

Recent Results from the MSL–RAD Experiment on the Curiosity Mars Rover

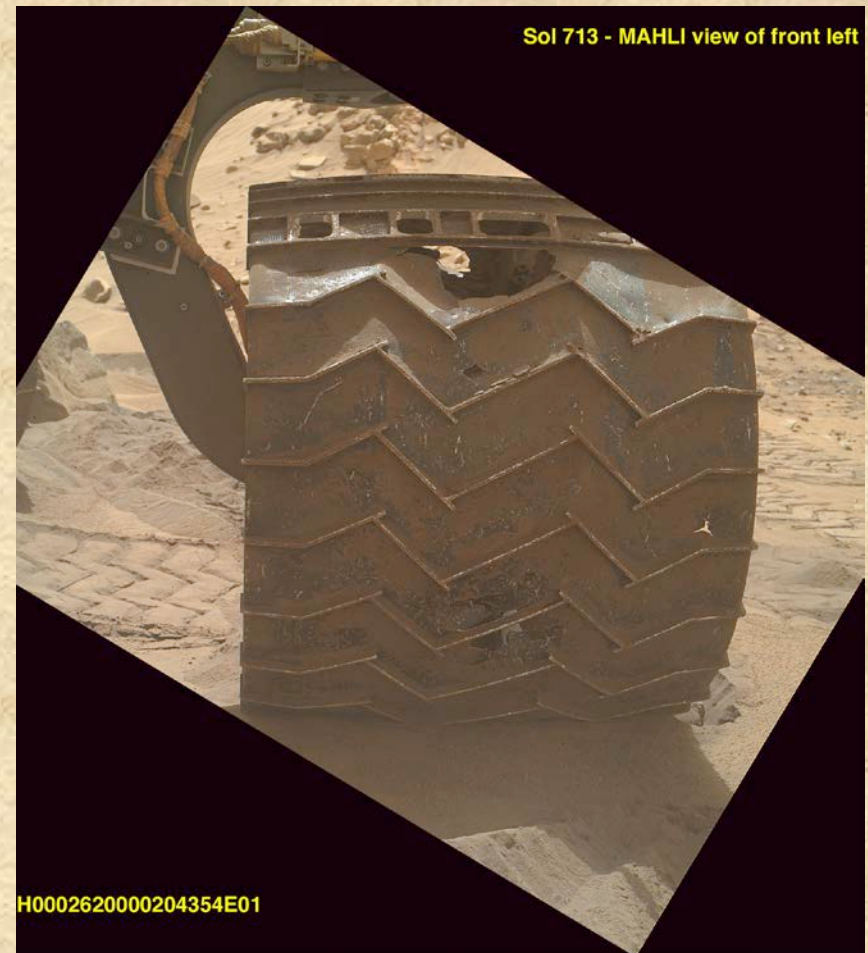
Cary Zeitlin, Southwest Research Institute
On behalf of the MSL–RAD Science Team

RAD Science Team

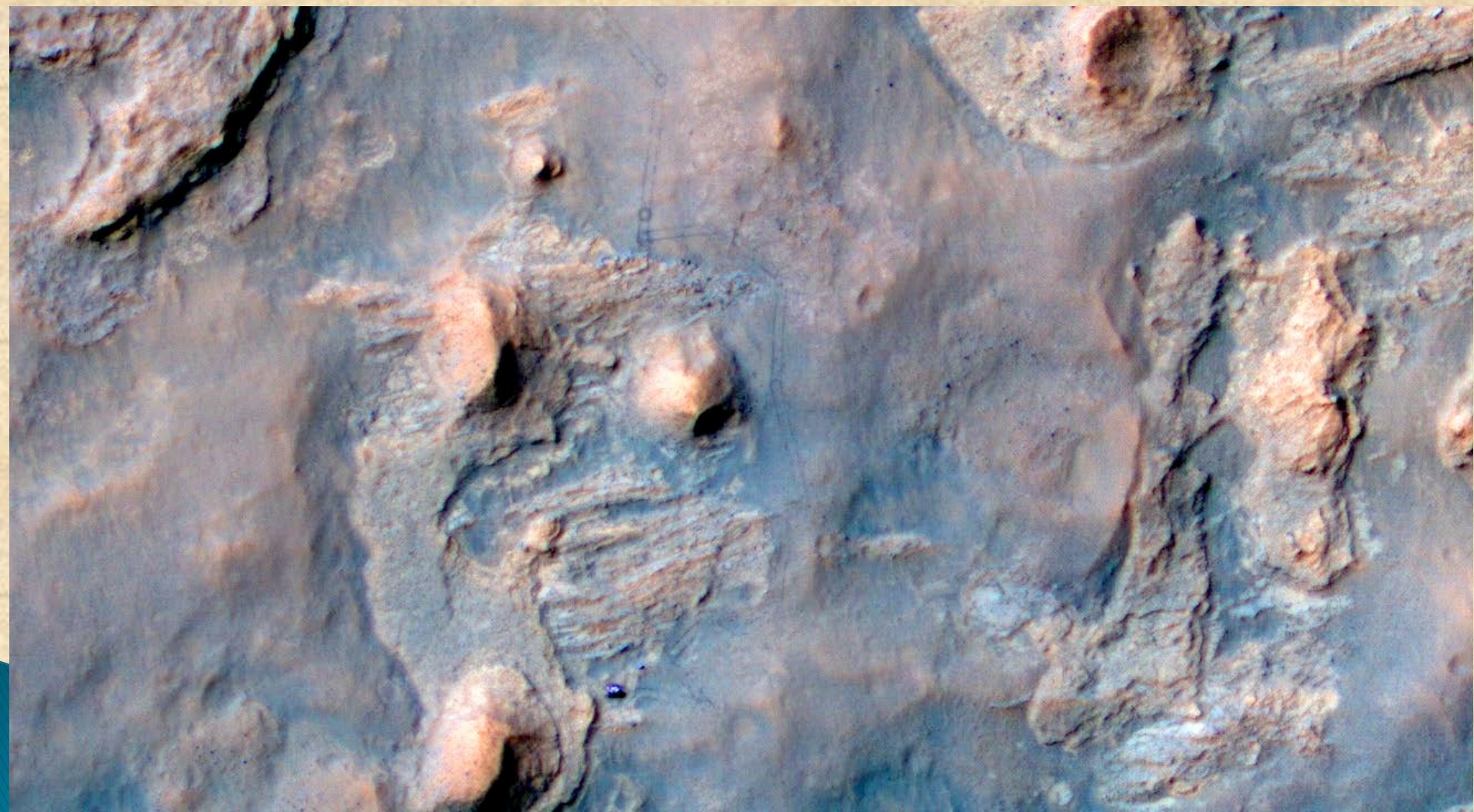
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 - ⁵University Space Research Associates
 - ⁶NASA Headquarters
 - ⁷German Aerospace Center

Curiosity Mission Update

- ▶ Reached base of Mt. Sharp yestersol.
- ▶ On schedule – goal was to reach Mt. Sharp at end of prime mission.
- ▶ Driving slower than expected at first, surprising wheel damage.



Path As Seen From Orbit

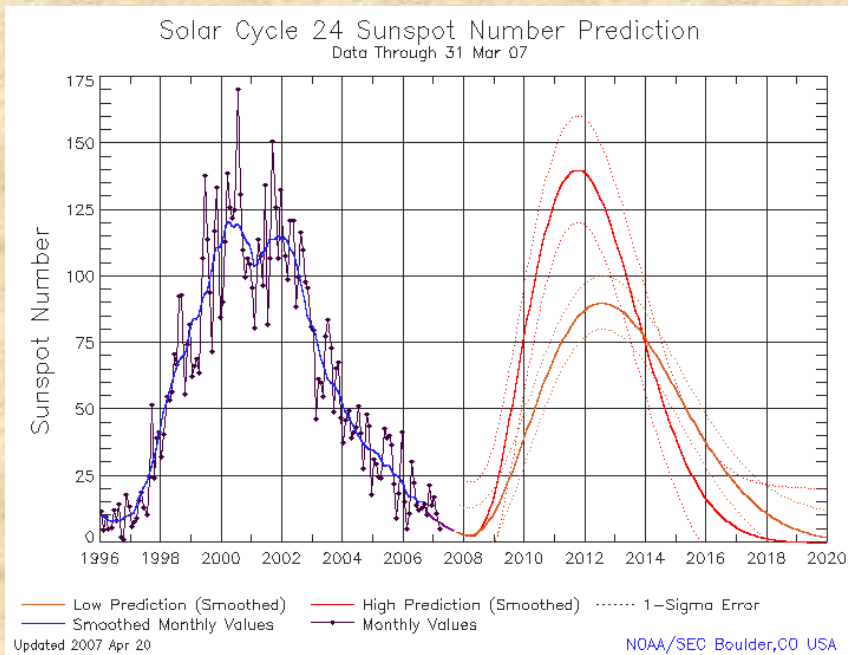


Path As Seen From Orbit

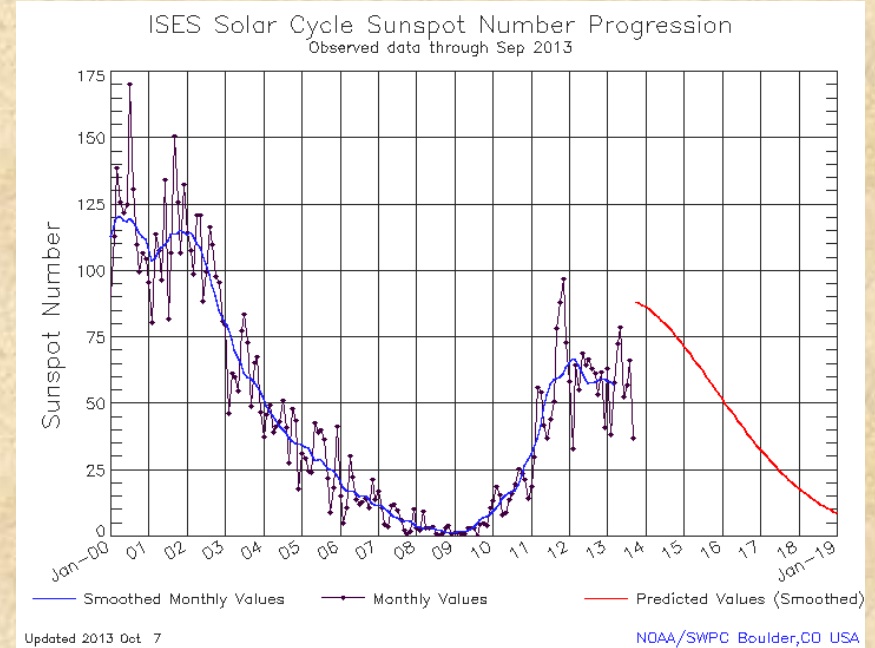


Solar Cycle Predictions

Cycle 23 Data + 2007 Predictions



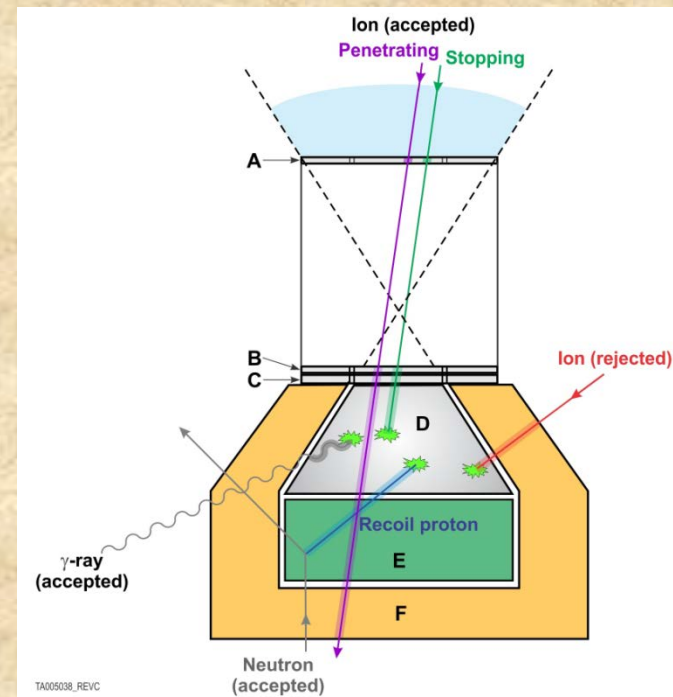
Actual Cycle 24



Cycle 24 is weakest in ~ 100 years (not predicted).

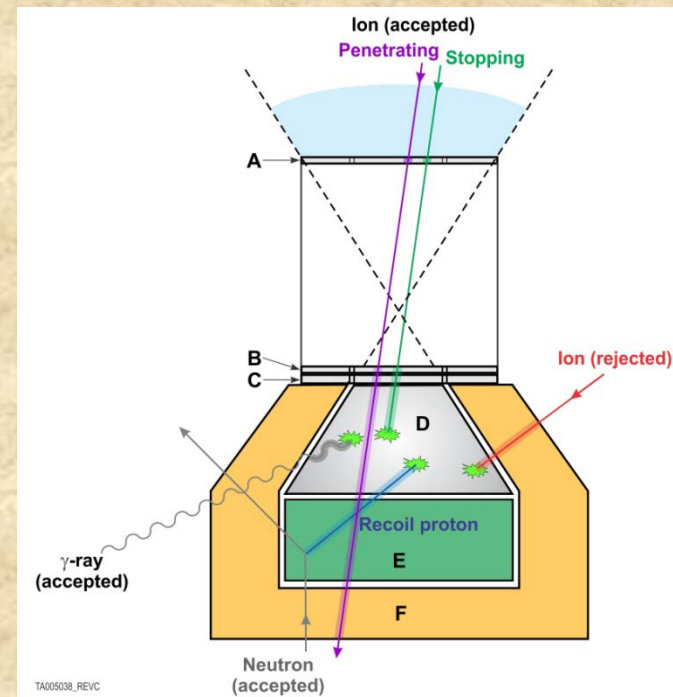
MSL-RAD

- ▶ Silicon detector telescope with 3 elements (A, B, C).
- ▶ CsI scintillator = D.
- ▶ Plastic scintillators: E = 1.8 cm, F = 1.2 cm.
- ▶ F = anticoincidence, upper (F1) and lower (F2).
- ▶ D & E each have 3 readout photodiodes attached.

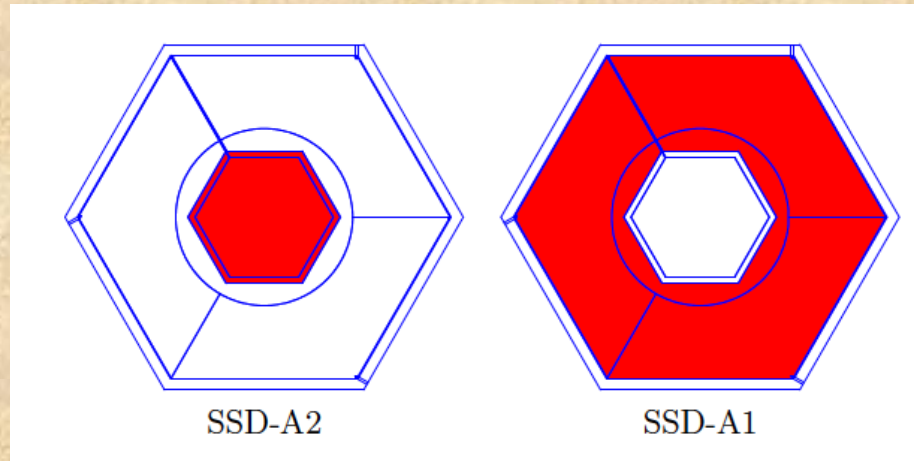


MSL-RAD

- ▶ Scintillator readout diodes are used in coincidence in triggers (avoids triggering on γ -rays that make a direct hit in diodes).
- ▶ $DH * DM * !F * !C = \text{neutral}$
- ▶ $EH * EM * !F * !C = \text{neutral}$
- ▶ $EH * EM = E \text{ dosimetry}$
- ▶ $BU = B \text{ dosimetry}$
- ▶ Dosimetry triggers accept omnidirectional radiation.

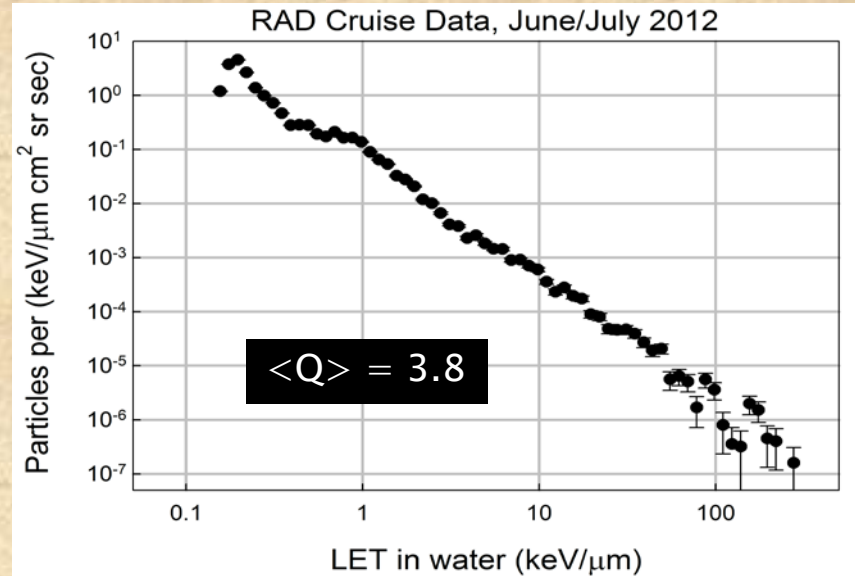
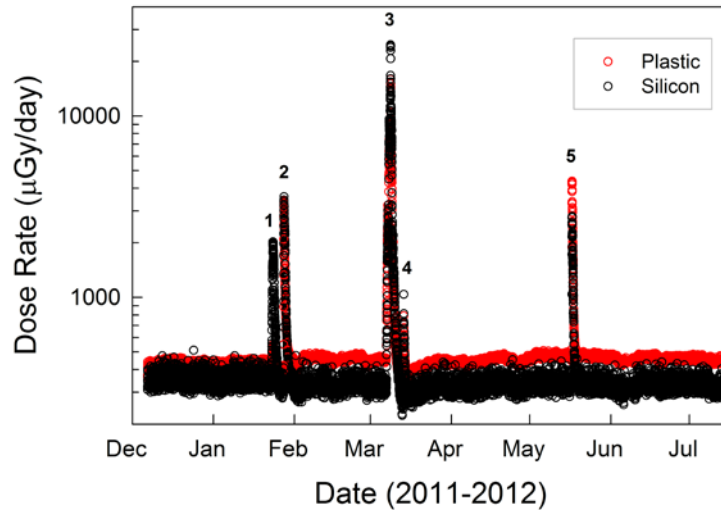


Diode Segmentation & FOV



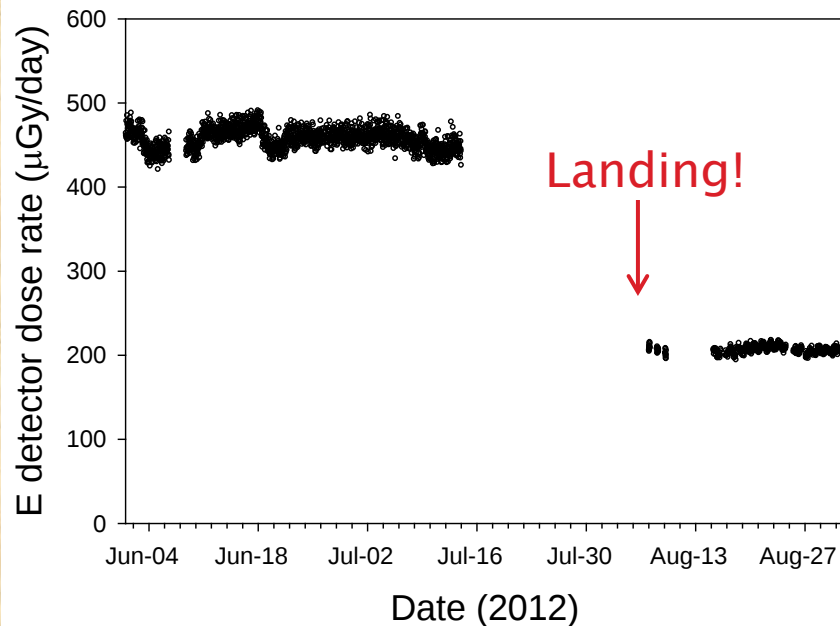
- ▶ A2, B, C use inner segment of diodes, A1 uses outer.
- ▶ Two fields of view, two geometry factors.
- ▶ A2*B cone has half-angle $\sim 18^\circ$, $G=0.17 \text{ cm}^2 \text{ sr}$.
- ▶ A1*B cone $\sim 30^\circ$, $G=0.72 \text{ cm}^2 \text{ sr}$.
- ▶ Use A2*B events for LET spectrum.

RAD Cruise Results



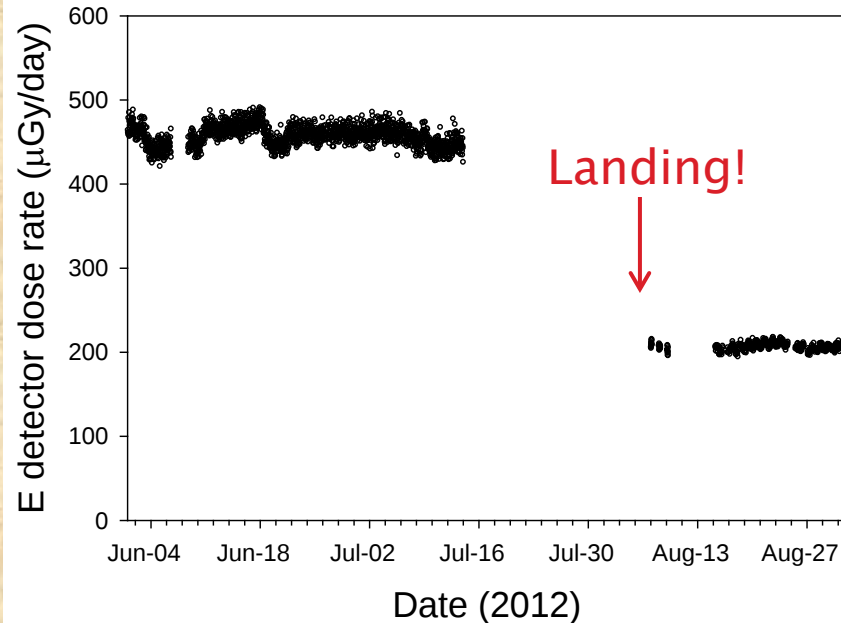
- ▶ Tissue dose rate = $0.48 \pm 0.08 \text{ mGy/day}$
- ▶ Dose equivalent rate = $1.8 \pm 0.3 \text{ mSv/day}$
- ▶ SEP event contribution ~ 14 days of GCR.

RAD Surface Results

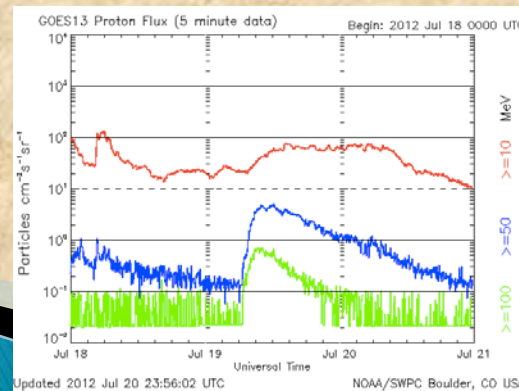
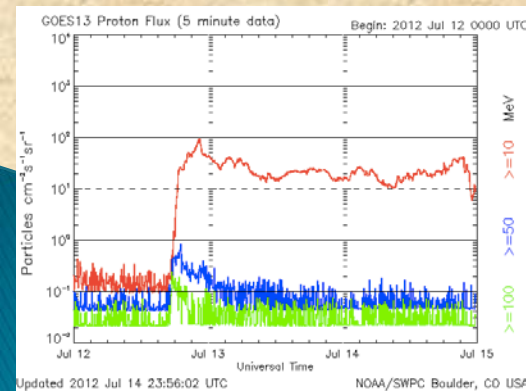


- ▶ Average E dose rate $\sim 210 \mu\text{Gy/day}$, $\sim 40\%$ of the cruise dose rate.
- ▶ On an airless body, expect dose rate to be 50% of free space.
- ▶ $\langle Q \rangle = 3.05 \pm 0.05$, $\sim 30\%$ lower on surface than in cruise.
- ▶ Atmospheric shielding & increased modulation decrease dose rates compared to cruise.

RAD Surface Results

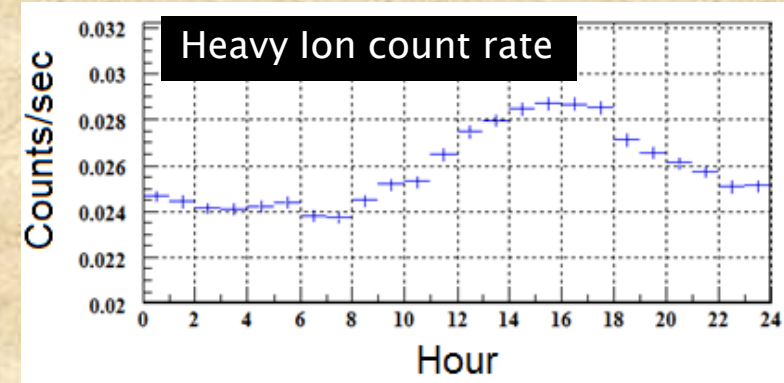
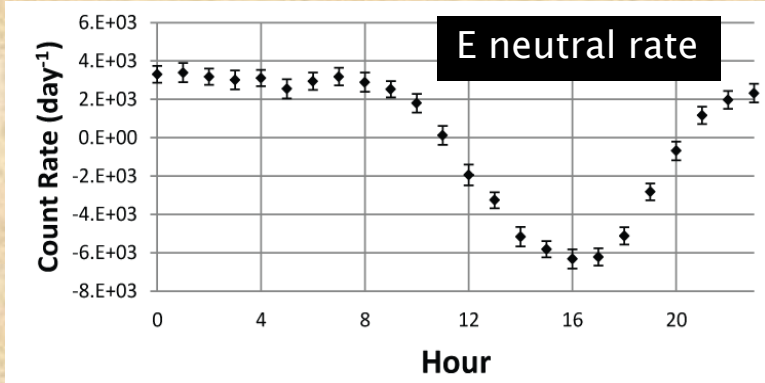
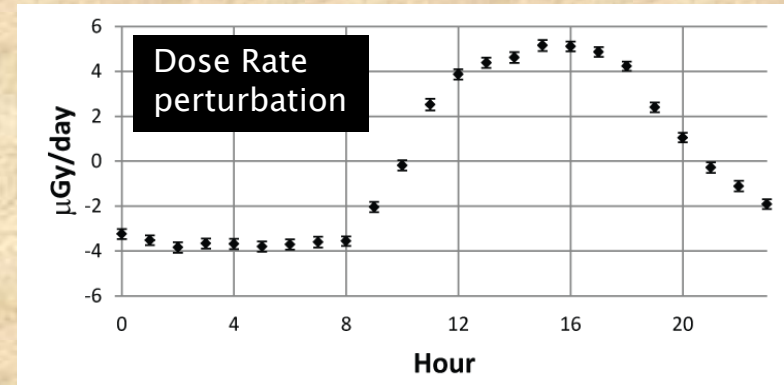
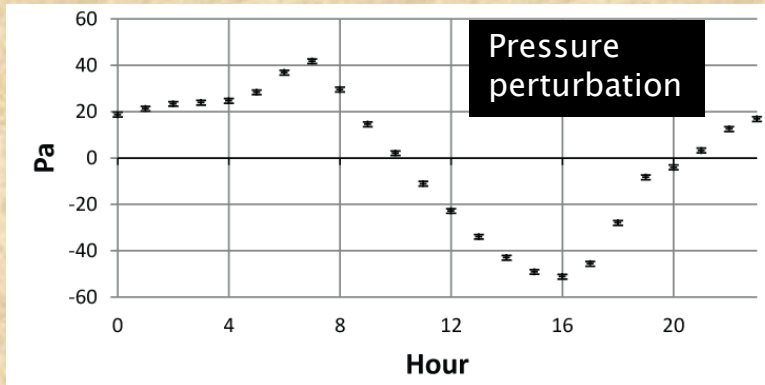


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Diurnal Variations

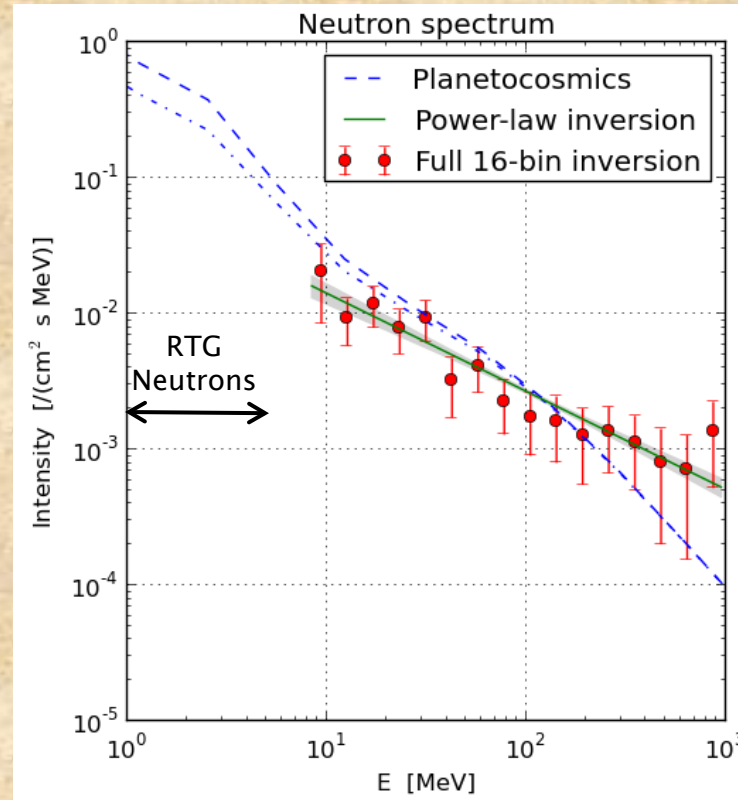
S. Rafkin



- ▶ “Thermal tide” $\rightarrow$ $\pm$ 5% daily variations in pressure $\rightarrow$ $\pm$ 2% variation in radiation dose rate, inversely correlated with P.
- ▶ Thinner atmosphere $\rightarrow$ fewer neutrons made + more heavy ions survive traversal $\rightarrow$ higher dose rate.
- ▶ Pressure data from the REMS team.

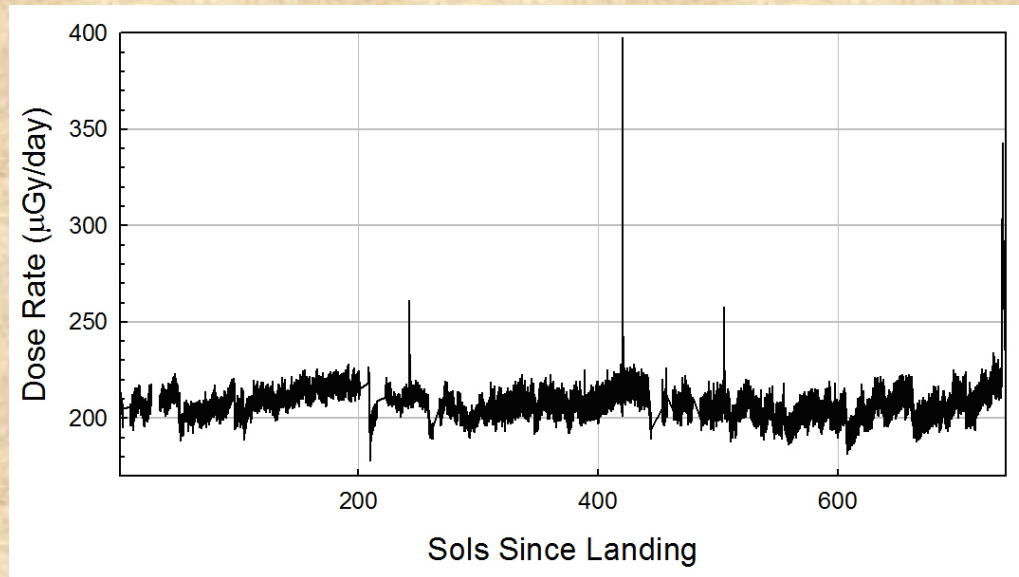
Neutron Spectrum

J. Köhler



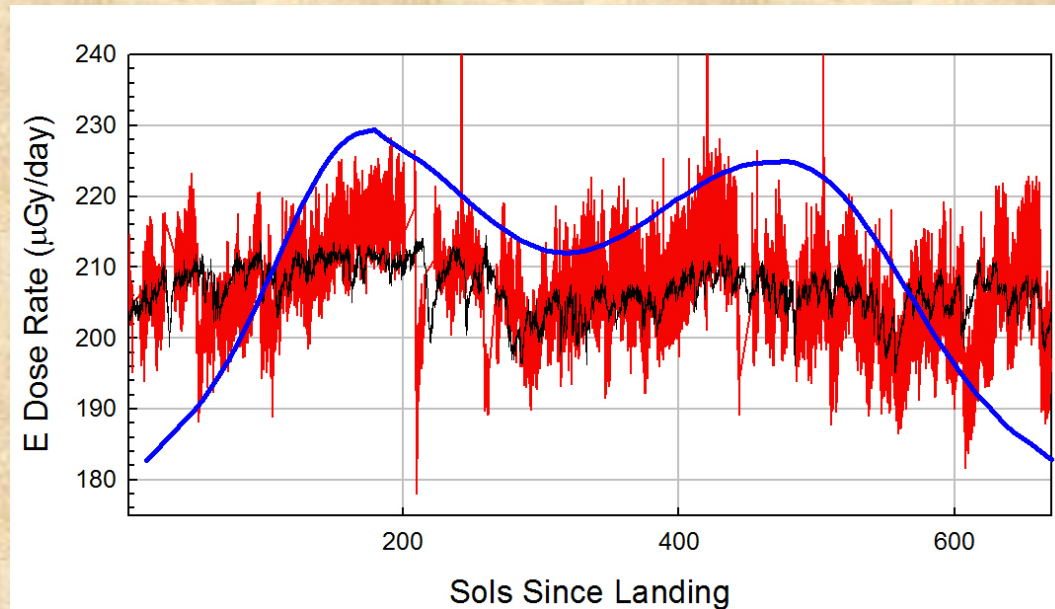
- ▶ D, E spectra inverted $\rightarrow \gamma$ and neutron spectra.
 - ▶ Neutron threshold energy ~ 8 MeV.
- ▶ $D = 14 \pm 4 \mu\text{Gy/day}$, about 7% of total.
- ▶ $H = 61 \pm 15 \mu\text{Sv/day}$, about 9% of total.
- ▶ E dose rate from RTG $< 1 \mu\text{Gy/day}$ in ground test.

E Dose Rate



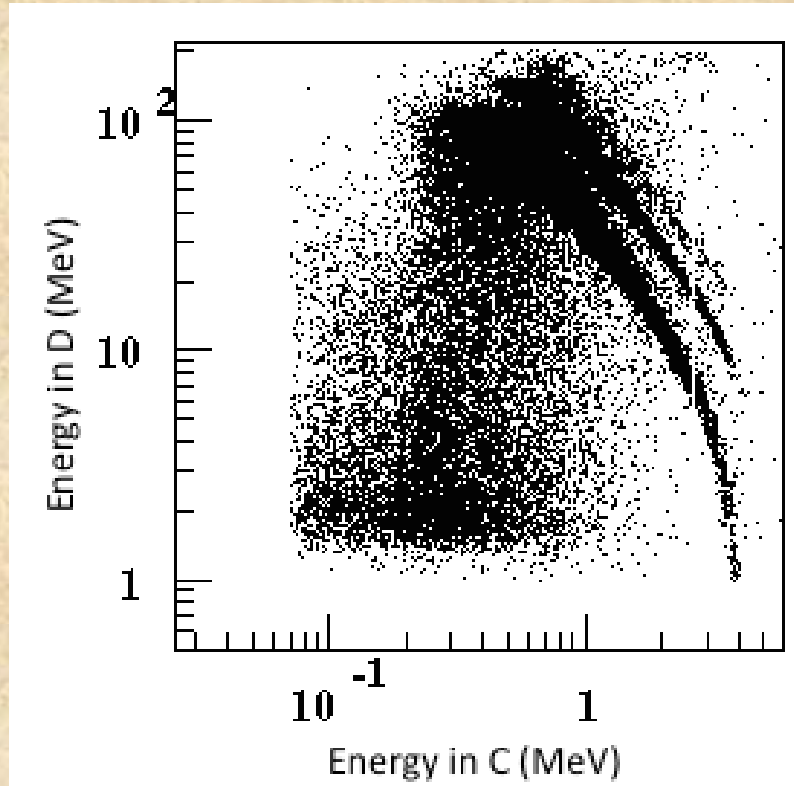
- ▶ Four small solar events, including one last week. (Another one on the way??)
 - RAD under avg. $21 \text{ g cm}^{-2} \text{ CO}_2 \rightarrow E_{\text{proton}} > 160 \text{ MeV}$
- ▶ Many Forbush decreases.
- ▶ SEP contribution to total dose $\sim$ negligible.

E Dose Rate – GCR



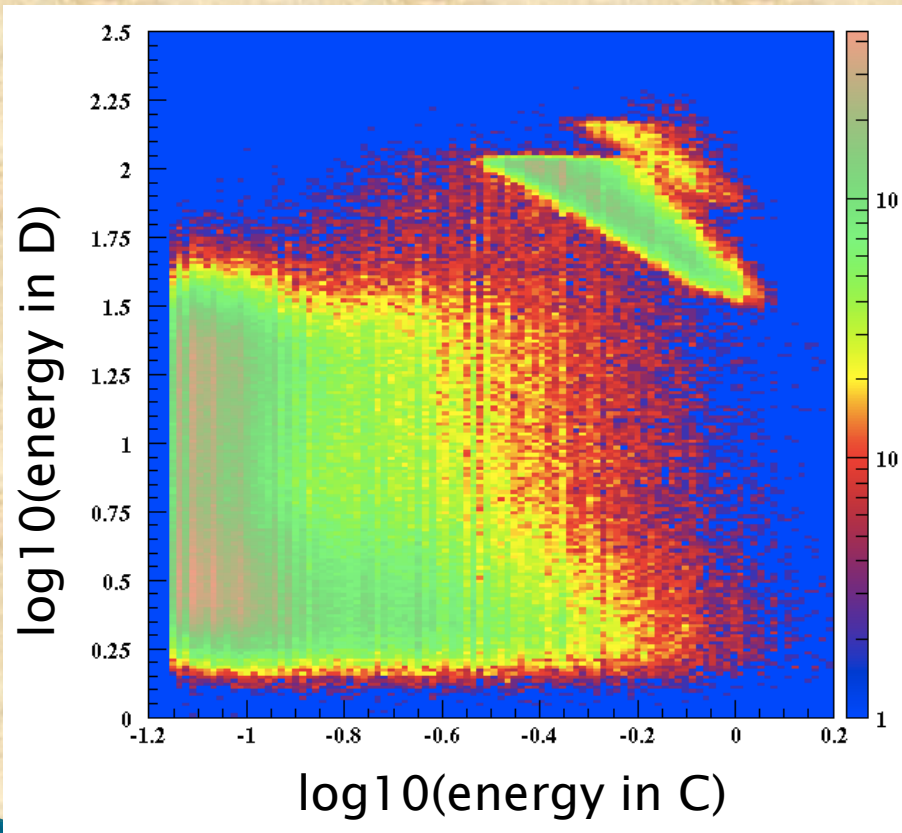
- ▶ Look for correlations w/seasonal atmospheric changes (scaled column depth in blue) and with heliospheric changes (scaled Oulu NM count rate in black).
- ▶ More influenced by heliosphere.
- ▶ Thanks to REMS team for pressure data.

$Z = 1$ Electrons & H Isotopes



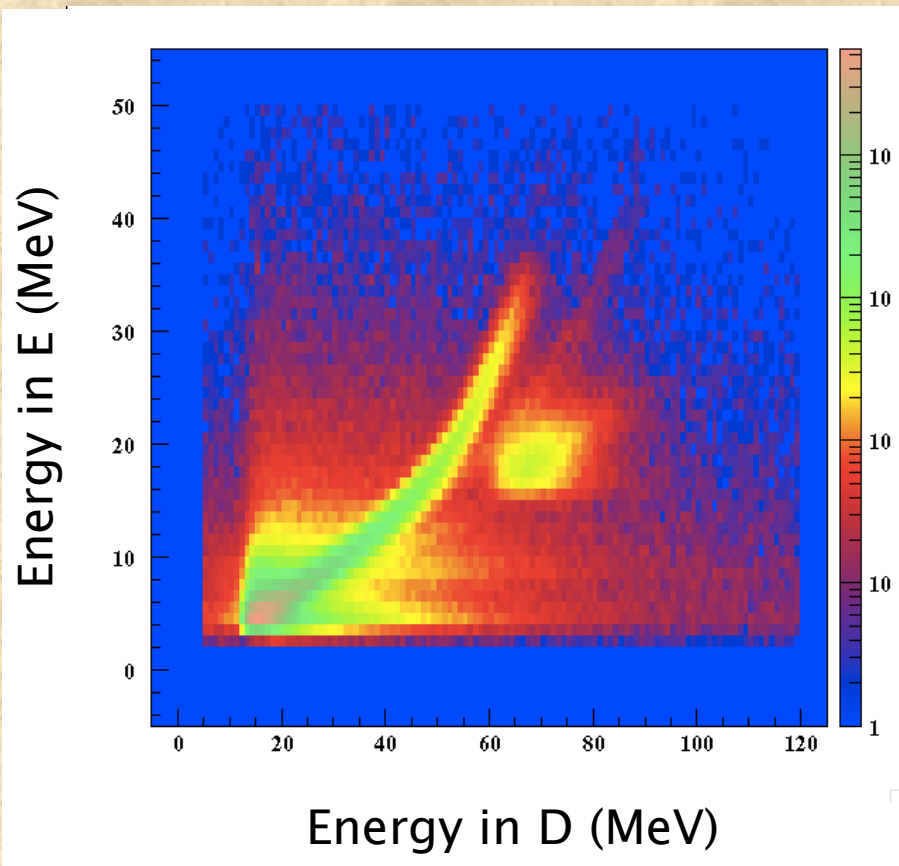
- ▶ Calorimetry useful for particle id.
- ▶ Select slow $Z=1$ particles that stop in D: hits in A2, B, C, D, but no energy in E or F2.
- ▶ See p, d, t.
- ▶ Electrons below the proton band.

More Stopping $Z=1$



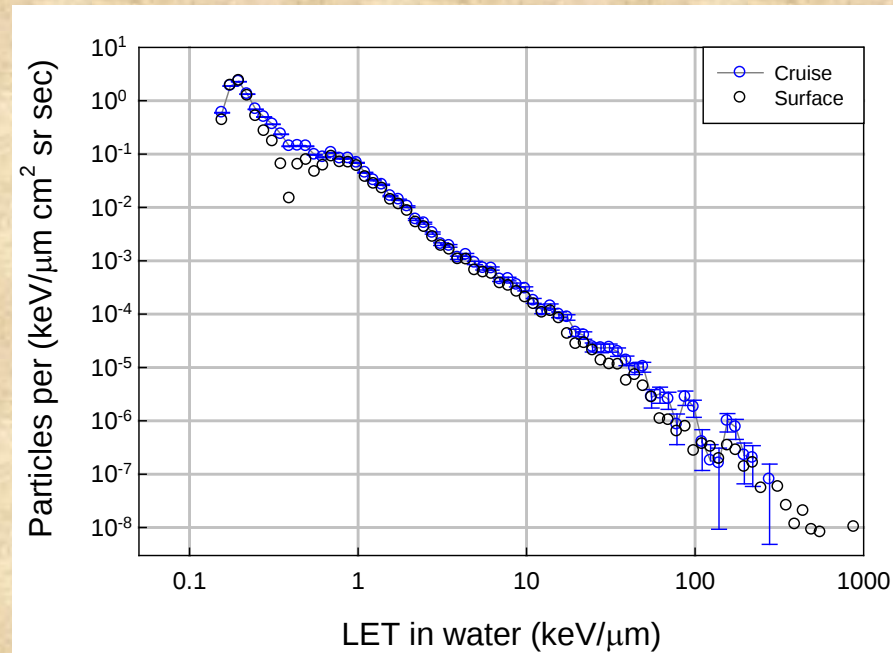
- ▶ D vs. C again but now include min-I in ABC.
- ▶ See low energy electrons and maybe pions stopping in D.

Penetrating $Z=1, 2$



- ▶ E vs. D, require $Z=1$ or 2 in ABC + energy in F2.
- ▶ See high-energy protons, deuterons, helium.
- ▶ Use to calculate integral fluxes.

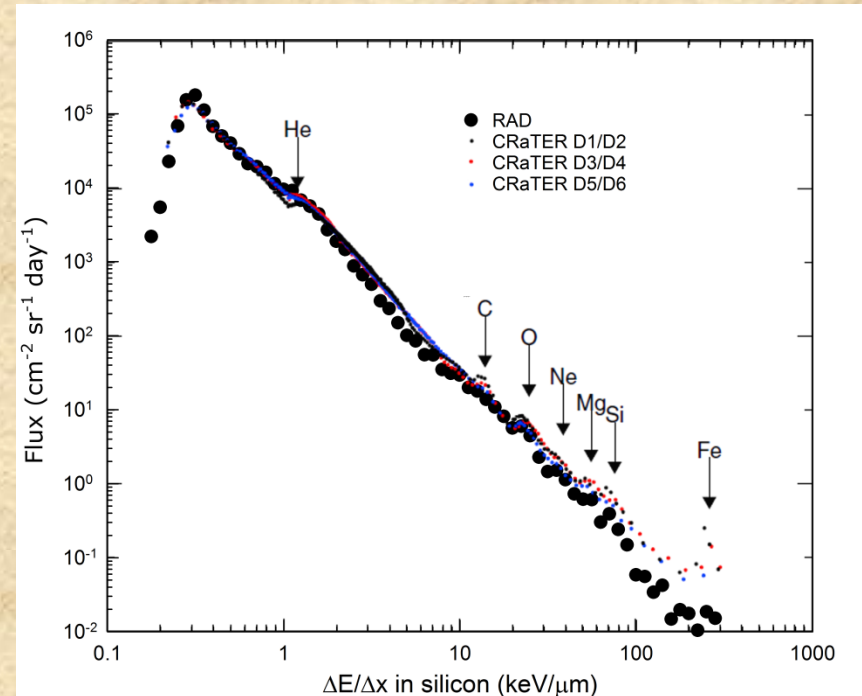
Cruise vs. Surface LET



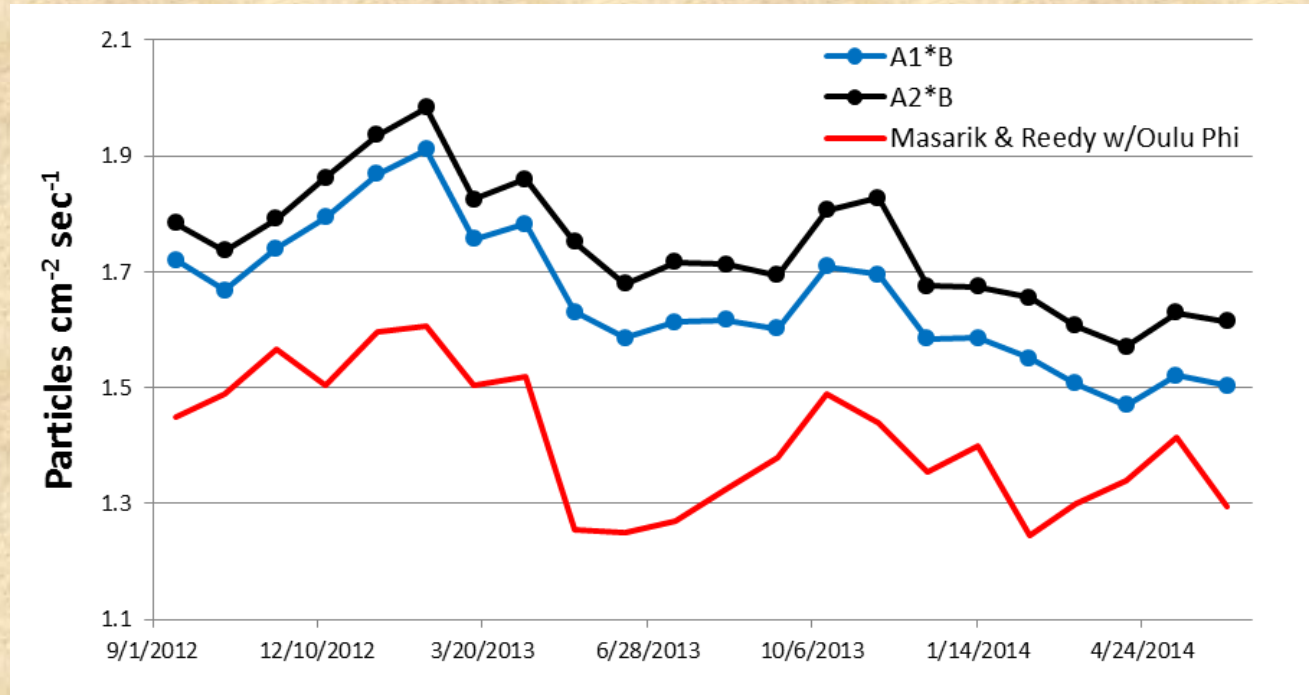
- ▶ Both spectra made with A2*B events.
- ▶ RTG background subtracted from both but less certain for surface.

RAD LET vs. CRaTER LET

- ▶ Compare silicon LET spectra ($\sim 21 \text{ g cm}^{-2} \text{ CO}_2$ shielding) to LET spectra from CRaTER with 0.2, 6, and 9 g cm^{-2} shielding.
- ▶ RAD sees slightly more min-I charge 1 particles and fewer heavies, as expected.



Flux Model w/Modulation



- ▶ Get Φ from Oulu NM count rates, use as input to Masarik & Reedy GCR flux model (red), compare to coincidence rates in A1*B and A2*B.
- ▶ Model is top of the atmosphere flux.

Summary and Conclusions

- ▶ RAD made the first measurement of radiation dose on a transit to Mars and continues to work well on the surface.
 - Diurnal and seasonal variations observed.
 - First SEP events observed on another planet.
 - Mars dose rate predictions span a factor of ~ 4 , from about $\frac{1}{2}$ of what RAD is measuring to about a factor of 2 higher.
 - For model validation, need to study spectral details & extend comparisons to include both flux models and transport models.

Acknowledgments

- ▶ RAD is supported at Southwest by NASA Advanced Exploration Systems.
- ▶ DLR supports Kiel team.
- ▶ JPL manages the MSL mission.
- ▶ Thanks to all for great support.