

Long Term Dose Monitoring Onboard the European Columbus Module of the International space Station in the frame of the DOSIS and DOSIS 3D Project-Results from the active Instruments

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# **DOSTEL Detector**



The DOSTEL Detecor Head consists two circular PIPS detectors by Canberra. These two are forming a telescope. The active area of the detectors is 6.93 cm<sup>2</sup>. Mounted in a distance of 1.5cm this leads to an opening angle of 120°.



2.97 cm

E 0

S

5

- Each DOSTEL can provide count rate profiles, dose rate profiles and energy deposition spectra

**DOSTEL** Data

- The energy spectra can be used to obtain LET-Spectra because of the path length limitation due to its telescope geometry
- The LET spectra can be used to get information such as average quality factors which leads to dose equivalent
- DOSTEL-1 has an additional mode 2 where the PHA data for every single particle is taken





### DOSTEL 1



DDPU – DOSTEL Data and Power Unit







DOSTEL



DDPU (DOSTEL Data and Power Unit)











# DOSIS - Launch 15.July 2009 STS-127







# DOSIS 3D - Launch 15. May 2012 Soyuz TMA 04M / 30S



# **DOSIS Installation Inside COLUMBUS**

1eap

![](_page_8_Figure_1.jpeg)

![](_page_9_Picture_0.jpeg)

### DOSIS & DOSIS 3D: DOSTEL

![](_page_9_Picture_2.jpeg)

![](_page_9_Picture_3.jpeg)

![](_page_9_Picture_4.jpeg)

![](_page_9_Picture_5.jpeg)

- Ethernet connection to EPM rack "Right Utility Distribution Panel"
- DOSIS MAIN BOX connected to EPM LAN like an external EPM instrument
- Data downlink is an EPM operation from ground performed once per month

# **DOSIS Main Box**

![](_page_10_Picture_4.jpeg)

![](_page_11_Picture_1.jpeg)

![](_page_11_Picture_2.jpeg)

ISS023E044747

# **Radiation Environment in Space**

![](_page_12_Figure_1.jpeg)

![](_page_12_Figure_2.jpeg)

Galactic Cosmic Rays (GCR)

lons from protons to iron

Trapped Radiation (Van Allen Belts)

 Low energy protons and electrons

Solar particle events

• Protons (in dependence of the solar cycle)

Variation of the radiation load onboard the ISS with altitude, latitude and time

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![](_page_13_Picture_1.jpeg)

![](_page_13_Picture_2.jpeg)

![](_page_13_Picture_3.jpeg)

![](_page_14_Figure_1.jpeg)

![](_page_15_Figure_1.jpeg)

![](_page_15_Figure_2.jpeg)

DoY 2009

![](_page_15_Picture_4.jpeg)

![](_page_15_Picture_6.jpeg)

![](_page_16_Figure_0.jpeg)

DoY 2012

![](_page_17_Figure_0.jpeg)

September Poland. Krakow. WRMISS.

2012

O

DoY 2011

DoY 2012

![](_page_18_Figure_1.jpeg)

# µGy/h in Si

CAU

![](_page_19_Figure_2.jpeg)

1eap

![](_page_19_Picture_3.jpeg)

# **Absorbed Dose Rates DOSTEL 1**

![](_page_20_Figure_1.jpeg)

# **Absorbed Dose Rates DOSTEL 2**

![](_page_21_Figure_1.jpeg)

![](_page_22_Picture_0.jpeg)

![](_page_22_Figure_1.jpeg)

![](_page_23_Picture_0.jpeg)

![](_page_23_Figure_1.jpeg)

![](_page_24_Picture_0.jpeg)

# LET Spectra DOSTEL-1 Detector 1

![](_page_24_Figure_2.jpeg)

# **Dose Equivalent Rates DOSTEL 1**

1ead

![](_page_25_Figure_1.jpeg)

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C | A

# **Dose Equivalent Rates DOSTEL 2**

lead

![](_page_26_Figure_1.jpeg)

![](_page_27_Picture_0.jpeg)

![](_page_27_Picture_1.jpeg)

# **DOSIS / DOSIS 3D Results**

DOSTEL-1: July 18, 2009 - May 28, 2010 DOSTEL-2: July 18, 2009 - June 16, 2011

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**C** |

DOSIS	Total				GCR				SAA			
	µGy/d in Si	µGy/d in tissue	Q	µSv/d	µGy/d in Si	µGy/d in tissue	Q	µSv/d	µGy/d in Si	µGy/d in tissue	Q	µSv/d
DOSTEL-1	195	240	2.5	609	130	159	3.2	510	66	81	1.22	99
DOSTEL-2	181	222	2.6	584	124	153	3.3	497	57	70	1.26	88

	Total				GCR				SAA			
DOSIS 3D	µGy/d in Si	µGy/d in tissue	Q	µSv/d	µGy/d in Si	µGy/d in tissue	Q	µSv/d	µGy/d in Si	µGy/d in tissue	Q	µSv/d
DOSTEL-1	236	291	2.2	627	117	144	3.1	451	119	147	1.20	176
DOSTEL-2	208	256	2.3	583	113	139	3.1	433	95	116	1.29	150

DOSTEL-1: May 21, 2012 – August 13, 2014 DOSTEL-2: May 21, 2012 – August 19, 2014

![](_page_28_Picture_0.jpeg)

### **GCR Dose Rates**

![](_page_28_Figure_2.jpeg)

![](_page_28_Figure_3.jpeg)

![](_page_29_Picture_0.jpeg)

### SAA Dose Rates

![](_page_29_Figure_2.jpeg)

![](_page_29_Figure_3.jpeg)

![](_page_29_Figure_4.jpeg)

![](_page_30_Picture_1.jpeg)

![](_page_30_Picture_2.jpeg)

![](_page_30_Picture_3.jpeg)

![](_page_31_Picture_0.jpeg)

10<sup>0</sup>

 $\Delta E / MeV$ 

10<sup>1</sup>

10<sup>2</sup>

10<sup>3</sup>

32

10<sup>-6</sup>

10<sup>-1</sup>

![](_page_32_Picture_0.jpeg)

DOSTEL 2 D1 SAA ascending DOSTEL 2 D2 SAA descending DOSTEL 2 D2 SAA ascending

10<sup>0</sup>

 $\Delta E / MeV$ 

10<sup>1</sup>

10<sup>2</sup>

10<sup>3</sup>

10<sup>-1</sup>

33

CAU

flux / arb. units

10<sup>-5</sup>

10<sup>-6</sup> 10-2 2014

19th WRMISS, Krakow, Poland, September 9

# DOSIS 3D SAA Mode 2 Data

![](_page_33_Figure_3.jpeg)

![](_page_33_Figure_4.jpeg)

![](_page_34_Picture_0.jpeg)

**DOSIS 3D Mode 2 Data** 

геар

GCR (mainly)

![](_page_34_Figure_4.jpeg)

### SAA (mainly)

![](_page_34_Figure_6.jpeg)

![](_page_34_Figure_7.jpeg)

![](_page_35_Picture_0.jpeg)

DOSIS 3D Mode 2 Data

*1еар* 

Ascending SAA (mainly)

![](_page_35_Figure_4.jpeg)

Descending SAA (mainly)

![](_page_35_Figure_6.jpeg)

![](_page_36_Picture_0.jpeg)

# **Conclusions**

- Up to now we have data for 1395 days of measurement with DOSTEL-2 (DOSTEL-1: 1046d)
- During the passes of the SAA region stopping Protons can clearly be seen in the second detector of DOSTEL-1, i.e. protons mainly reaching the DOSTEL-1 telescope from aft direction
- Contribution of GCR to the dose rate is obviously correlated with the GCR flux measured by NM while the contribution of trapped particles is correlated with the stations altitude

С

# **Acknowledgements**

- Special Thanks to the ICCHIBAN Workinggroup and the HIMAC facility for a lot of support

Supported by:

![](_page_37_Picture_5.jpeg)

Federal Ministry of Economics and Technology under grants 50WB0826, 50WB1026, and 50WB1232

on the basis of a decision by the German Bundestag

Thank you very much for your attention!

![](_page_38_Picture_1.jpeg)

![](_page_38_Picture_2.jpeg)

![](_page_39_Picture_0.jpeg)

![](_page_39_Picture_1.jpeg)

### **DOSIS 3D: 38S**

![](_page_39_Picture_3.jpeg)

**DOSIS: 2009** 

![](_page_40_Picture_1.jpeg)

![](_page_40_Picture_2.jpeg)

![](_page_41_Picture_1.jpeg)

![](_page_42_Picture_0.jpeg)

![](_page_42_Picture_1.jpeg)

![](_page_43_Picture_1.jpeg)

![](_page_44_Picture_1.jpeg)

![](_page_45_Picture_1.jpeg)

![](_page_45_Picture_2.jpeg)

![](_page_46_Picture_1.jpeg)

![](_page_46_Picture_2.jpeg)

CAU

![](_page_47_Picture_1.jpeg)

### **Absorbed Dose Rates SAA**

![](_page_47_Figure_3.jpeg)

# C A U ISS configuration

![](_page_48_Figure_1.jpeg)

![](_page_48_Figure_2.jpeg)

![](_page_48_Figure_3.jpeg)

### Service Module

![](_page_48_Picture_5.jpeg)

![](_page_48_Picture_6.jpeg)

### Thanks to S.G. Drobyshev, V.V. Benghin

![](_page_48_Figure_8.jpeg)

![](_page_49_Figure_1.jpeg)

![](_page_50_Figure_0.jpeg)

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![](_page_51_Picture_1.jpeg)

# **Dose Values During Shuttle Docking**

- The mean contribution to the daily dose drops roughly 30-40% from ~100  $\mu Sv/d$  to ~60-70 $\mu Sv/d$  during the Shuttle docking phases
- The effect occurs mainly during the south-eastward passages through the SAA
- It is most likely due to the attitude change of the station from +Xvv to -Xvv

![](_page_52_Picture_1.jpeg)

![](_page_52_Picture_2.jpeg)

![](_page_52_Picture_3.jpeg)

[European Users Guide to Low Gravity Platforms, UIC-ESA-UM-0001]

![](_page_53_Picture_0.jpeg)

# **Mean Quality Factor**

1eap

![](_page_53_Figure_2.jpeg)

54

![](_page_54_Picture_0.jpeg)

![](_page_54_Figure_1.jpeg)

DoY 2001/2009/2011

μGy/h in Si

![](_page_55_Picture_0.jpeg)

# LET Spectra DOSTEL-1 Detector 2

![](_page_55_Figure_2.jpeg)

![](_page_56_Figure_1.jpeg)

keV/ $\mu$ m (H<sub>2</sub>O)

геар

![](_page_57_Figure_1.jpeg)

keV/µmH2O

![](_page_58_Picture_0.jpeg)

# **DOSTEL-2 Count Rate Profile**

![](_page_58_Figure_2.jpeg)

DoY 2009

![](_page_59_Picture_0.jpeg)

# **DOSTEL-2 Count Rate Profile**

![](_page_59_Figure_2.jpeg)

DoY 2009

![](_page_60_Figure_1.jpeg)

![](_page_60_Figure_2.jpeg)

ieap

![](_page_60_Picture_4.jpeg)

![](_page_61_Picture_1.jpeg)

![](_page_61_Picture_2.jpeg)

### **DOSIS & DOSIS 3D: Scientific Goals**

The main objective of the **DOSIS & DOSIS 3D** experiment is the determination of the absorbed dose and dose equivalent using a variety of active and passive radiation detector devices distributed throughout the ISS.

- Monitor the radiation environment inside Columbus with active and passive radiation detectors (ESA) for the determination of the temporal and spatial dose distribution
- Combine data gathered by NASA, JAXA, IMBP and ESA into a 3D radiation map of the International Space Station

![](_page_61_Picture_7.jpeg)

![](_page_61_Picture_8.jpeg)

![](_page_62_Picture_1.jpeg)

![](_page_62_Picture_2.jpeg)

Thermoluminescence detectors (TLD)

First usage of LiF (Lithiumfluoride) for the measurement of radiation following an atomic weapon test

Measurement of internal radiation dose received by cancer patients treated with radioactiv isotopes at Oak Ridge Institute for Nuclear **Studies** 

**F. Daniels** Science **117**, 343, **1953** 

![](_page_62_Picture_7.jpeg)

**Nuclear Track Etch Detectors (CR-39)** 

Material : CR-39 = allyl diglycol carbonate

Heavy charged particles break chemical bonds in the material. This trail can be made visible by etching the material.

R. P. Henke and E. V. Benton. Nucl.Instr.Meth. 97 (1971) 483-9

The combination of passive thermoluminescence detectors and nuclear track etch detectors allows to determine the absorbed dose (in Gray) and the dose equivalent (in Sievert).

# **DOSIS Passive Detectors**

11 passive radiation detector packages (PDP) distributed within Columbus

![](_page_63_Picture_3.jpeg)

**DOSIS** Passive Detector Packages (PDP)

![](_page_63_Picture_5.jpeg)

![](_page_63_Picture_6.jpeg)

64

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C | A

# DOSIS / DOSIS 3D

![](_page_64_Picture_2.jpeg)

Agency	Name of the detectors	Quantity	Position on ISS				
ESA	PDP	10 + 1	Columbus (red rectangle)				
NASA	RAMs	24	All over the ISS (yellow rectangle)				
JAXA	PADLES	12	KIBO (blue rectangle)				
IBMP	SPD	6	Russian part of the ISS (green rectangle)				
IDMP	Pille	10					

Fig. 3: Passive detectors on the ISS. The RAMs, PADLES, SPD and Pille detectors are permanently on board of the ISS. The PDPs from ESA are not. To reach the scientific goals set by Dosis 3D the PDPs shall be re-introduced on Columbus.