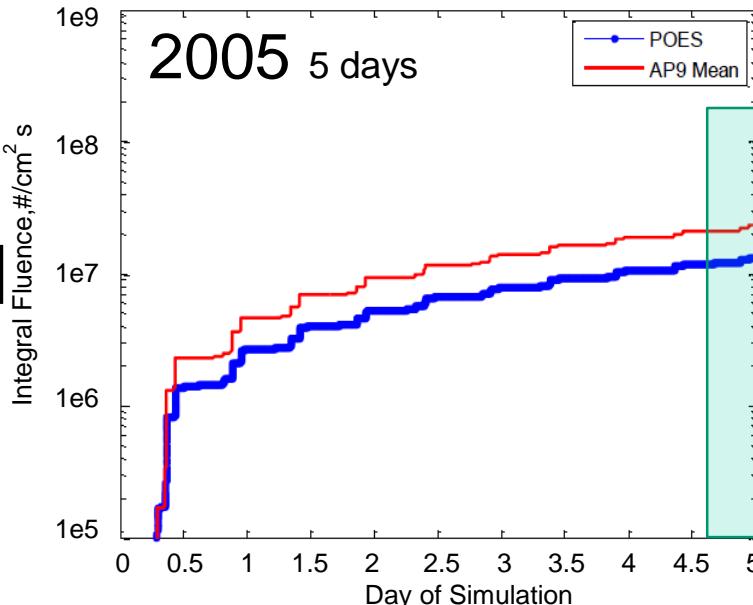
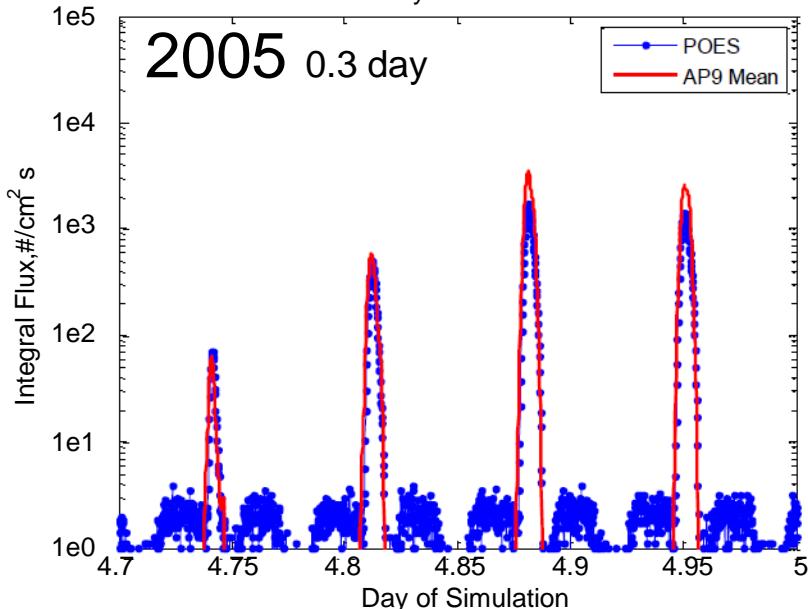
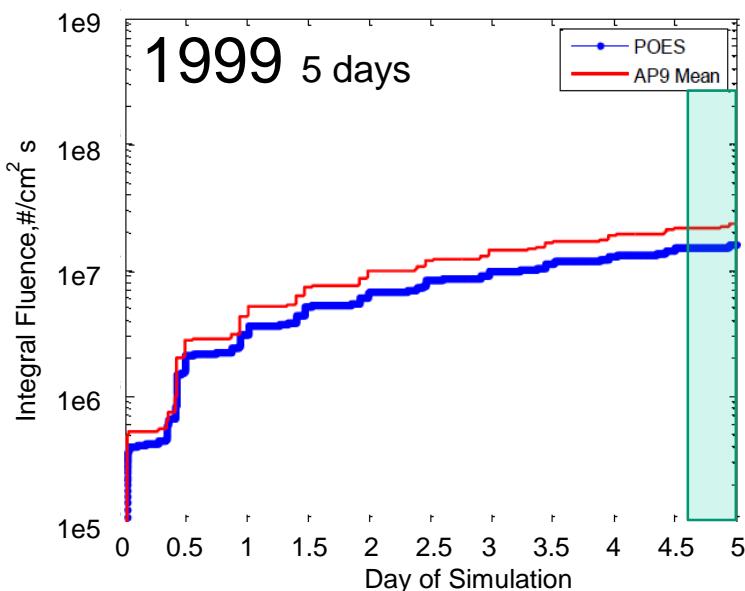
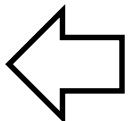
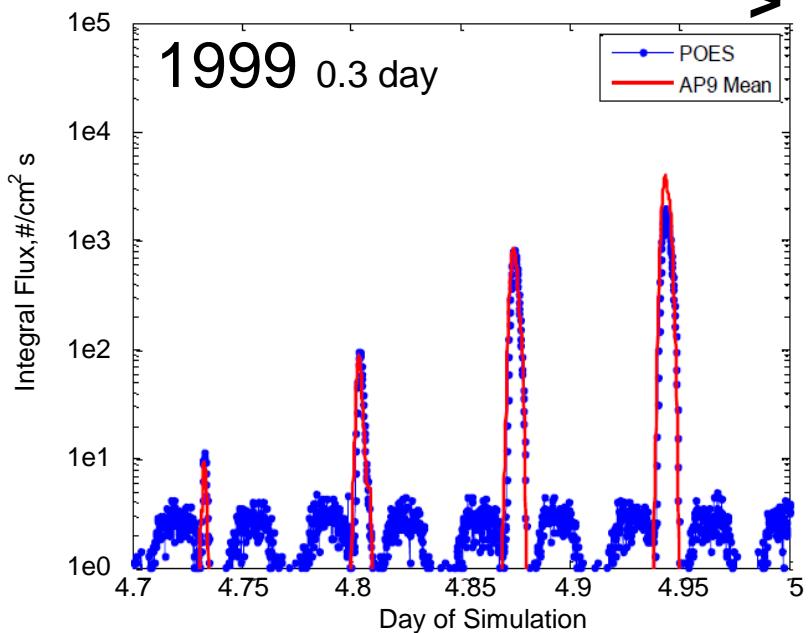




January 1999/2005 AP9 Validation, POES (part IV)



> 140 MeV

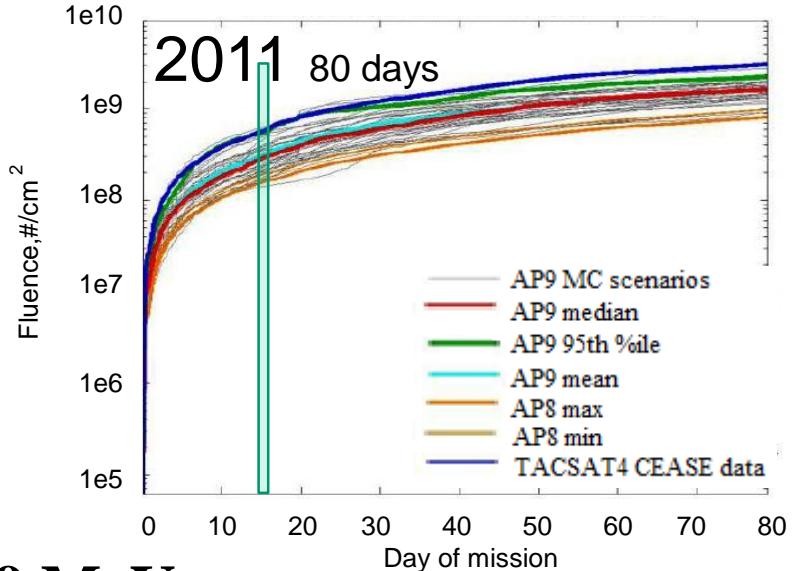
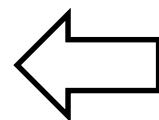
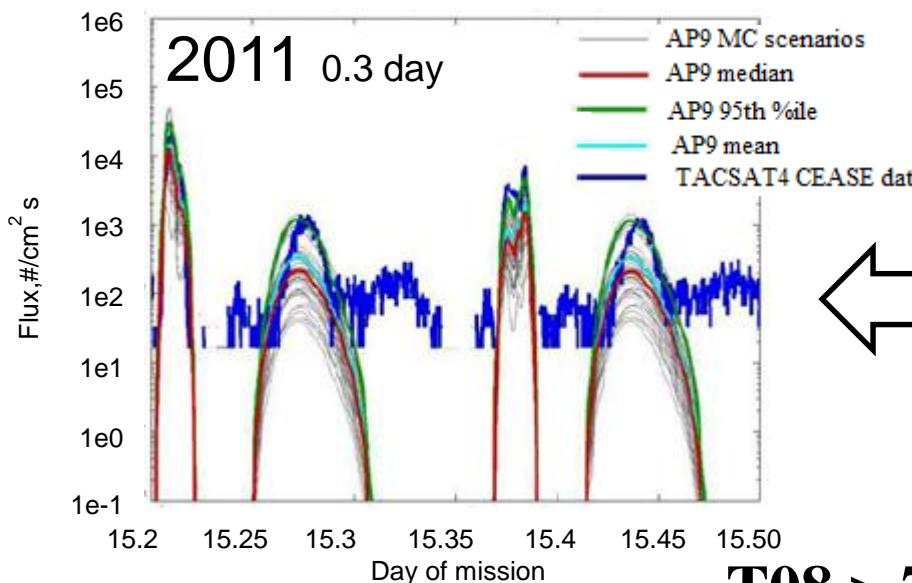




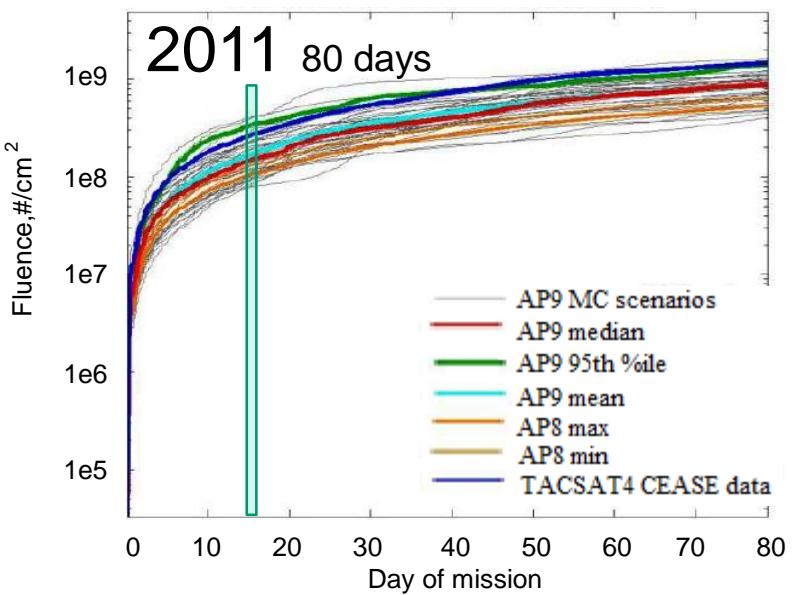
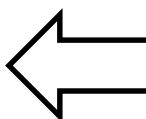
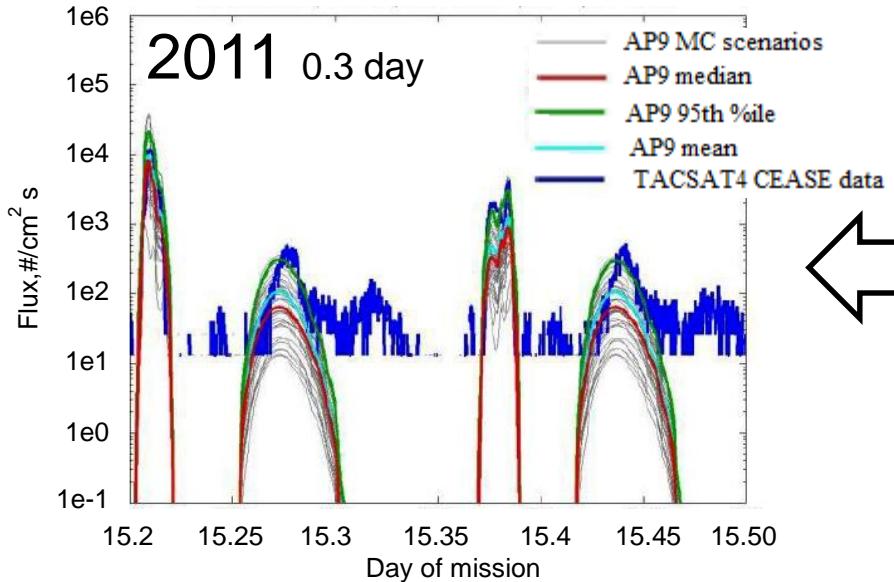
October 2011 AP9 Validation, TACSAT4 (part V)



T09 > 43.7 MeV



T08 > 71.9 MeV





AP9 (PM/MC) Run Times



Transit Trajectory - Crew1out to 50,000 km			123 orbit positions
	Perturbed Mean	Monte Carlo	
10	1.763	2.005	
50	4.532	5.585	
100	7.77	10.063	
250	17.652	23.28	
500	35.757	47.814	
999	66.618	89.445	

1 Day - ISS Orbit			1440 orbit positions
	Perturbed Mean	Monte Carlo	
10	16.092	17.269	
50	41.857	45.414	
100	73.495	80.682	
250	168.289	186.369	
500	327.021	362.435	
999	643.048	714.52	



Summary and Current/Future Works



- Using IRENE (SPENVIS) version of AE9/AP/SPM, I discussed:
 - I. Deficiencies of AE8/AP8 model
 - II. Plans for implementing AE9/AP9 model on SPENVIS-4 and SPENVIS-NG
 - III. Statistical capabilities of AP9 model in **PM** and **MC** modes
 - IV. For ISS, verification among AP8, AP9 and CRRES-proton models
 - V. Validation of AP9 model using POES and TACSAT4 satellites measurements
- Current version (1.20, July 2014) not available on SPENVIS yet
 - I. Updated flux maps for electrons/protons
 - II. Partial inclusion of data from Van Allen twin satellites to:
 - I. Study GeV protons using relativistic proton spectrometer (RPS) measurements
 - II. Separate temporal/spatial anomalies
- On going work to release Version 2 of AE9/AP9/SPM (2015 ???)