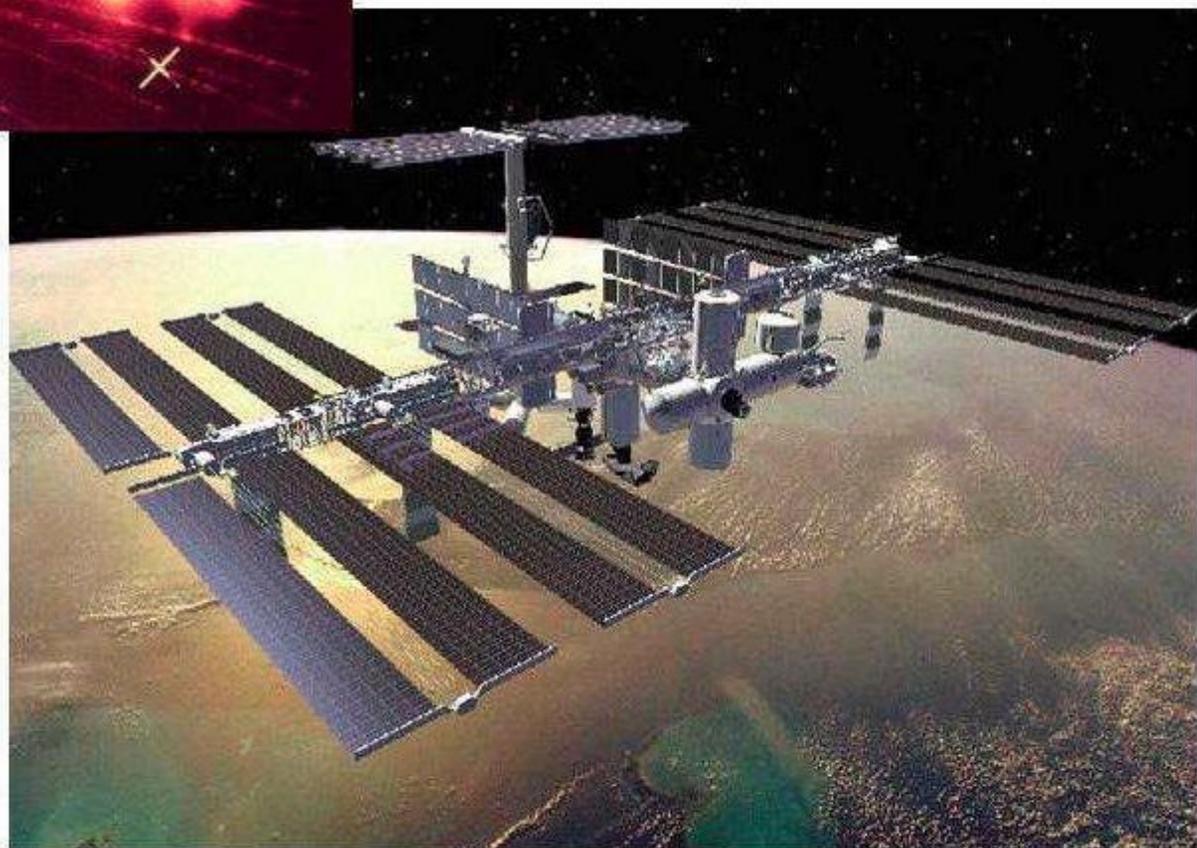
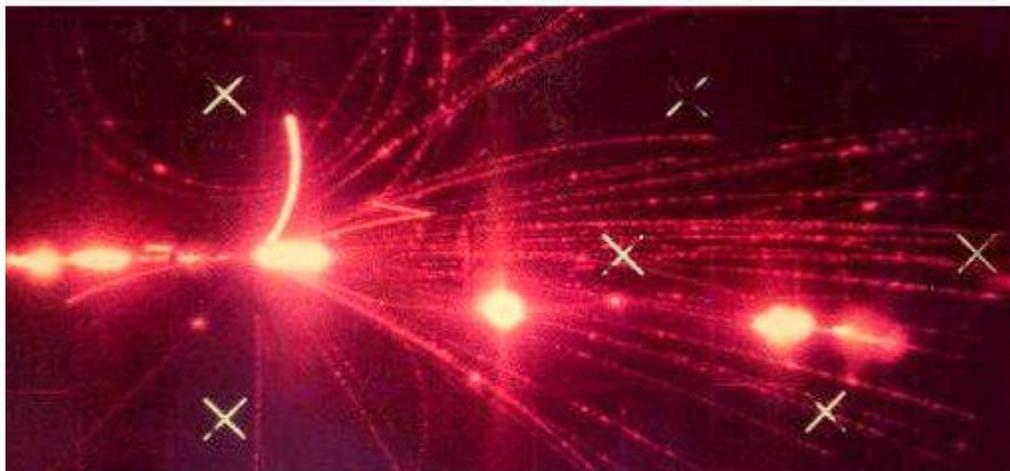


Measurements Consortium NASA Space Radiation Shielding Program



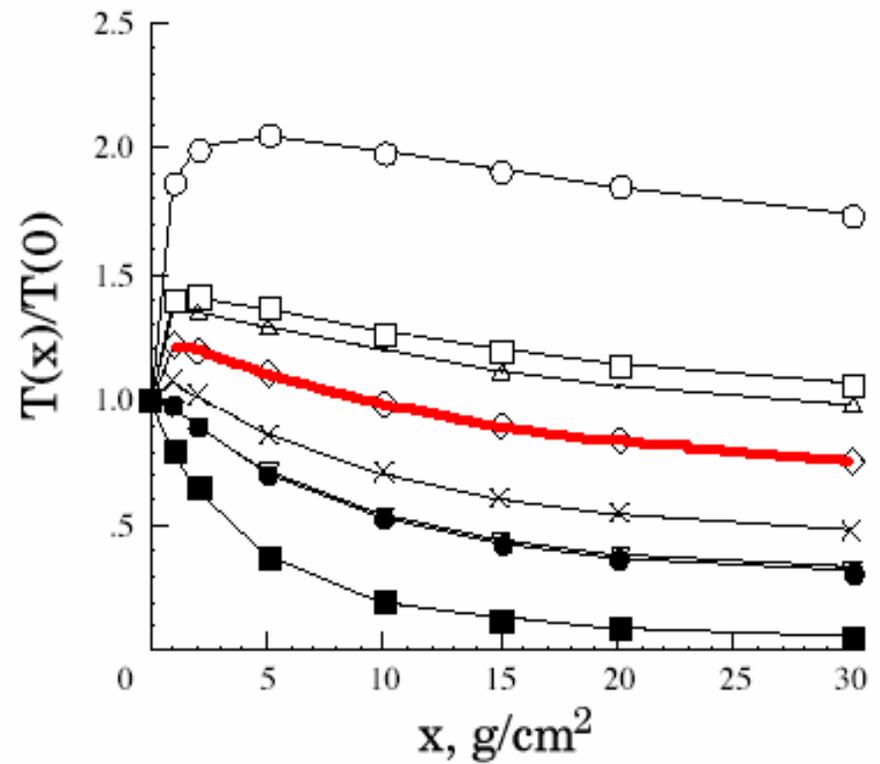
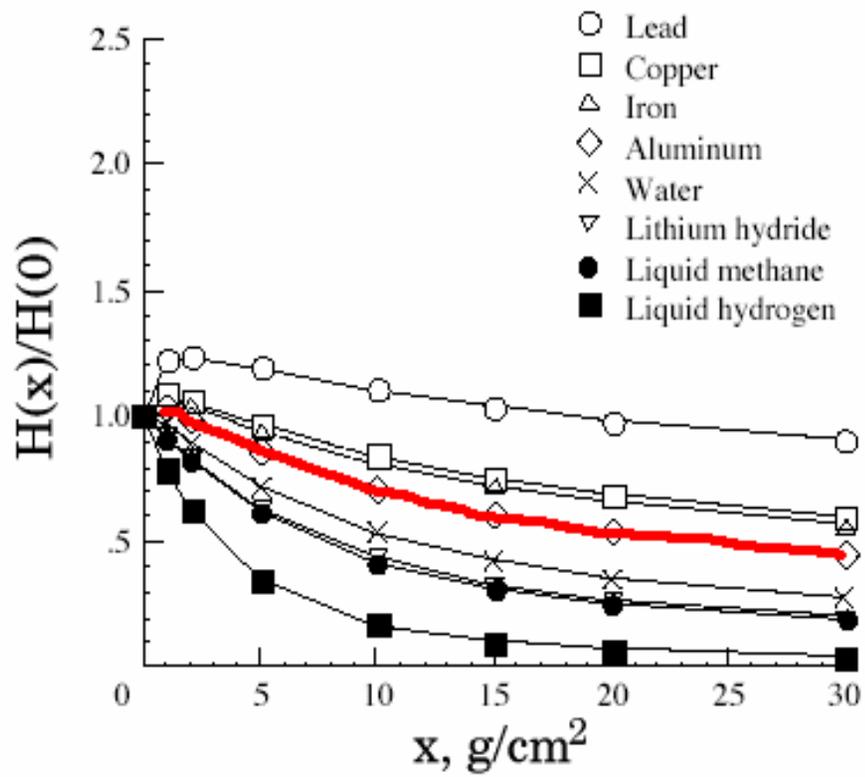
NSCL



CHALMERS



Chalmers University of Technology, Sweden



- For some materials, modest thicknesses (5-10 g-cm⁻²) may *increase* risk from GCR (Wilson *et al.*)

NASA-SRSP Objectives:

- **Use model calculations and experiments at particle accelerators to evaluate the radiation transport properties of potential spacecraft and habitat materials as a function of composition and thickness.**
- **From the accelerator measurements, develop a database of nuclear fragmentation data for use as model source terms and benchmarks.**
- **Ultimately, develop the models into a design tool for use by mission planners, spacecraft and instrument designers, flight surgeons and radiation health specialists.**

Berkeley Workshop Recommendations:

1. Use existing cross section data sets to ascertain the accuracy of the present radiation transport models.
2. Measure a recommended list of nuclear cross sections.
3. Use these new measurements and improved cross section calculations to improve radiation transport codes.
4. Incorporate measured cross section data into transport models under development. Establish a feedback mechanism with the model developers.
5. Determine whether models are sufficiently precise, by means of both ground-based measurements and in the natural cosmic ray environment using a suite of instruments to measure the radiation field with and without shielding.

Measurements Consortium Mission:

To provide the data to make the models as accurate and precise as required.

- **Nuclear fragmentation cross sections** are needed to ensure that the models get the underlying physics right, and **radiation transport measurements with proposed shielding materials** are needed to validate the models and to evaluate promising materials.
- The Consortium works with the SRSP Transport Consortium (PI, Prof. L Townsend) to assign priorities for the data gathering and analysis, and is pursuing new detector methods to optimize the use of accelerator resources.

Measurements Program

Fundamental Physics Needed for Code Development

- Cross sections for charged-particle production in HZE reactions using elemental targets (Zeitlin)
- Cross sections for neutron production in HZE reactions (Heilbronn)

Shielding Materials Evaluation

- Heavy charged particle fragmentation and transport in proposed shielding materials (Guetersloh, Zeitlin)

Instrument Responses

- TEPC (with T. Borak *et al.*)
- PNTD (Eril Research)
- ICCHIBAN

Quantities to be measured:

- **inclusive single and double differential cross sections** as a function of particle charge energy and emission angle
- **semi-inclusive cross sections** over large solid angle
- **particle yields** as a function of projectile and target mass, particle charge, energy (or LET) and emission angle
- **absorbed dose and dose equivalent** per incident ion

Facilities

	Z_{proj} (max)	E_{proj} (MeV/u)	$E_{\text{proj}} (^{56}\text{Fe})$ (MeV/u)
NSRL (BNL)	79	40-3000	100-1100
AGS (BNL)	79	600-30000	600-10000
HIMAC (NIRS- Chiba)	36	100-800	400-500
LLUPTC (Loma Linda)	1	40-250	—

NSRL – 3-4 beam opportunities/year. Each running period provides several ion-energy combinations decided upon by a scientific program advisory committee in consultation with the user community and NASA program managers.

AGS – (7/05) 3, 5, 10 GeV/u C, Si, Fe

HIMAC – Operates ten months per year; shut down in March and September. Applications for beam time are due 6-8 months in advance.

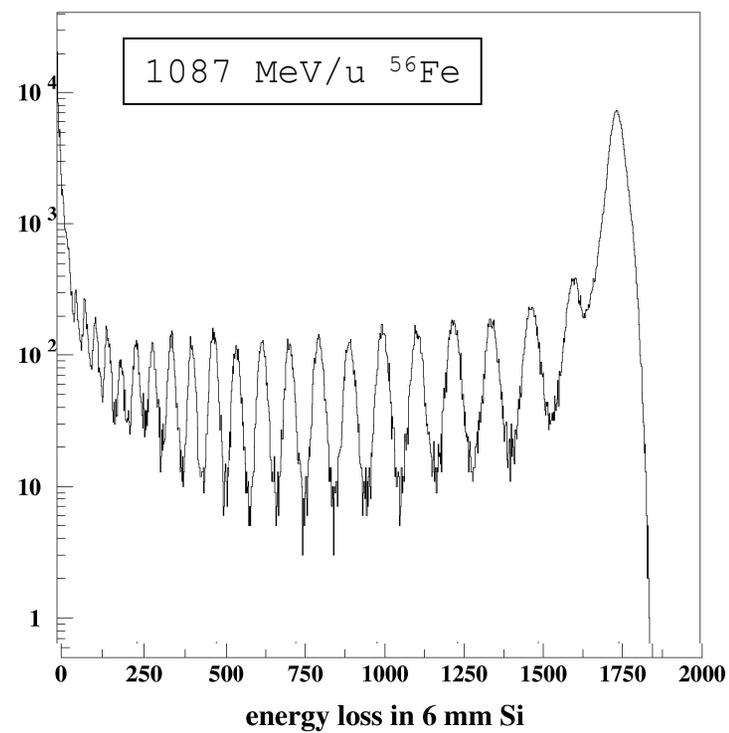
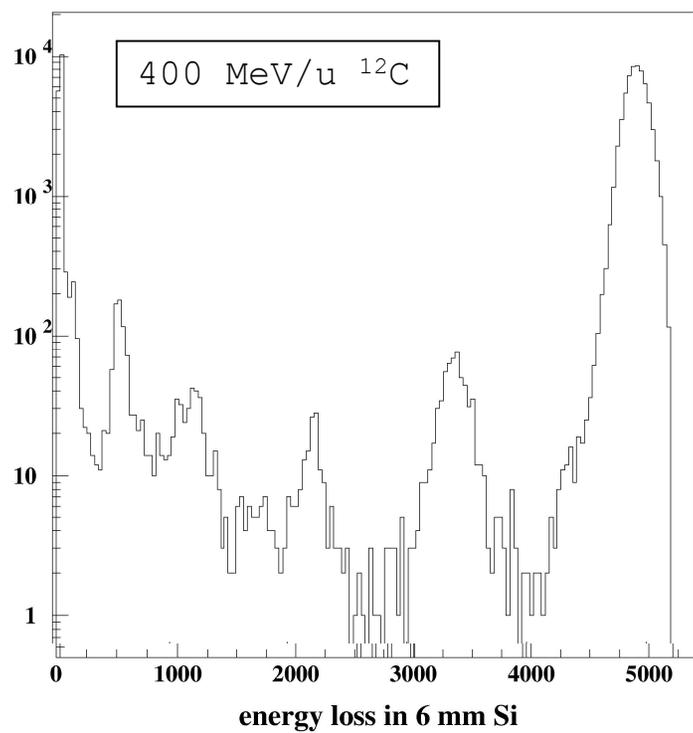
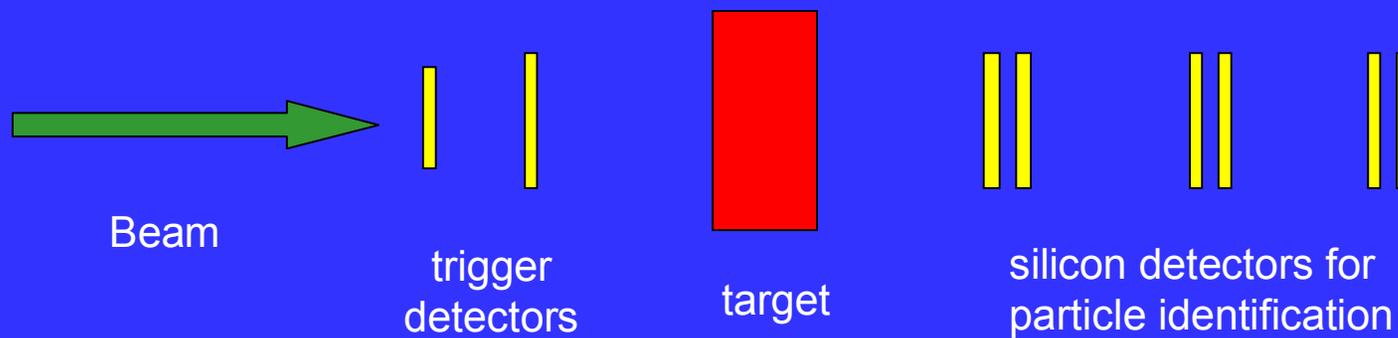
LLUPTC – Beam time potentially available throughout the year, usually during nights and weekends when patient treatments are not taking place.

Charged Particle Detectors

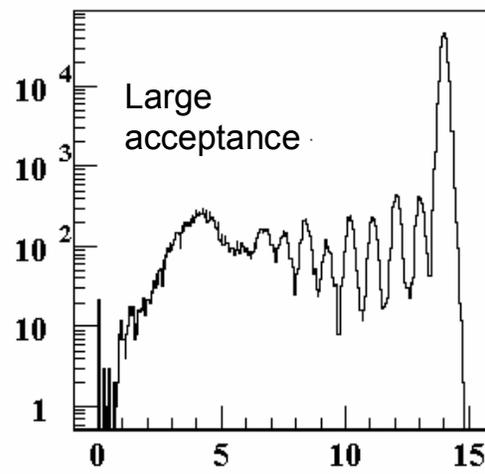
- **LBNL SSD stack**
 - small acceptance
 - measurements along beam axis
 - off-axis measurements possible
 - limited granularity
- **MSFC ZDDS**
 - larger acceptance (w/ hole at 0°)
- **higher granularity (multiple particle resolution)**
- **UH strip detectors**
 - largest acceptance
 - singly charged particles only

0° Charged Particle Cross Sections (6/05)

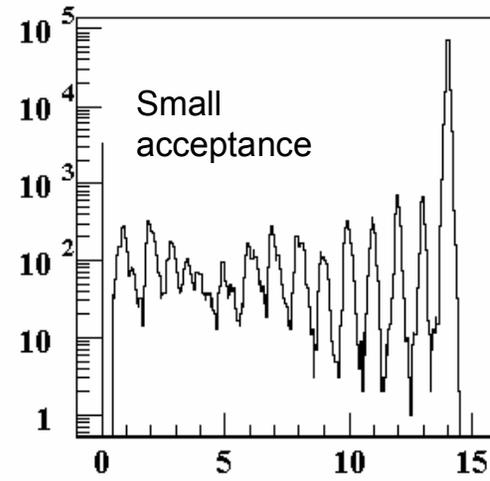
Ion	E (MeV/A)	E (MeV/A)	E (MeV/A)	E (MeV/A)	E (MeV/A)	Large acceptance	Small acceptance
⁵⁶ Fe	400	500	600	800	1000	1000 done	1000 done
⁴⁸ Ti	1000					Done	Done
⁴⁰ Ar	290	400	650			In progress	In progress
²⁸ Si	290	400	600	800	1200	All done	Some done
²⁰ Ne	290	400	600			All done	600 done
¹⁶ O	290	400	600	1000		Done	Done
¹⁴ N	290	400				Done	Done
¹² C	290	400				Done	Done
⁴ He	230						



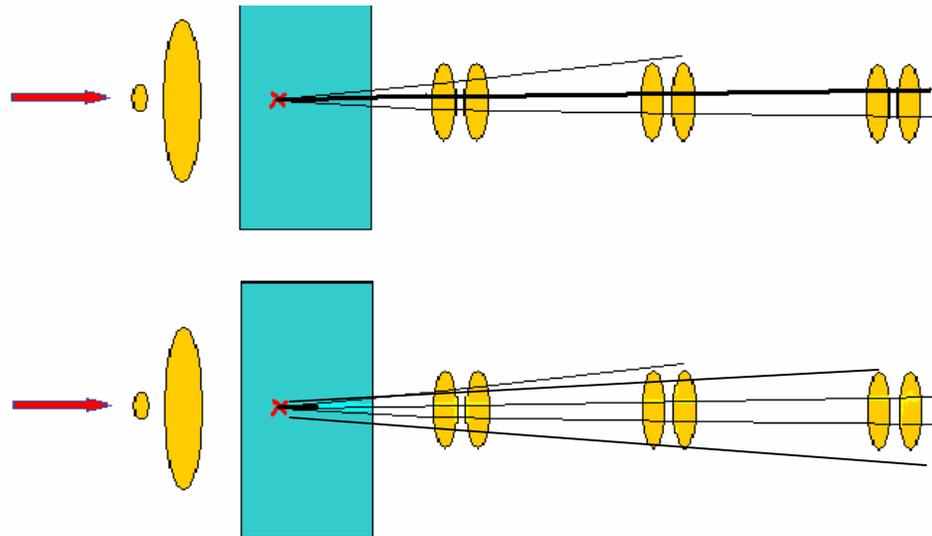
Si+C at 600 MeV/u



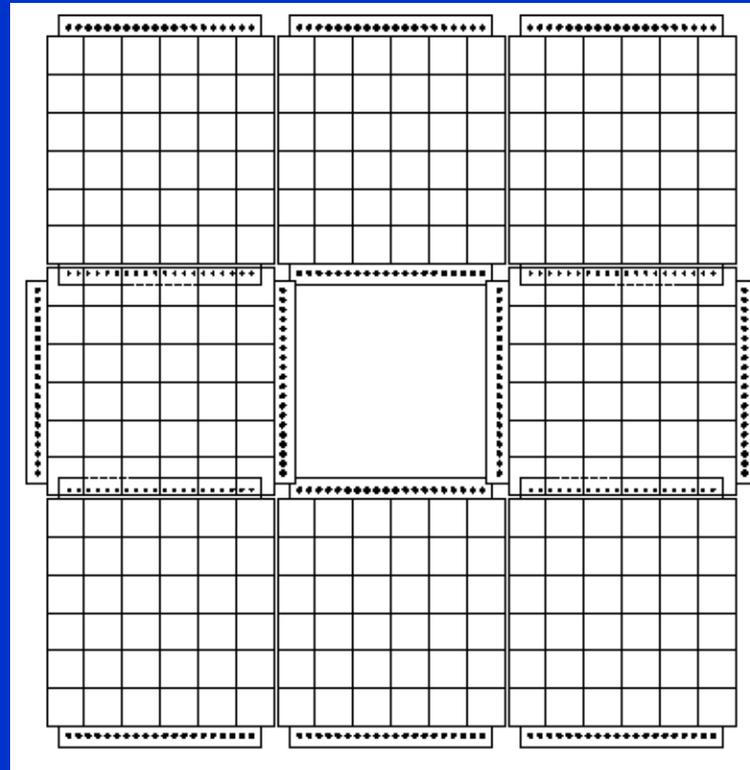
Detected Z



Detected Z

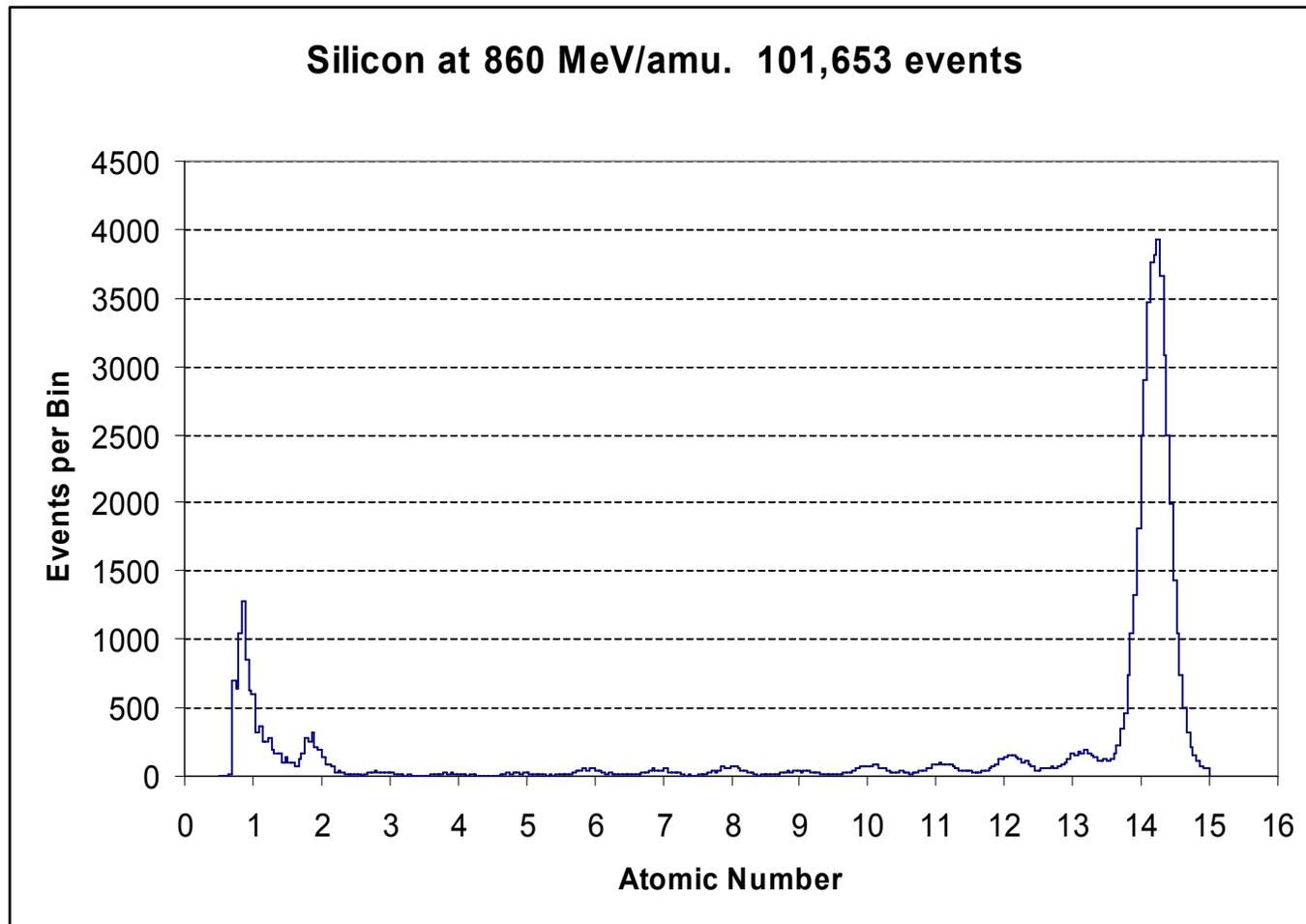


ZDDS (NASA-MSFC)

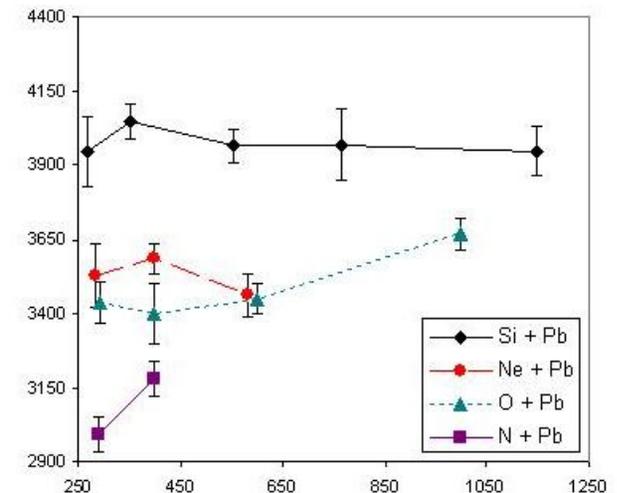
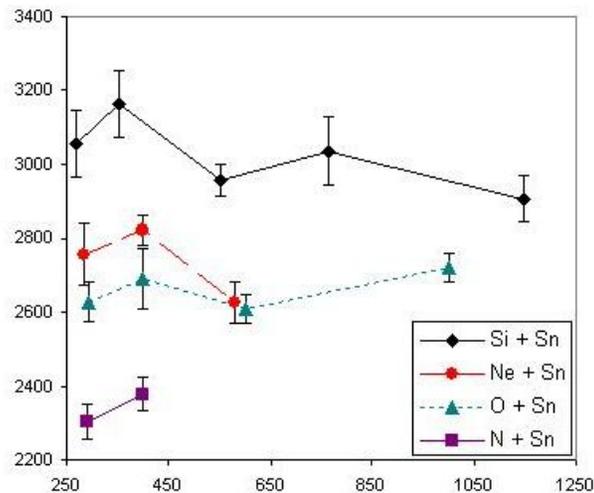
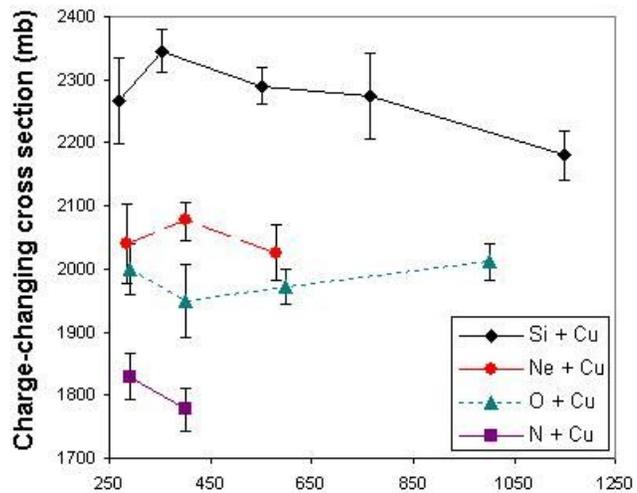
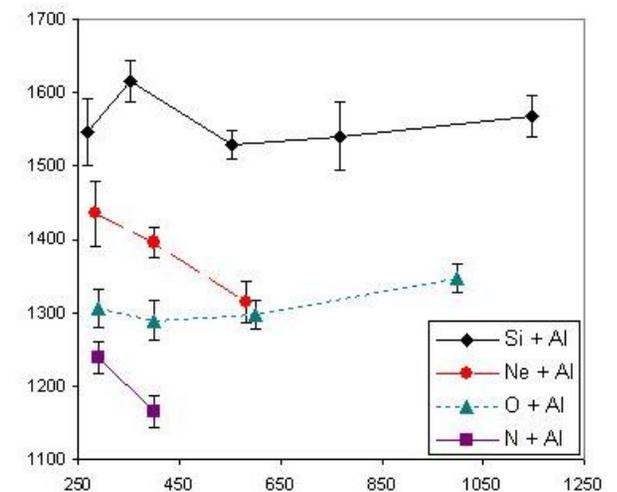
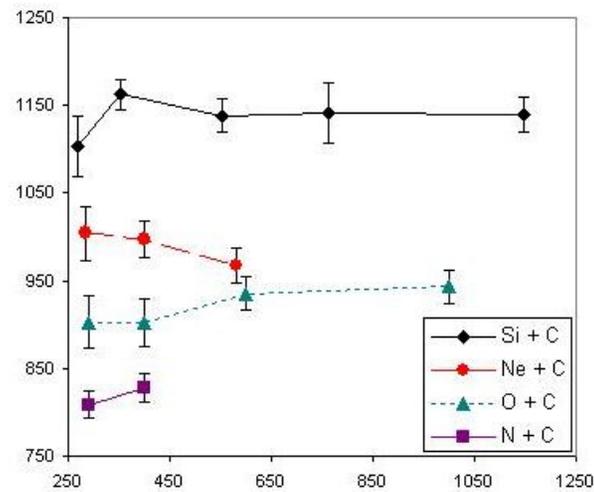
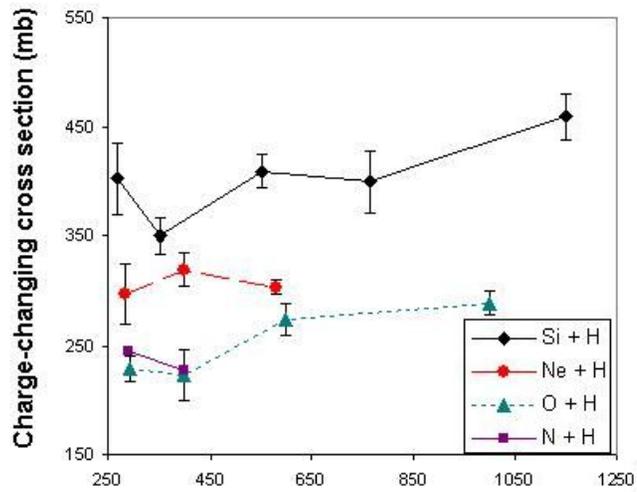


← 18 cm →

ZDDS Charge Resolution



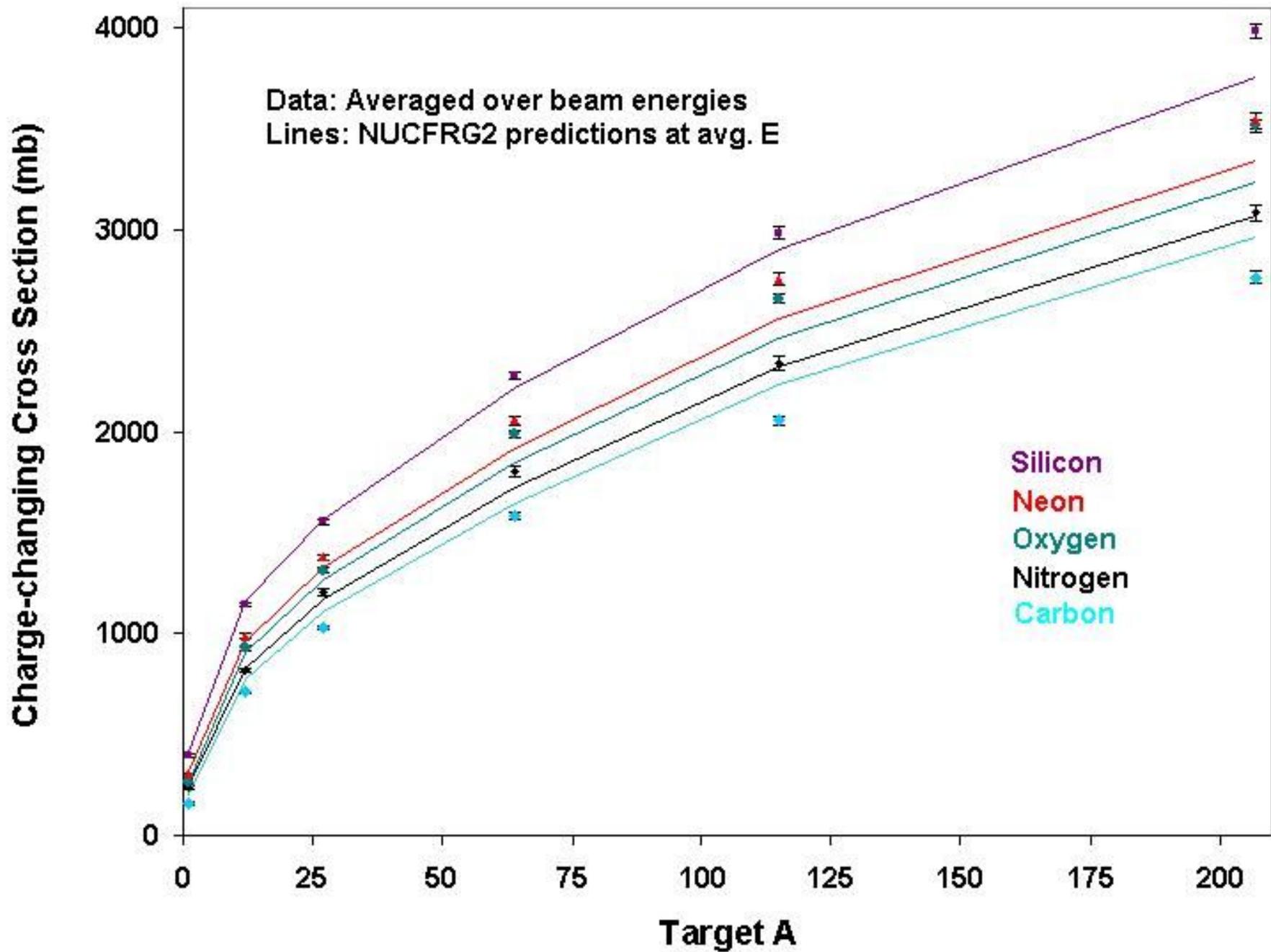
Charge-changing cross section systematics – beam energy dependence



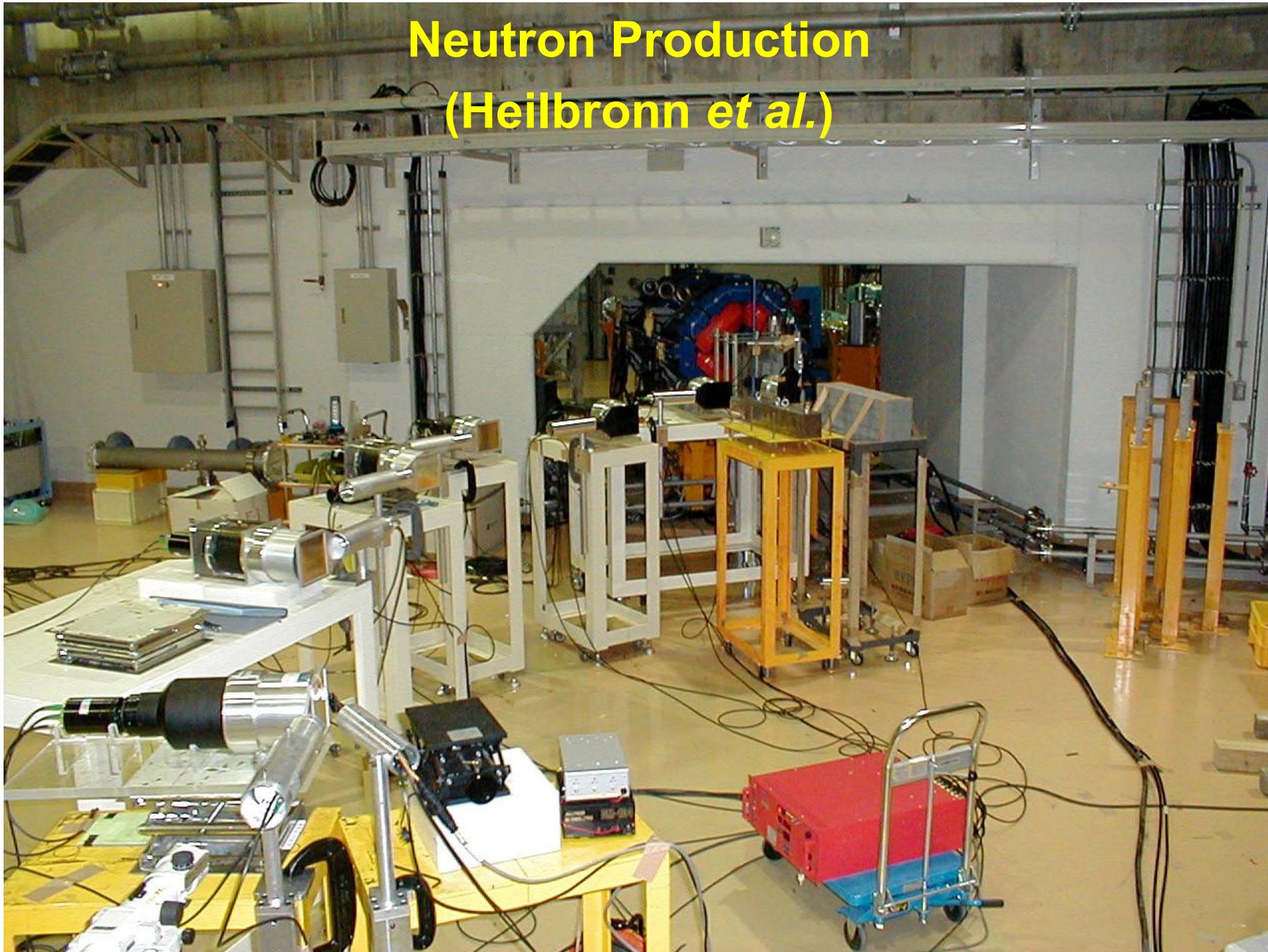
Beam Energy (MeV/amu)

Beam Energy (MeV/amu)

Beam Energy (MeV/amu)



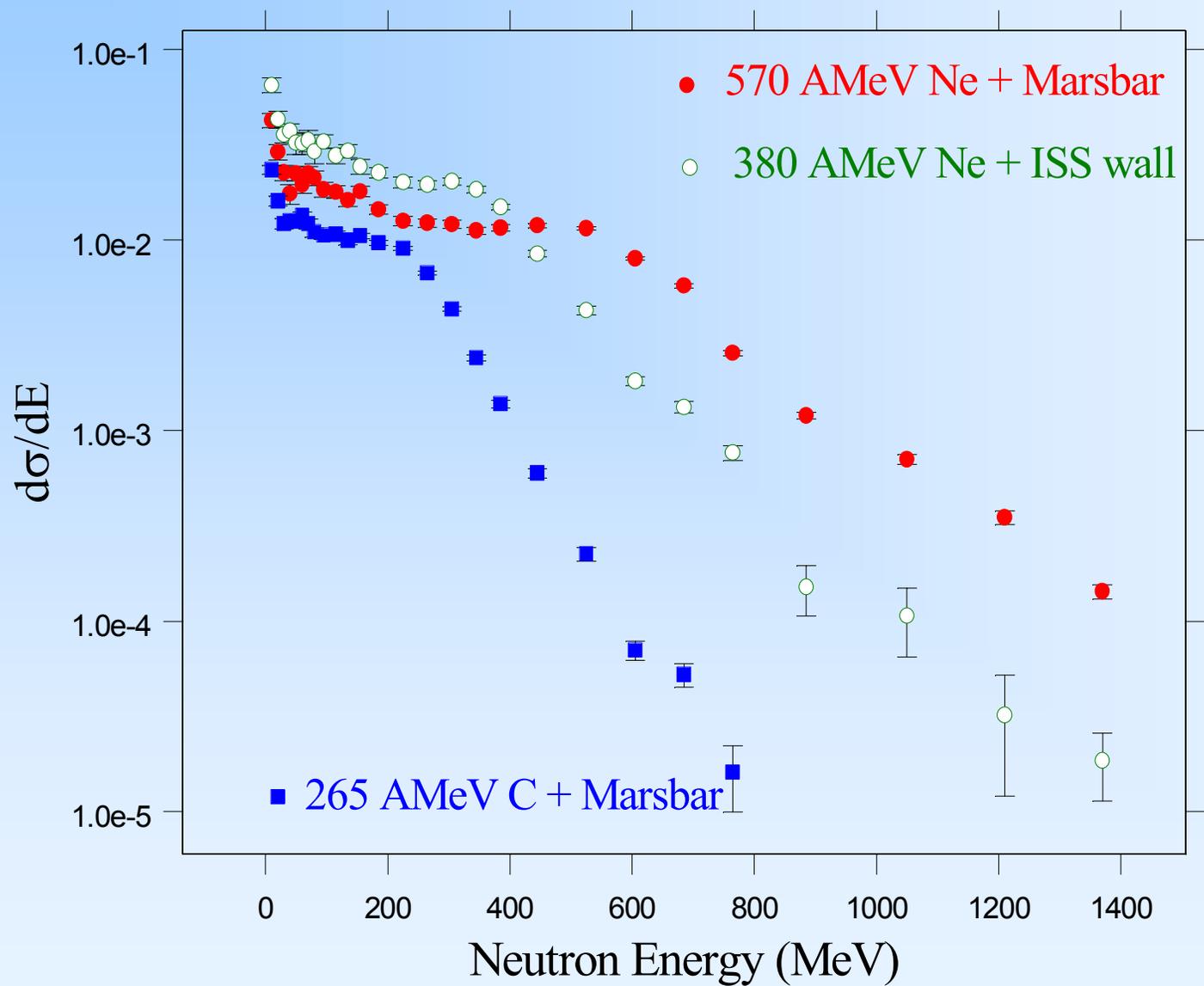
Neutron Production (Heilbronn *et al.*)



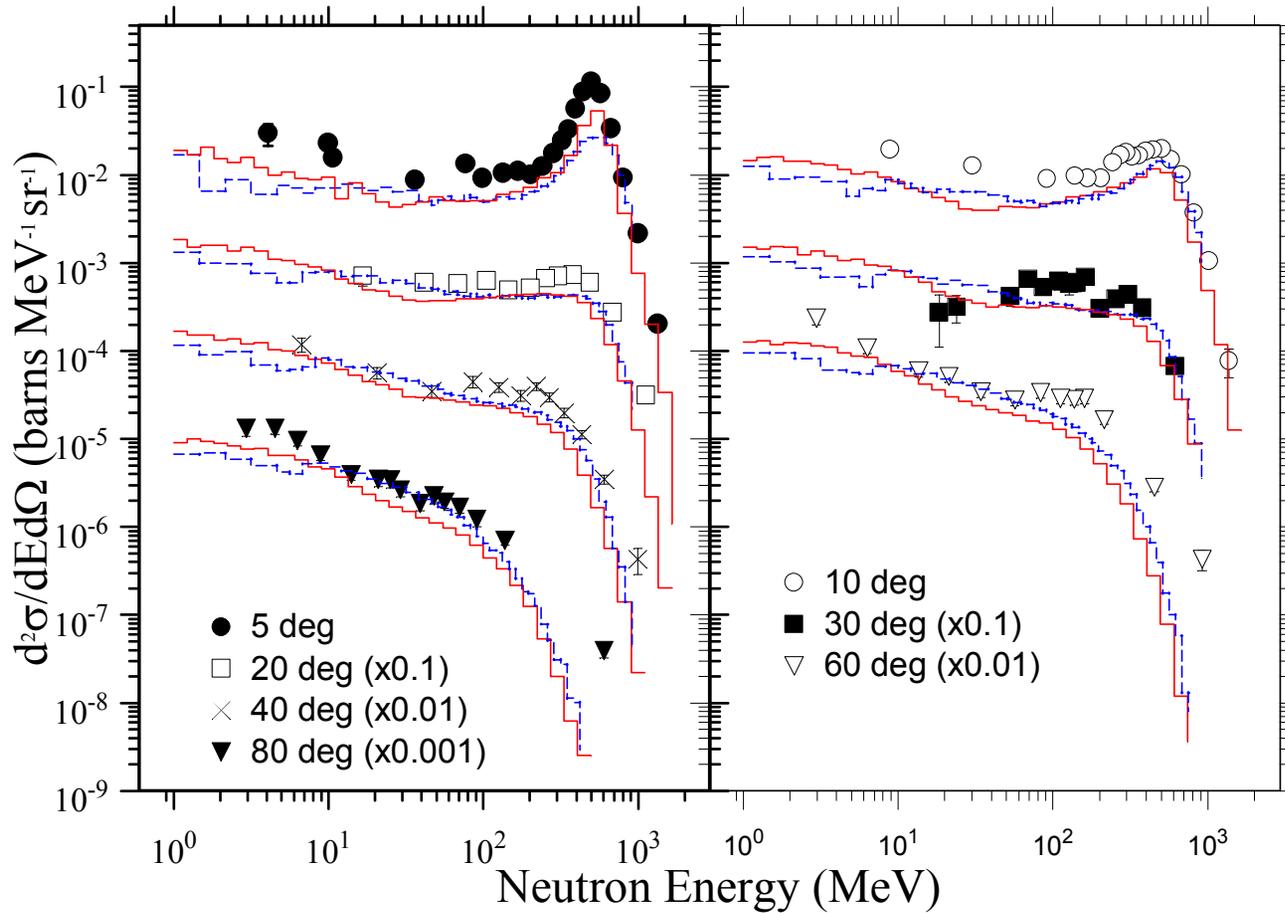
Neutron Yields and Production Cross Sections

- **a comprehensive set of 3-dimensional data available for the development of HZE transport models**
- **low energy (< 20 MeV) neutrons at back angles provide information on target fragmentation**
- **data will be available in a handbook on secondary particle production from heavy-ion interactions**

Energy Distributions



600 MeV/nucleon Ne + "Mars bar"



- PHITS and SHIELD-HIT both reproduce the spectra shapes very well

- Both models slightly underestimate the yield

Neutron Yields and Cross Sections

Ion	E (MeV/A)	E (MeV/A)	E (MeV/A)	Targets	Type	Pub.
^{131}Xe	400			Li, C, CH ₂ , Al, Cu, Pb	σ	draft
^{93}Nb	272	435		Nb, Al (272 only)	y	Y
^{84}Kr	400			Li, C, Al, Cu, Pb	σ	draft
^{56}Fe	400	500		Li,C,Cu,Pb Li,CH ₂ ,Al	σ,y	draft
^{40}Ar	400	560		C, Cu, Pb	σ	Y
^{28}Si	600			C, Cu, Pb	σ	draft
^{20}Ne	400	600		C, Cu, Pb, ISS, Mars Reg.	σ	Y
^{14}N	400			C, Cu	σ	draft
^{12}C	155	290	400	Al (155), C, Cu, Pb Mars Reg (290)	y (Al), σ	Y
^4He	155	230		Al, Cu (290)	σ,y	Y

“Handbook on Secondary Particle Production and Transport by High Energy Heavy Ions”

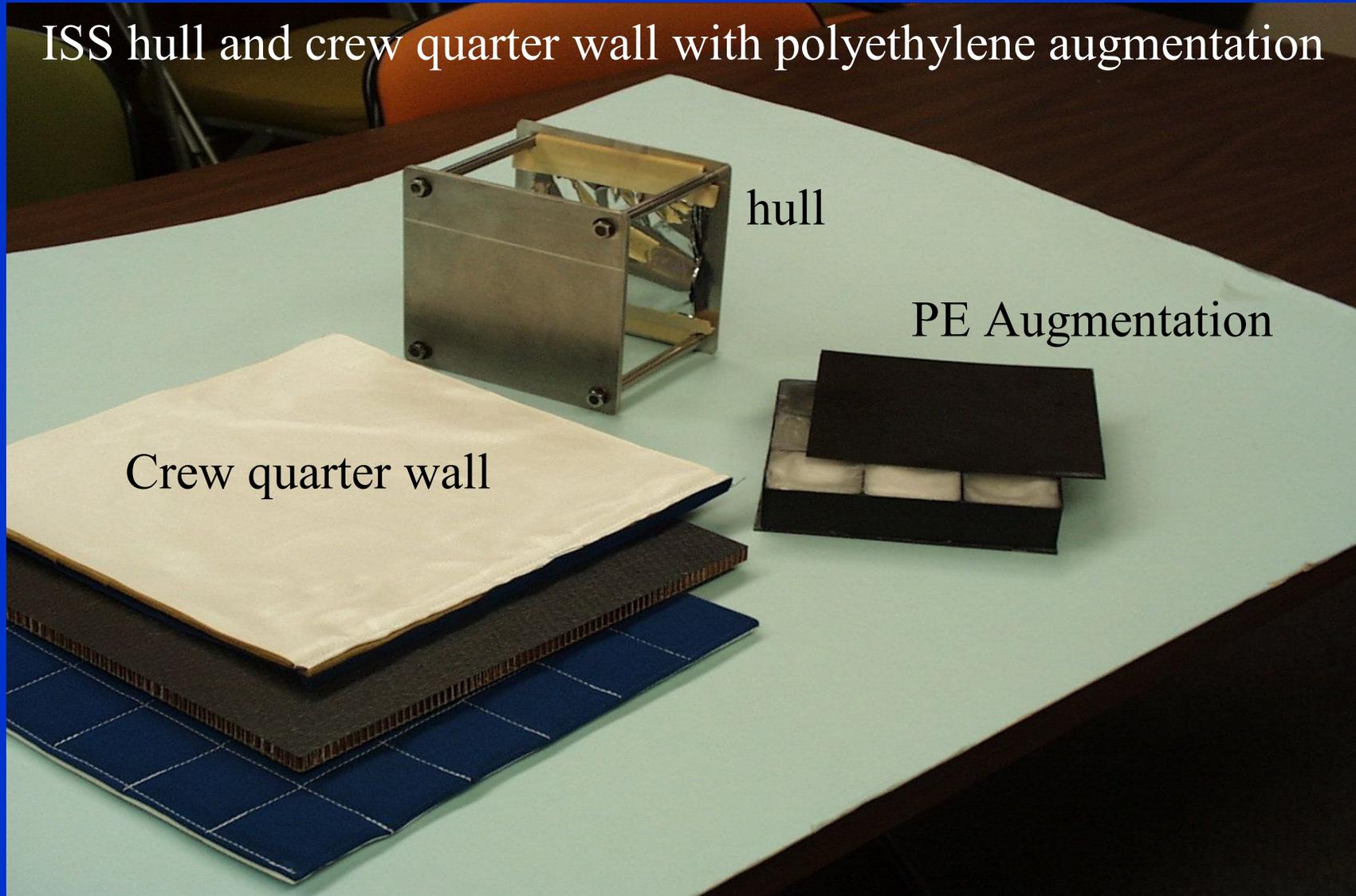
T. Nakamura and L. Heilbronn, Eds.

(World Scientific, Singapore, to be published late 2005)

- neutron production cross section data
- thick-target neutron yields
- neutron production behind shielding
- spallation production cross sections
- moving-source parameterizations of the neutron-production cross sections and thick-target yields
- detailed descriptions of the experiments and analysis

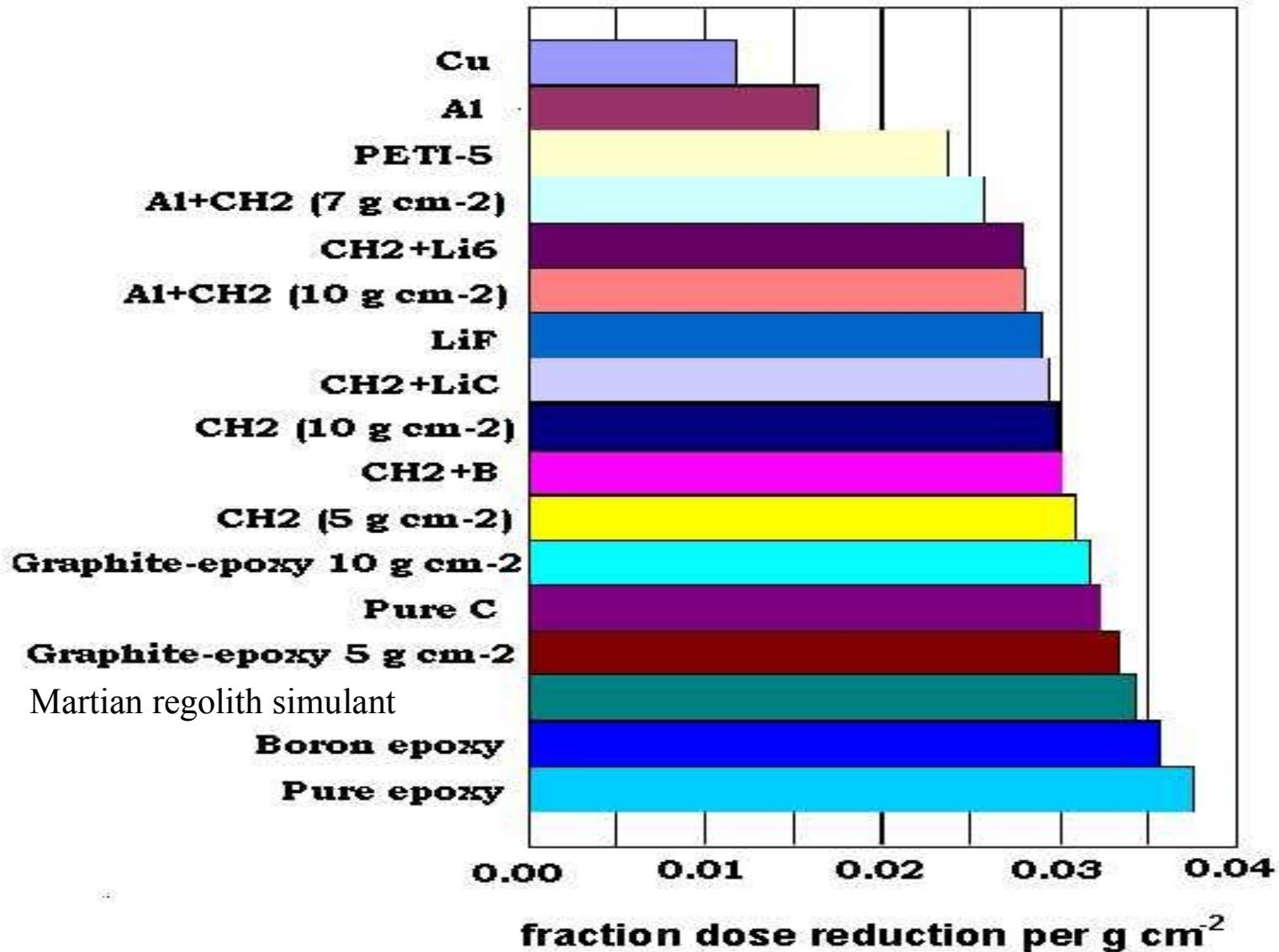
Radiation Transport in Materials

ISS hull and crew quarter wall with polyethylene augmentation



Dose reduction (1 GeV/u ^{56}Fe)

(Zeitlin, *et al.*)



Testing and Data Analysis of Composite Radiation Shielding Materials

To date, 127 samples tested and data analysed:

NASA-MSFC	C (290)	2
	N (290)	6
	Ne (290)	6
	Si (290)	6
	Si (800)	6
	Ti (1000)	2
	Fe (800)	2
	Fe (1000)	6
NASA-LaRC	O (600)	15
	O (1000)	2
	Fe (1000)	17
ORNL	O (600)	3
	O (1000)	15
U. Of Utah	O (600)	9
	H (40)	30

NASA Measurements Consortium: LBNL Cross-sections.



Credits

SRHP

NSRL

ENTER USER NAME :

ENTER PASSWORD :

Log In

(And by the way, your I.P. address has been logged. °_°)

NASA Measurements Consortium: LBNL Cross-sections.

Main Page

SRHP

NSRL

Credits

Green cells are active links.

Ion	Energy (MeV/nucleon)					
⁵⁶ Fe	400	500	600	800	1000	5000
⁴⁸ Ti					1000	
⁴⁰ Ar	290	400	650			
²⁸ Si	290	400	600	800	1200	
²⁰ Ne	290	400	600			
¹⁶ O	290	400	600		1000	
¹⁴ N	290	400				
¹² C	290	400				
⁴ H	230					

NASA Measurements Consortium : LBNL Cross-sections

Click [here](#) for the entire silicon data set in .pdf format

Main Page

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Credits

Silicon at 290 MeV/nucleon

(Cross sections listed by target. All values are in mBarn)

	H	C	Al	Cu	Sn	Pb
σ_{cc}	402 ± 33	1103 ± 34	1546 ± 46	2276 ± 68	3056 ± 91	3943 ± 120
Z = 13	90 ± 9	144 ± 8	175 ± 10	219 ± 14	259 ± 18	380 ± 28
Z = 12	105 ± 11	161 ± 9	190 ± 11	226 ± 13	268 ± 17	338 ± 24
Z = 11	53 ± 6	89 ± 5	111 ± 7	141 ± 9	166 ± 11	200 ± 15
Z = 10	49 ± 6	91 ± 5	106 ± 7	143 ± 9	151 ± 11	199 ± 15
Z = 9	18 ± 3	49 ± 3	67 ± 5	69 ± 5	90 ± 7	97 ± 10
Z = 8	31 ± 5	91 ± 5	112 ± 7	157 ± 9	185 ± 12	185 ± 15

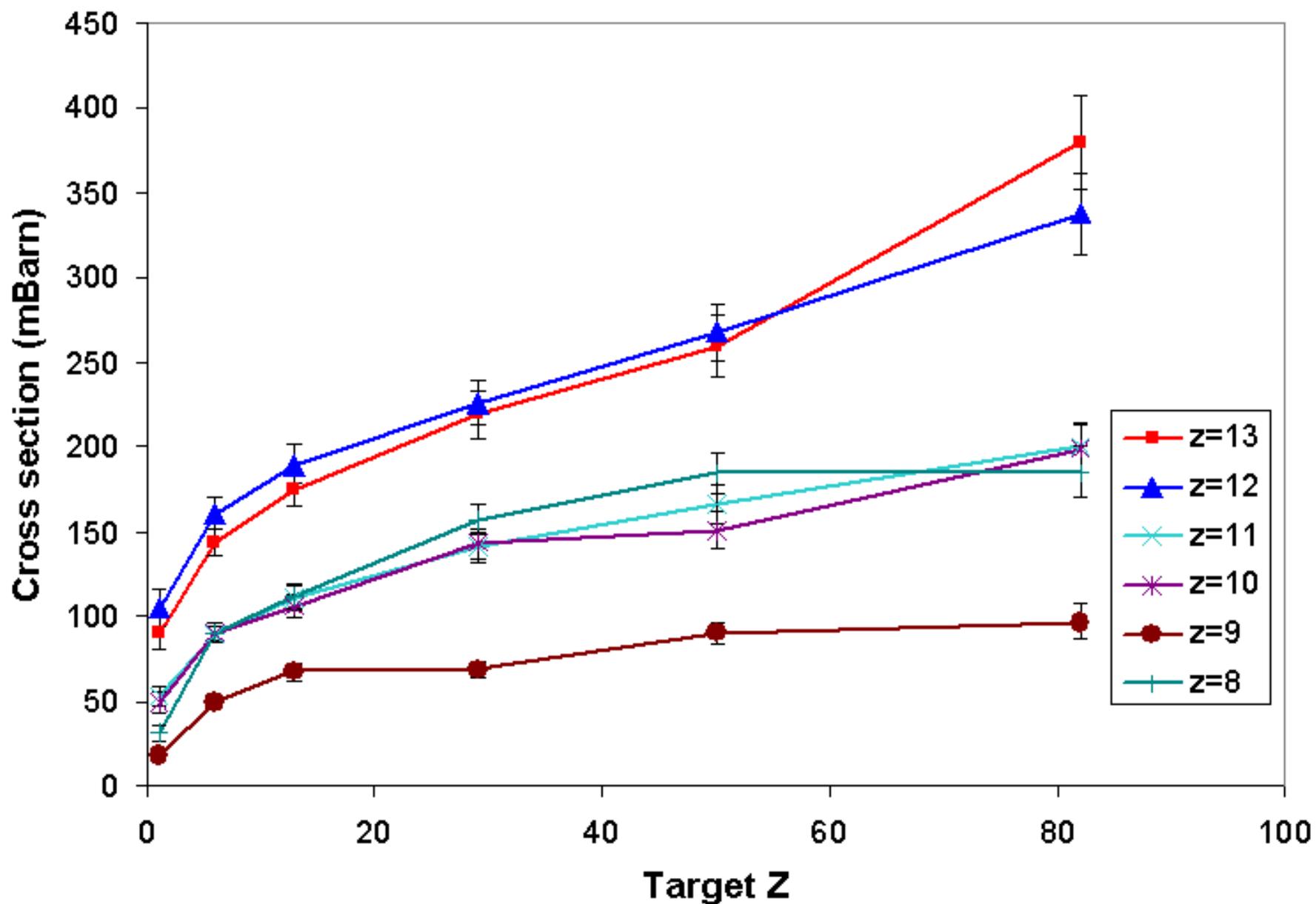
Charge-Changing Cross Sections of Si Beams

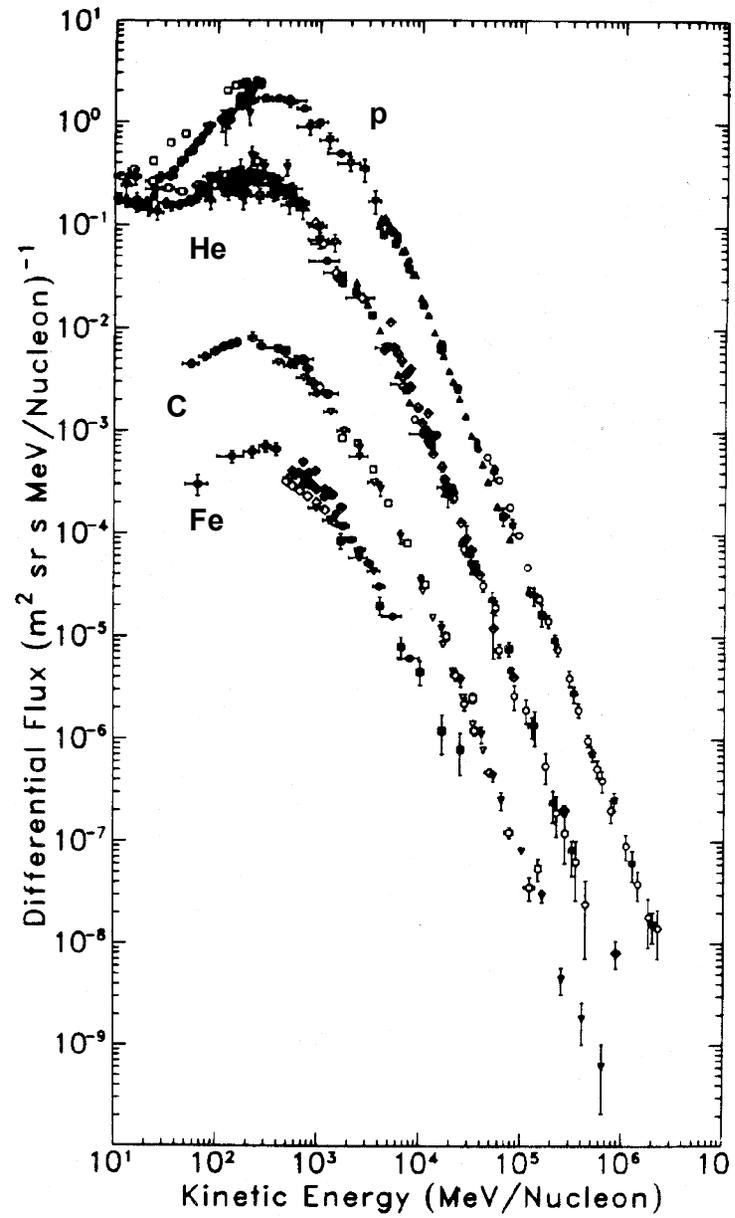
(E_{beam} , MeV/amu)	$\sigma_{\text{cc}}(\text{Si}+\text{H})$	$\sigma_{\text{cc}}(\text{Si}+\text{C})$	$\sigma_{\text{cc}}(\text{Si}+\text{Al})$	$\sigma_{\text{cc}}(\text{Si}+\text{Cu})$	$\sigma_{\text{cc}}(\text{Si}+\text{Sn})$	$\sigma_{\text{cc}}(\text{Si}+\text{Pb})$
270 ± 8	402 ± 33	1103 ± 34	1546 ± 46	2276 ± 68	3056 ± 91	3943 ± 120
352 ± 12	350 ± 16	1162 ± 17	1615 ± 29	2345 ± 34	3163 ± 91	4047 ± 59
552 ± 16	409 ± 15	1138 ± 19	1528 ± 19	2289 ± 29	2955 ± 44	3962 ± 57
765 ± 5	399 ± 28	1141 ± 34	1540 ± 46	2273 ± 68	3035 ± 91	3967 ± 120
1147 ± 11	459 ± 21	1139 ± 20	1567 ± 28	2179 ± 39	2905 ± 62	3945 ± 84

Fragment Production Cross Sections at Large Acceptance *Si + H*

E_{beam}	263	350	560	765	1160
$\sigma(Z=13)$	90 ± 9	76 ± 4	75 ± 3	74 ± 7	61 ± 4
$\sigma(Z=12)$	105 ± 11	93 ± 5	88 ± 3	83 ± 8	74 ± 5
$\sigma(Z=11)$	53 ± 6	48 ± 3	46 ± 2	43 ± 4	43 ± 3
$\sigma(Z=10)$	49 ± 6	43 ± 3	49 ± 2	45 ± 5	44 ± 3
$\sigma(Z=9)$	18 ± 3	16 ± 3	21 ± 1	22 ± 2	23 ± 3
$\sigma(Z=8)$	31 ± 5	31 ± 7	41 ± 2	41 ± 5	44 ± 5
$\sigma(Z=7)$		15 ± 5	26 ± 2	22 ± 3	36 ± 5
$\sigma(Z=6)$		17 ± 6	21 ± 2	30 ± 4	28 ± 5
$\sigma(\leq 5)$		26 ± 14	46 ± 7	51 ± 19	92 ± 20

Fragment Cross Sections for Si-290 on Elemental Targets





High Energy Fragmentation Measurements at the AGS (7/05)

C, Si, Fe at 3, 5 and 10 GeV/A.

- 0-3 deg. - plastic scintillators and silicon SSD (LBNL)
- 3-9 deg. - segmented silicon detectors (NASA-MSFC)
- 3-45 deg. - silicon strip detectors to cover single charged particles (UH)
- 15-115 deg – neutron counters (LBNL)
- FLUKA Simulations of the entire setup including all detector systems. (UH)

AGS Detector Layout - 7/05

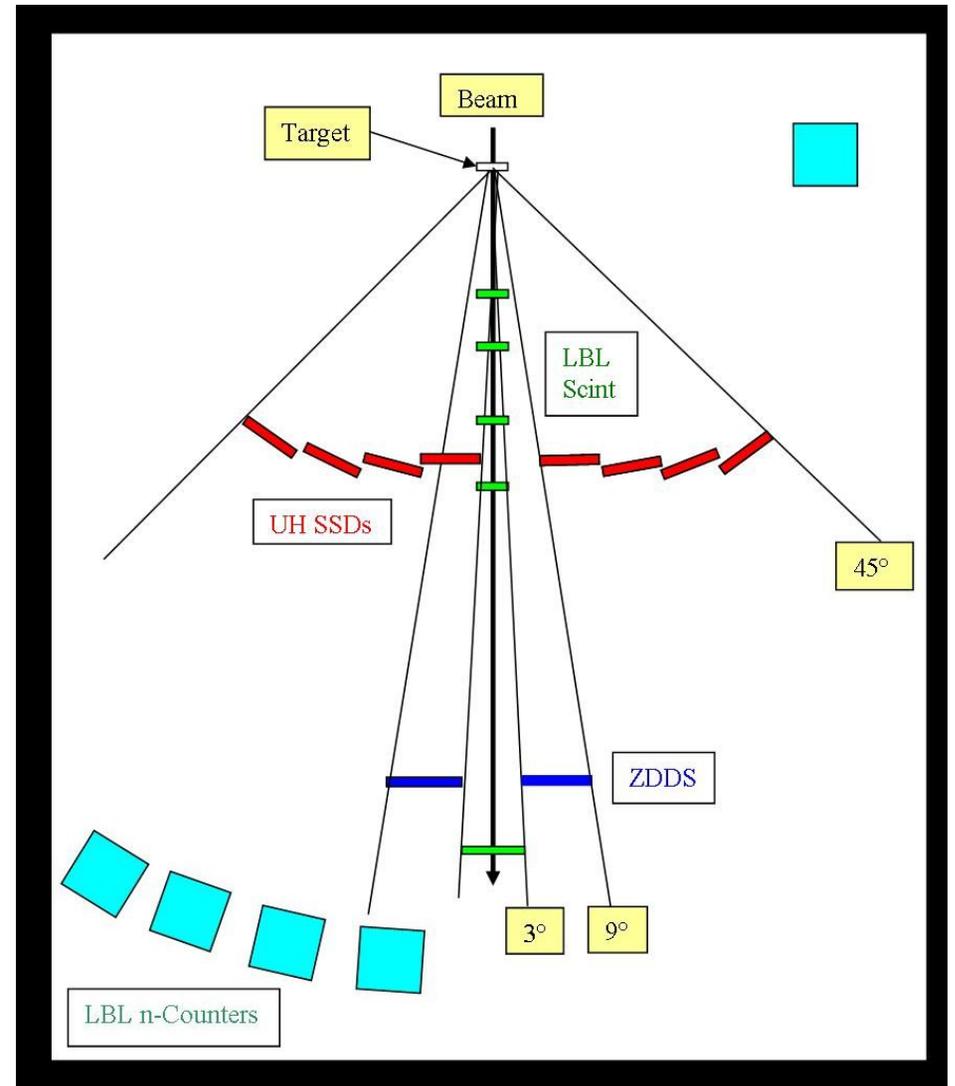
UH-MSFC-LBNL

HIMAC

Test Run

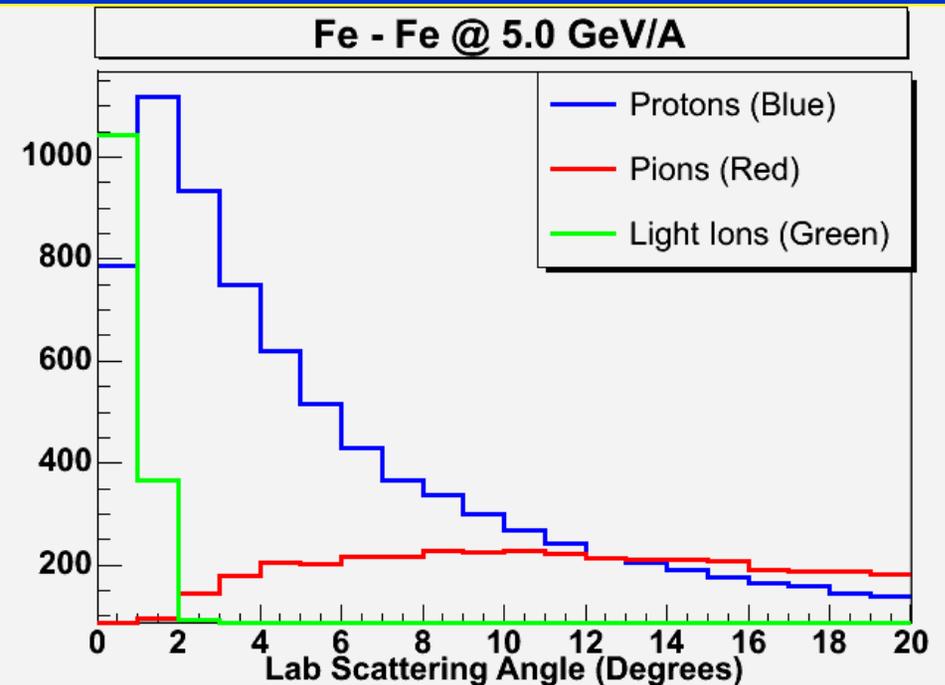
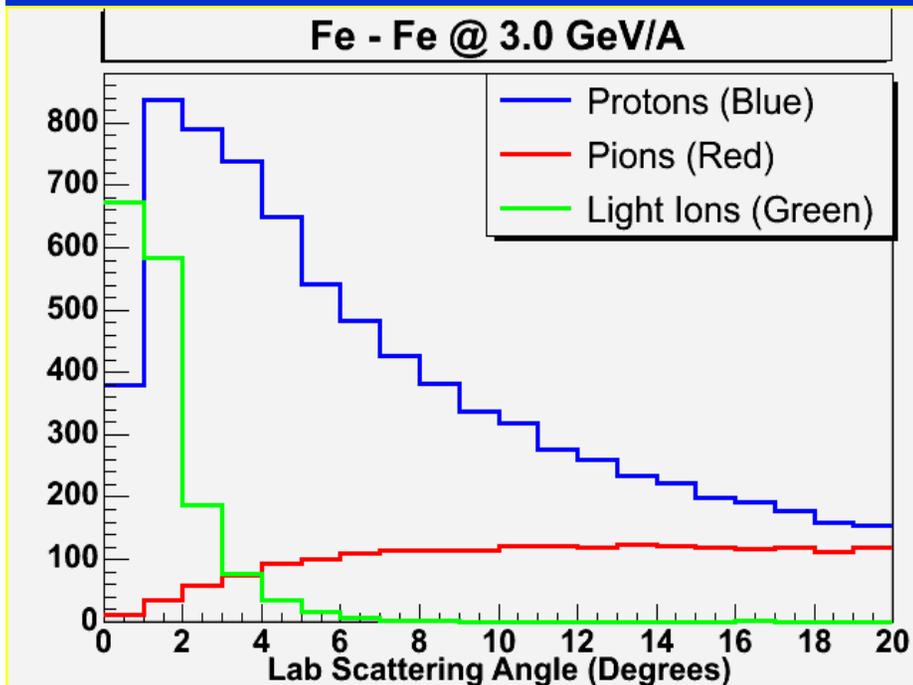
Setup

Feb. '05



Planned AGS Layout

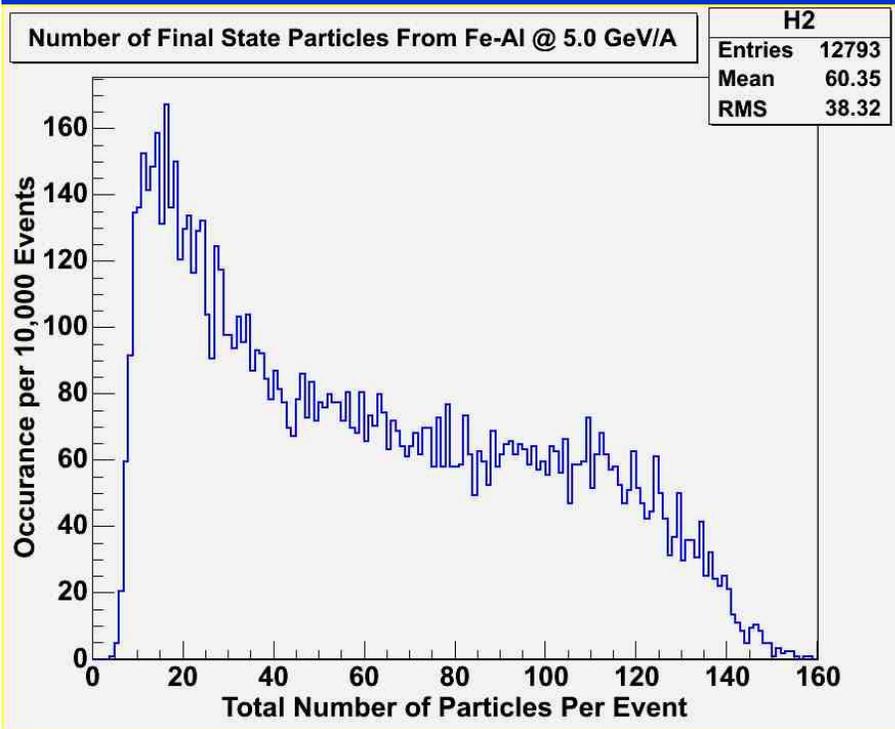
FLUKA RQMD AGS Predictions (UH)



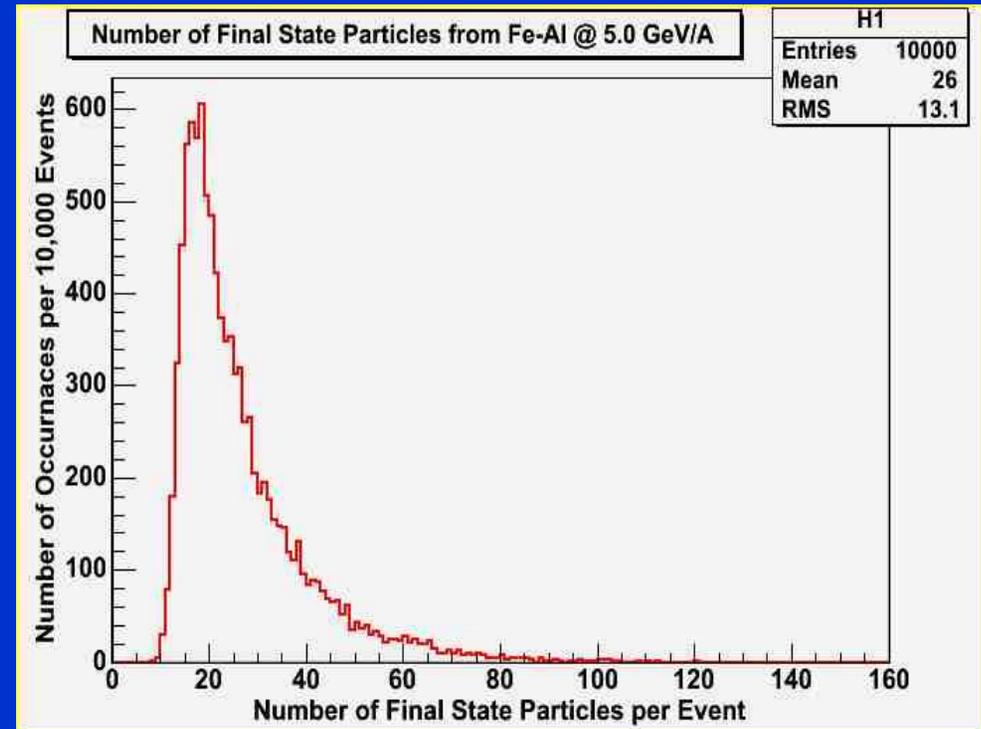
- RQMD Generator Predictions of Scattering Angle Distributions at 3 & 5 GeV/A Fe-Fe Collisions

One of the Primary Motivations for the AGS Measurements...

Predicted Number of Particles Per Event RQMD v DPMJET for Fe-Al @ 5 GeV/A



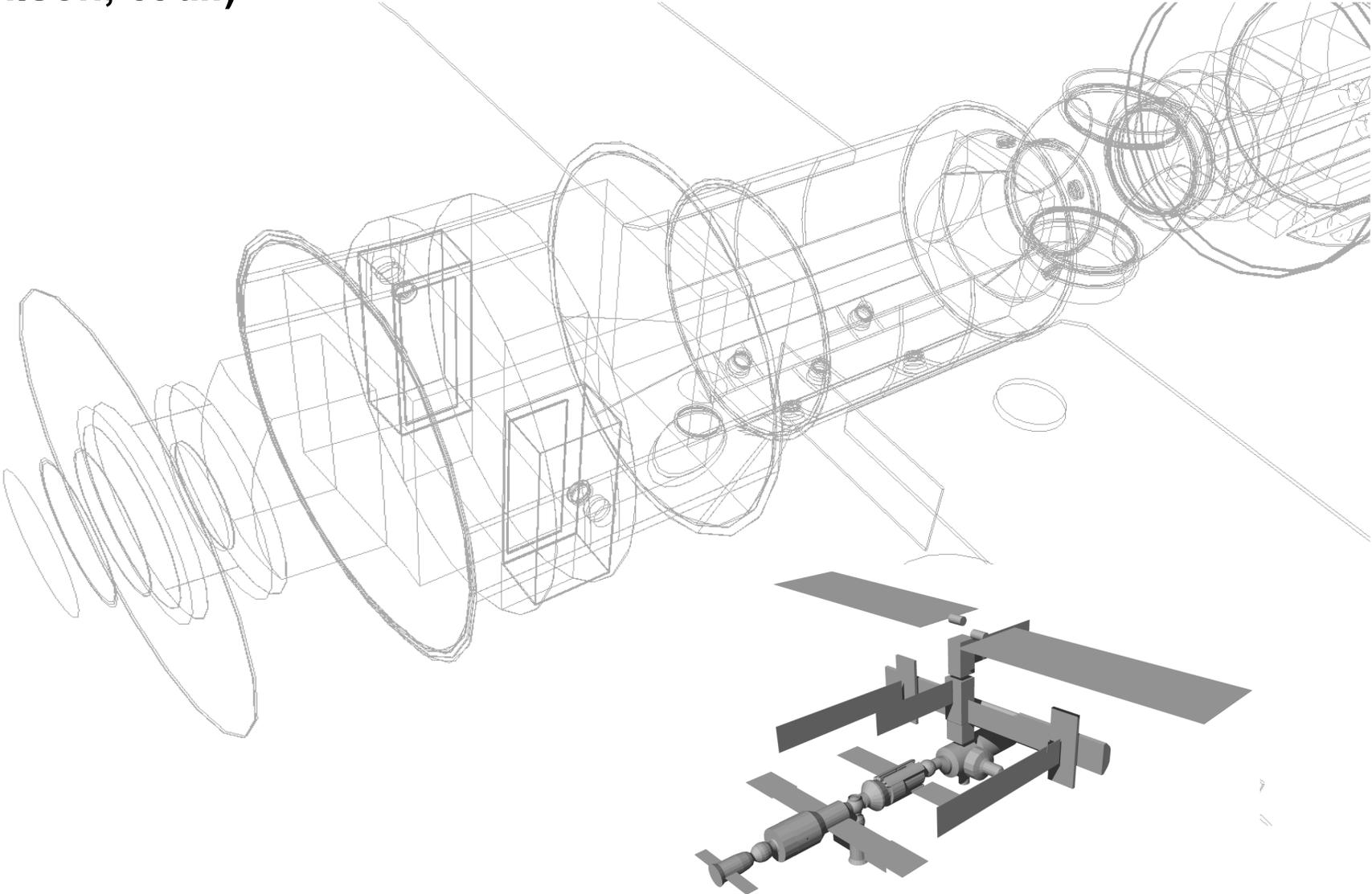
RQMD



DPMJET

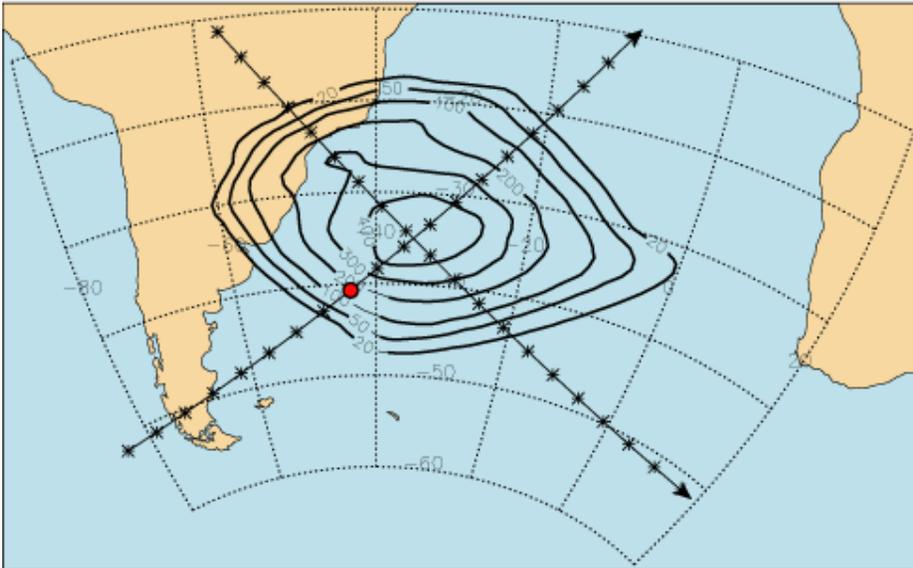
Radiation Shielding Model: Service Module Crew Quarters

(Wilson, *et al.*)

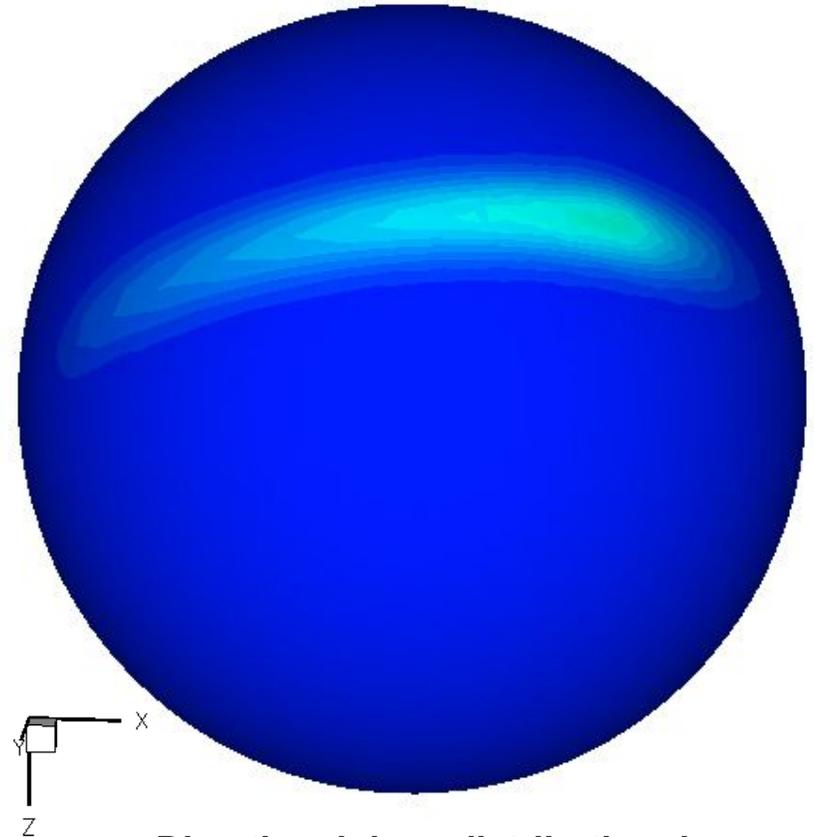


Radiation Environment: South Atlantic Anomaly, Protons

400-km Solar Min. Flux (greater than 100 MeV), protons/(cm²-sec)



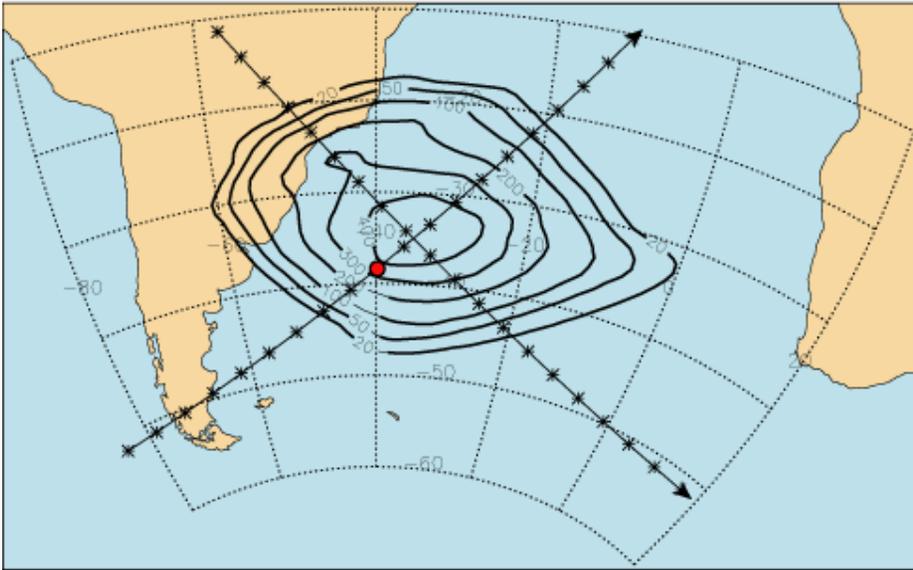
ISS ascending pass through center of SAA



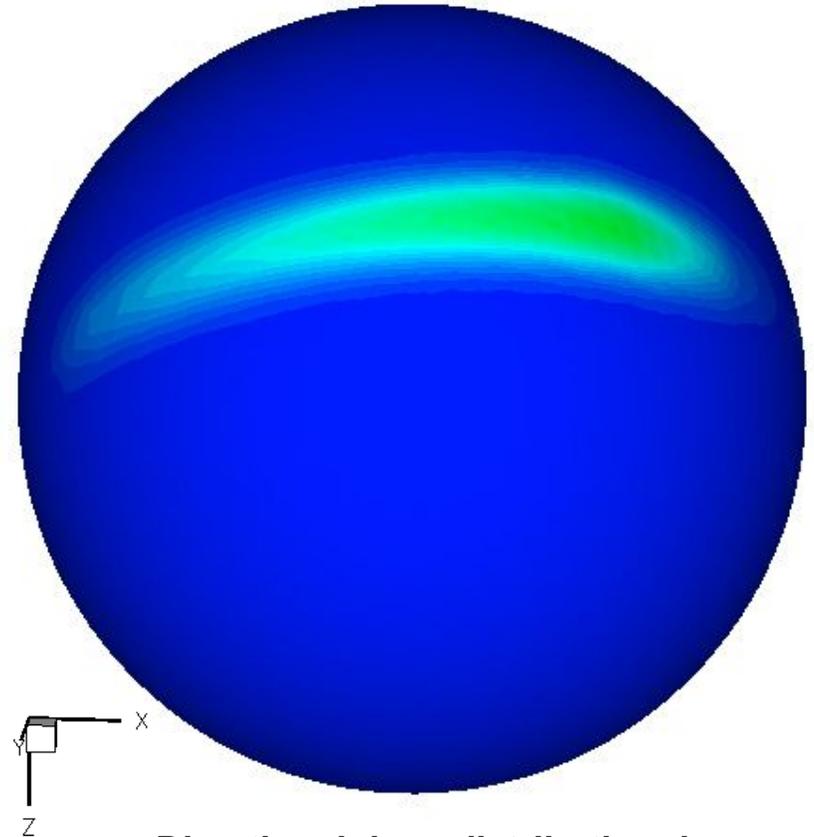
Directional dose distribution due to protons within the SM starboard Crew Quarters

Radiation Environment: South Atlantic Anomaly, Protons

400-km Solar Min. Flux (greater than 100 MeV), protons/(cm²-sec)



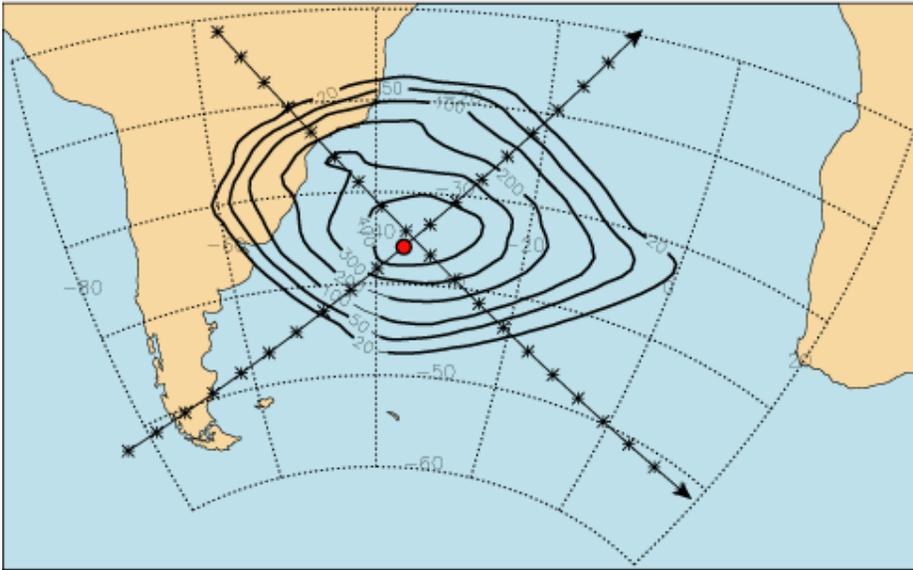
ISS ascending pass through center of SAA



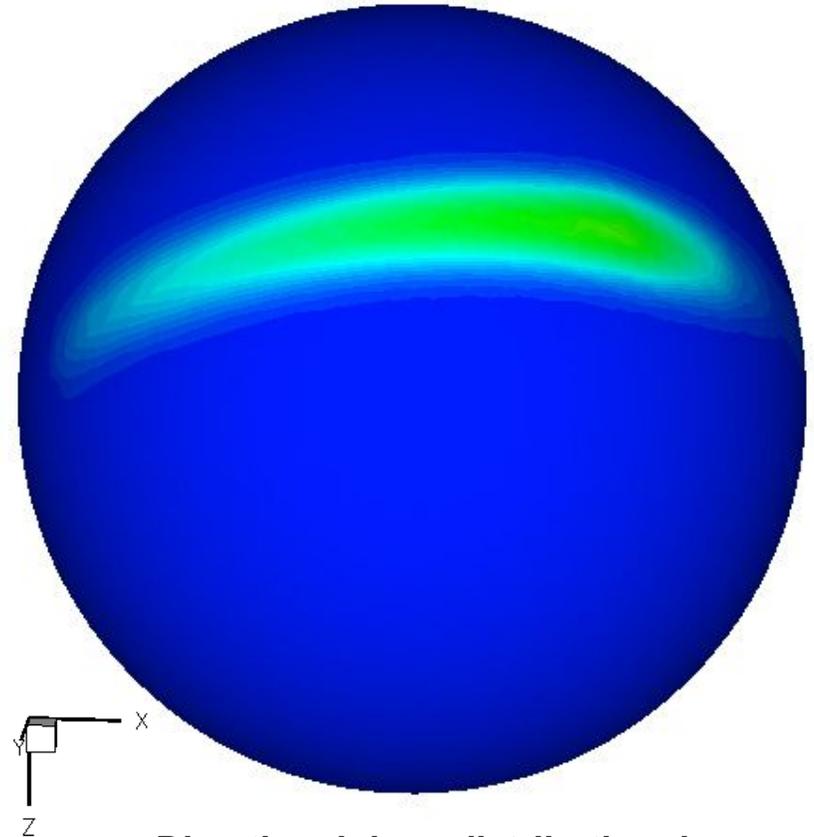
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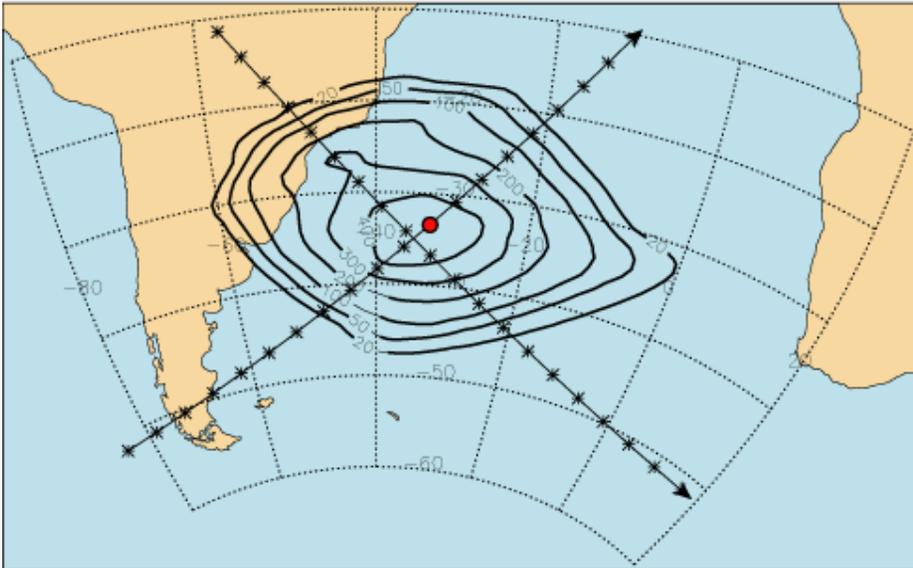
ISS ascending pass through center of SAA



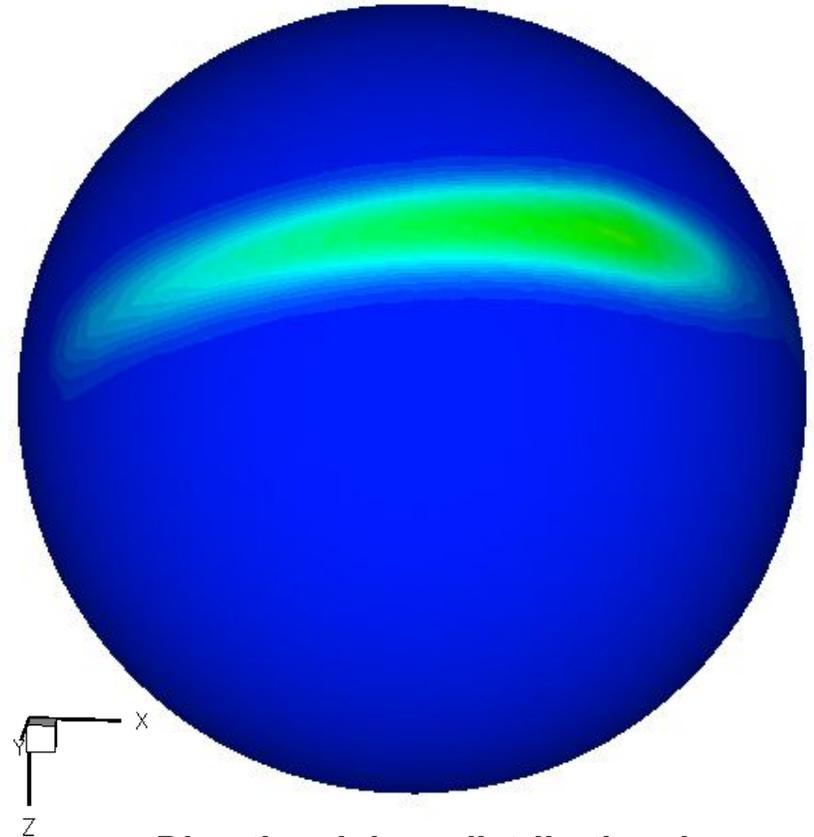
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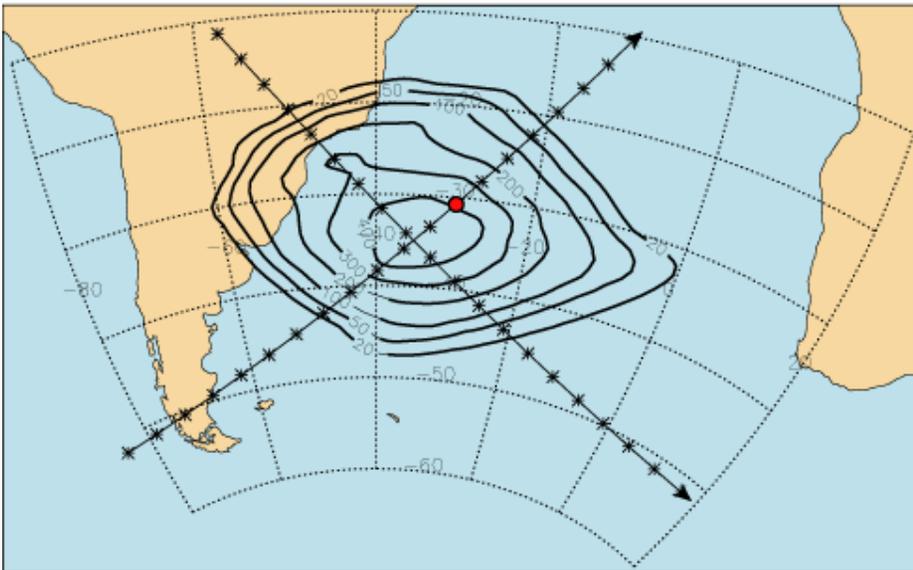
ISS ascending pass through center of SAA



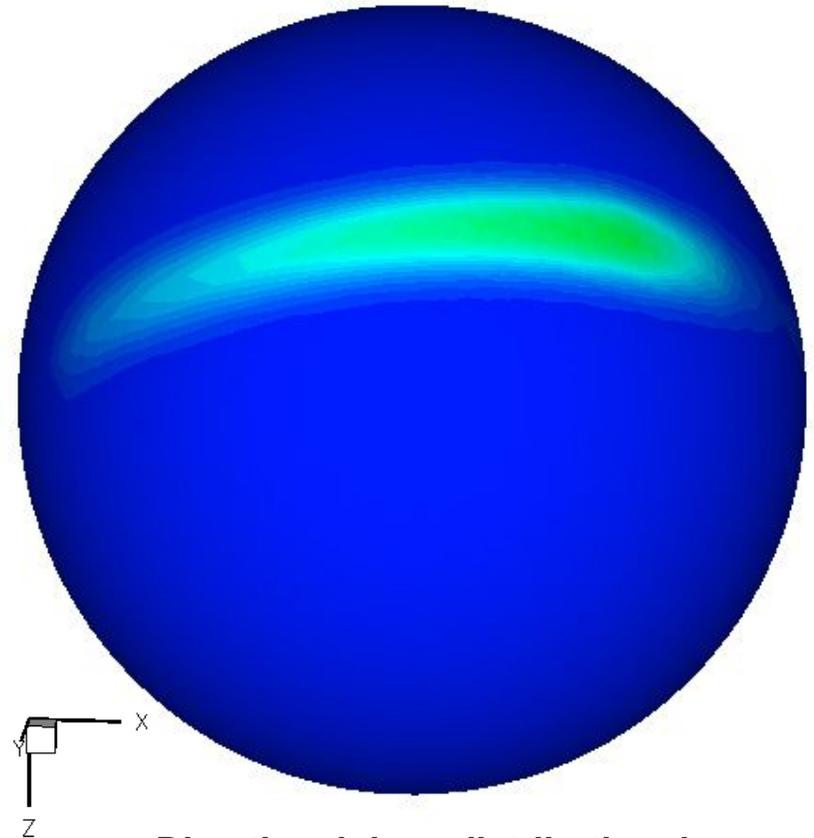
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400-km Solar Min. Flux (greater than 100 MeV), protons/(cm²-sec)



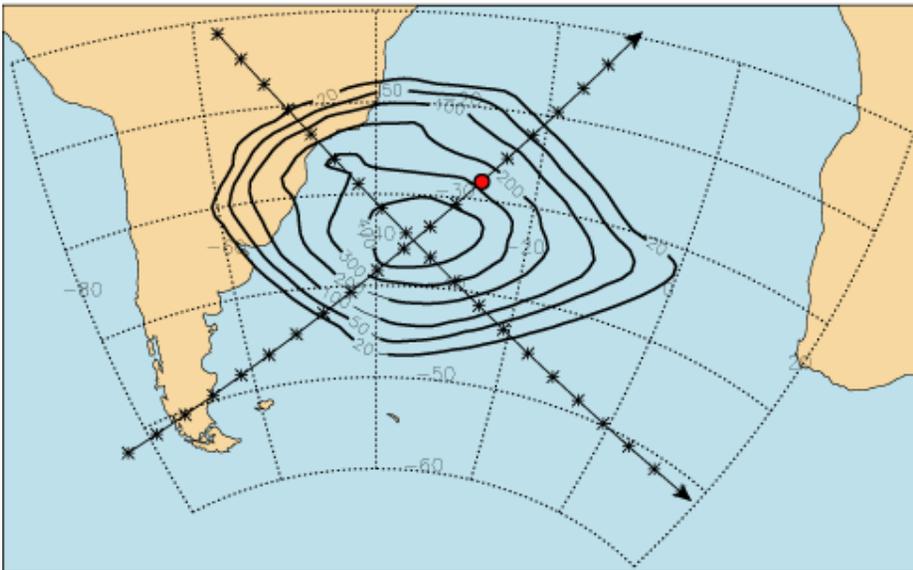
ISS ascending pass through center of SAA



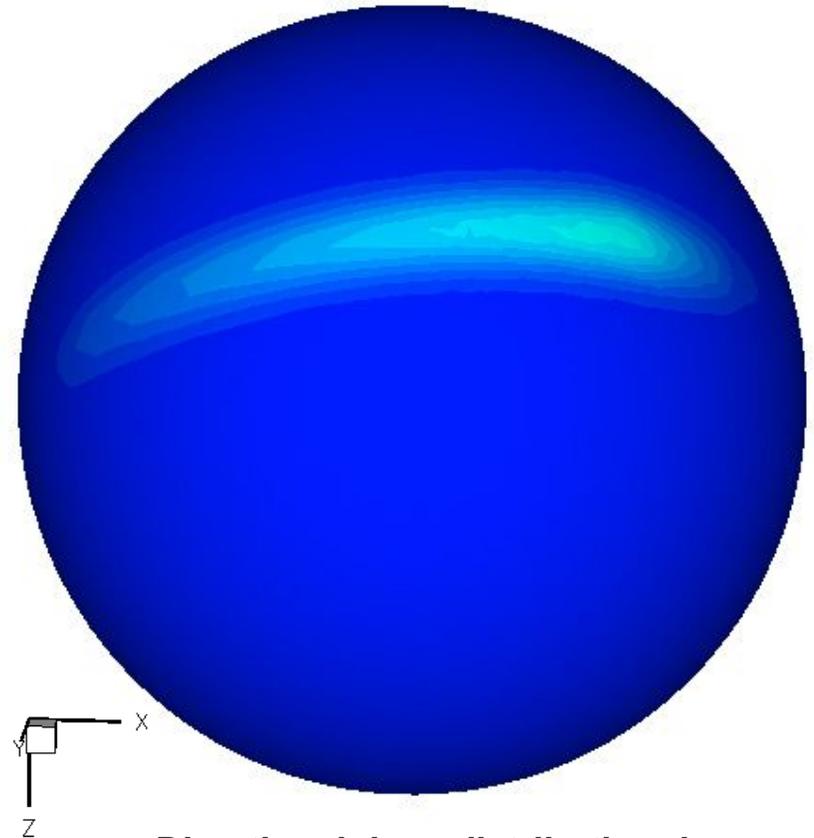
Directional dose distribution due to protons within the SM starboard Crew Quarters

Radiation Environment: South Atlantic Anomaly, Protons

400-km Solar Min. Flux (greater than 100 MeV), protons/(cm²-sec)



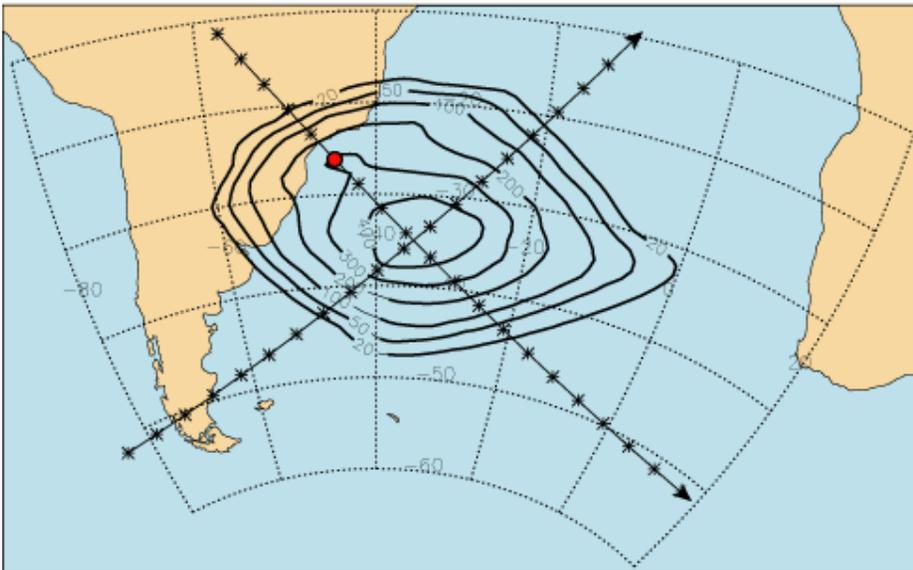
ISS ascending pass through center of SAA



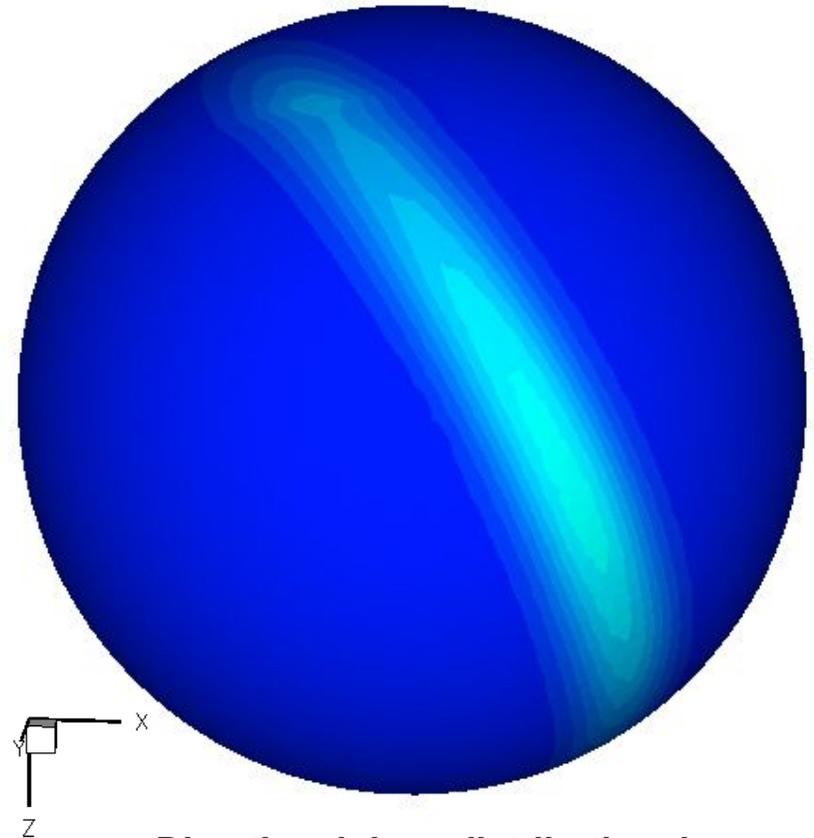
Directional dose distribution due to protons within the SM starboard Crew Quarters

Radiation Environment: South Atlantic Anomaly, Protons

400-km Solar Min. Flux (greater than 100 MeV), protons/(cm²-sec)



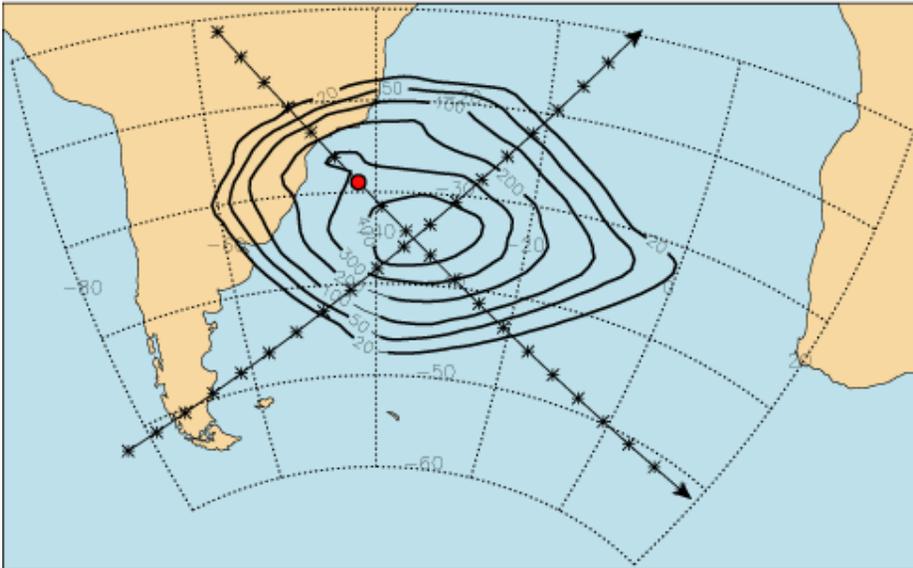
ISS descending pass through center of SAA



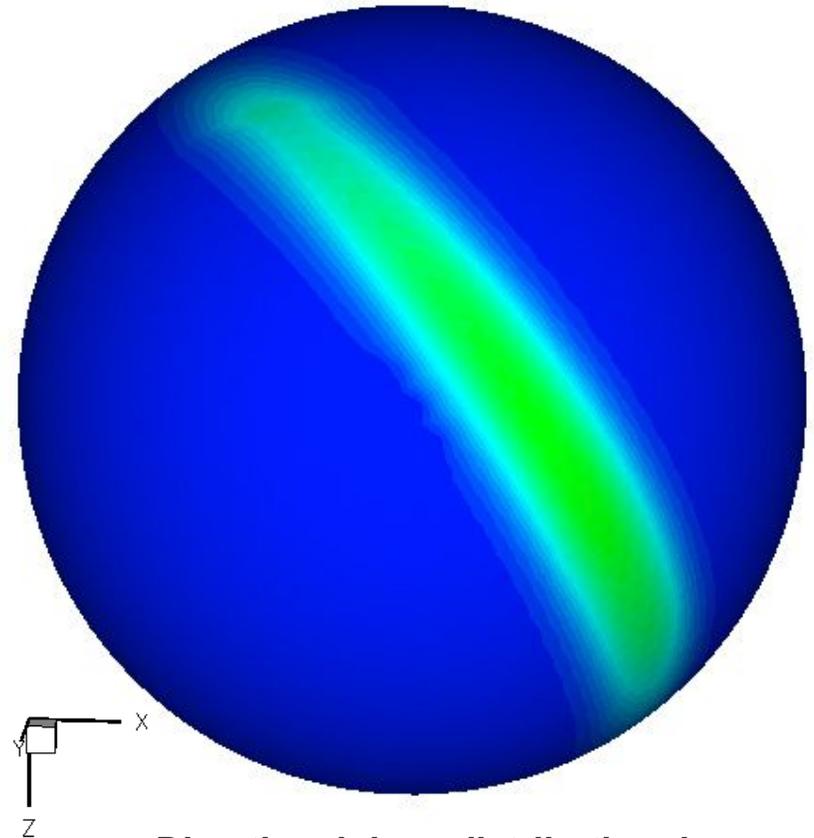
Directional dose distribution due to protons within the SM starboard Crew Quarters

Radiation Environment: South Atlantic Anomaly, Protons

400-km Solar Min. Flux (greater than 100 MeV), protons/(cm²-sec)



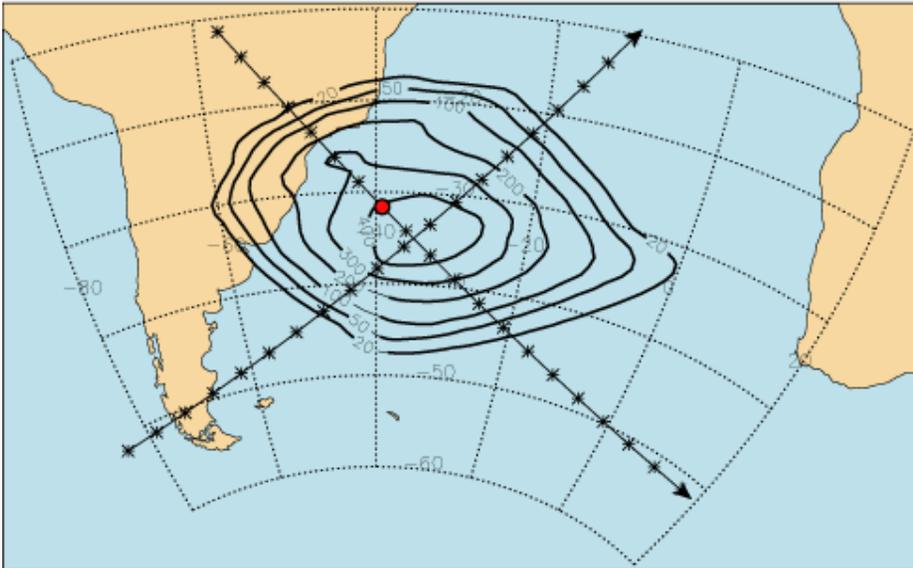
ISS descending pass through center of SAA



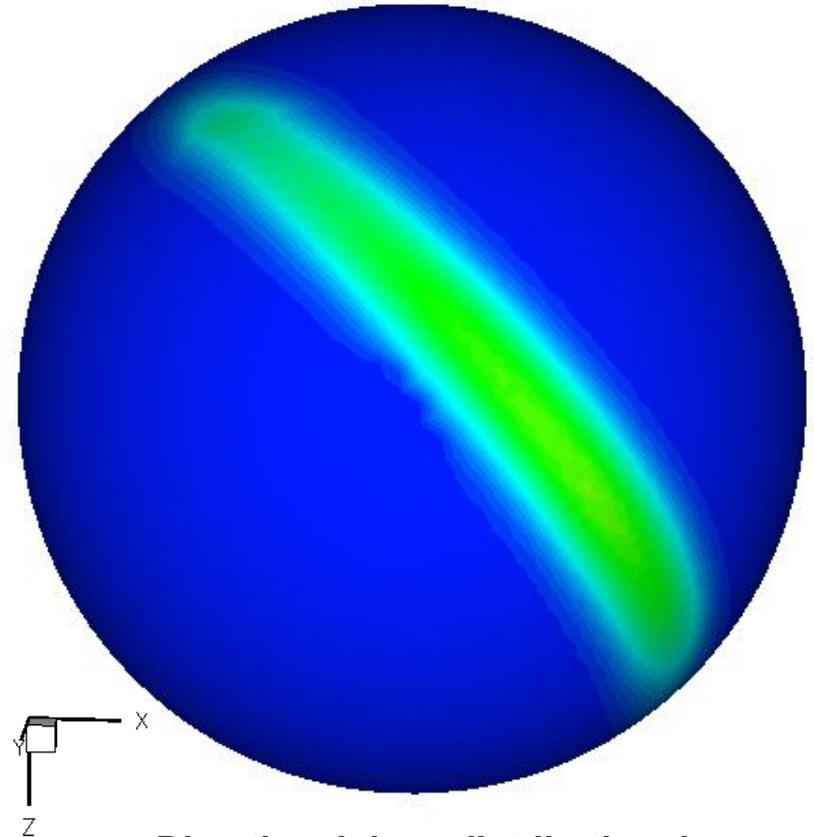
Directional dose distribution due to protons within the SM starboard Crew Quarters

Radiation Environment: South Atlantic Anomaly, Protons

400-km Solar Min. Flux (greater than 100 MeV), protons/(cm²-sec)



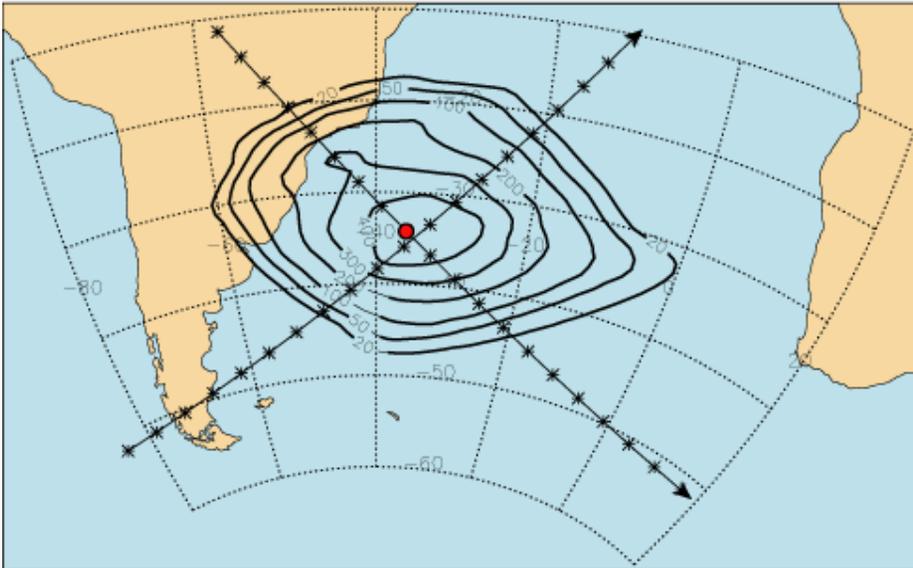
ISS descending pass through center of SAA



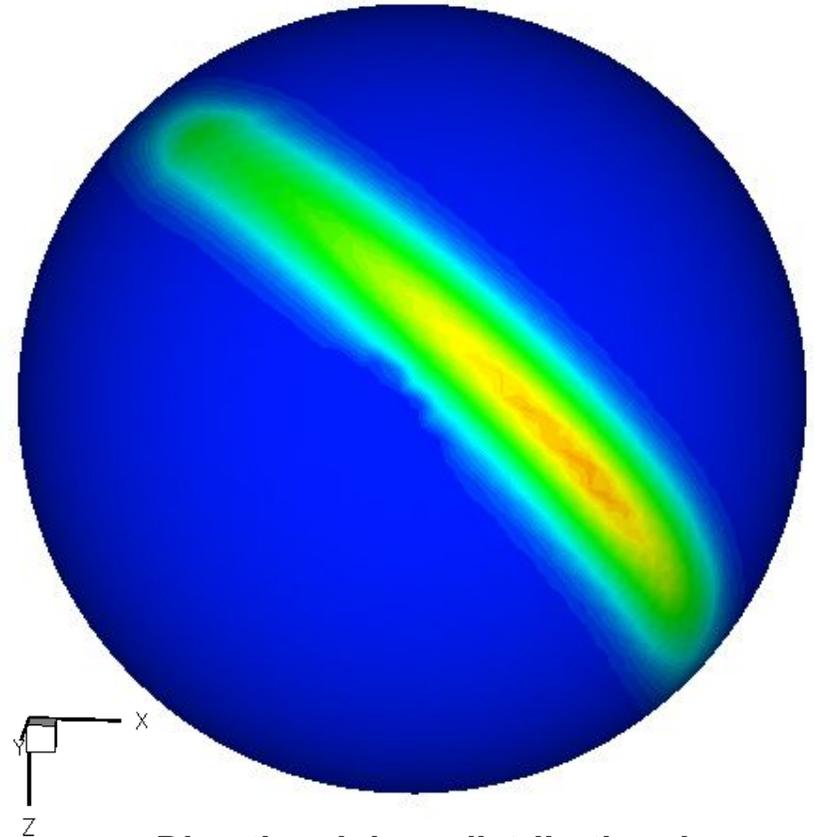
Directional dose distribution due to protons within the SM starboard Crew Quarters

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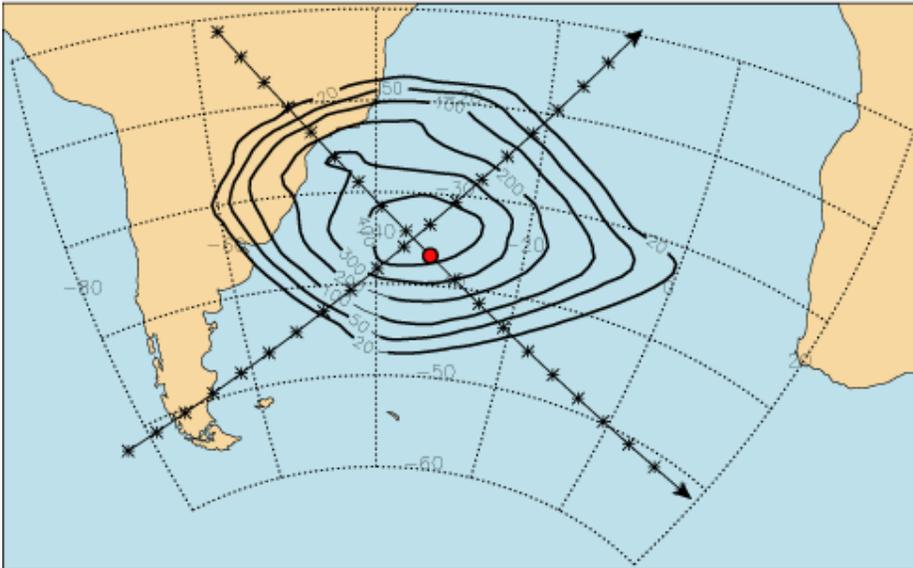
ISS descending pass through center of SAA



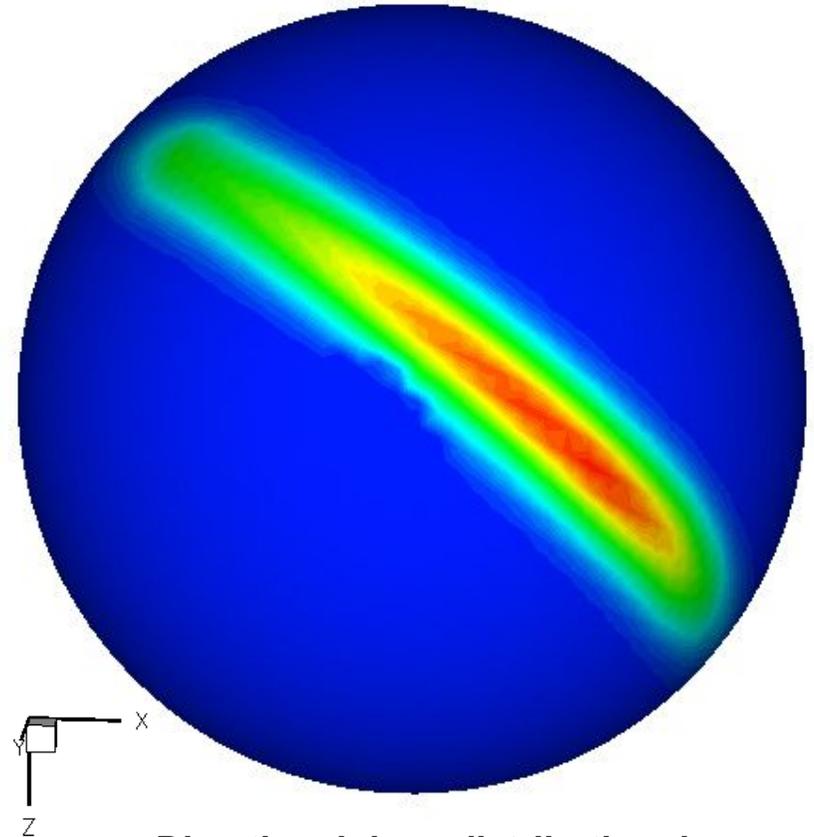
Directional dose distribution due to protons within the SM starboard Crew Quarters

Radiation Environment: South Atlantic Anomaly, Protons

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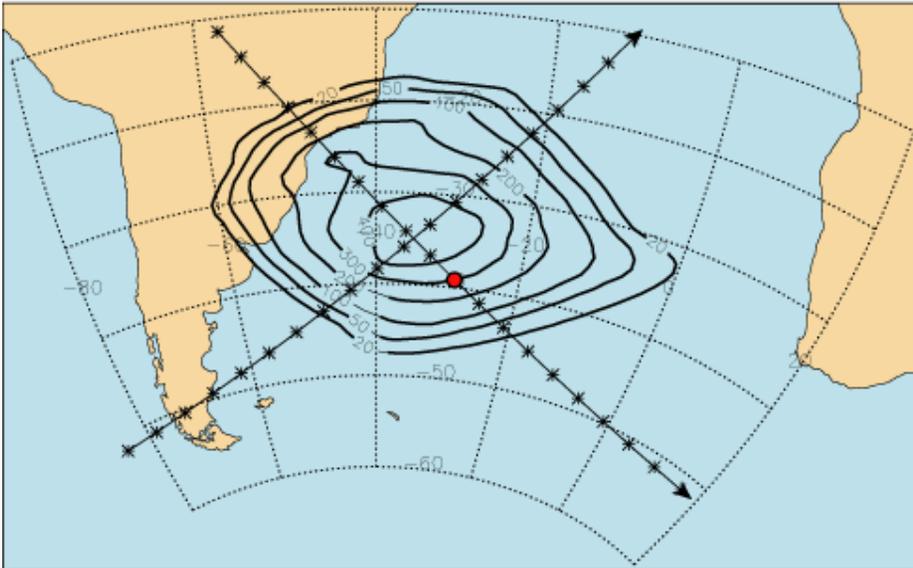
ISS descending pass through center of SAA



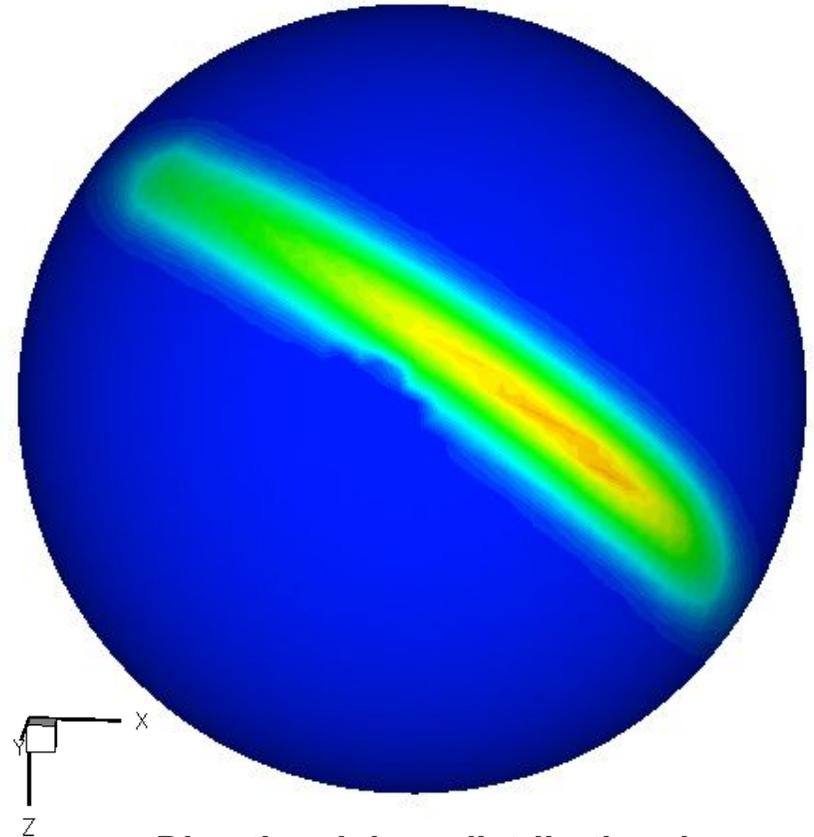
Directional dose distribution due to protons within the SM starboard Crew Quarters

Radiation Environment: South Atlantic Anomaly, Protons

400-km Solar Min. Flux (greater than 100 MeV), protons/(cm²-sec)



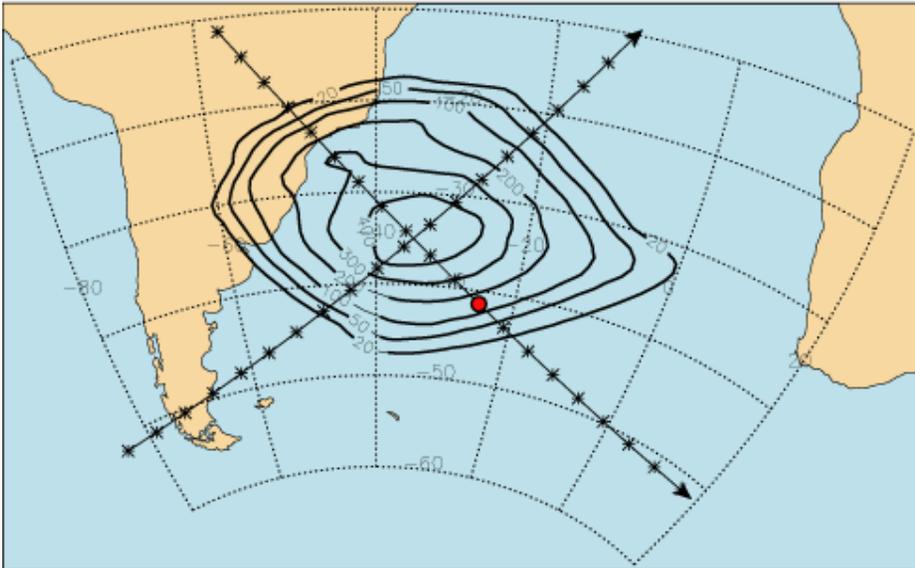
ISS descending pass through center of SAA



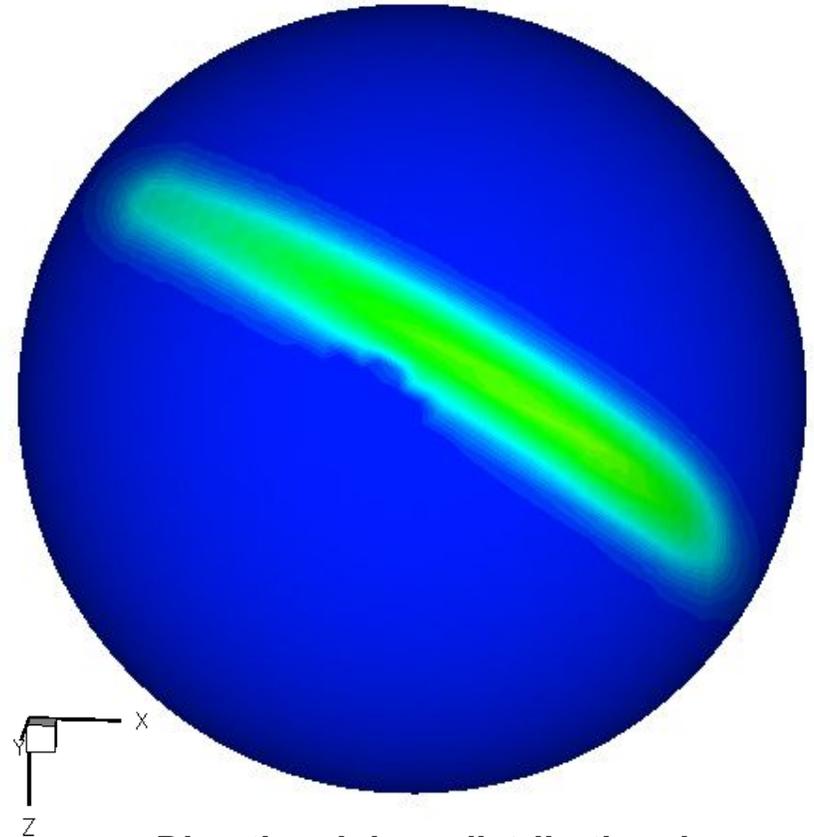
Directional dose distribution due to protons within the SM starboard Crew Quarters

Radiation Environment: South Atlantic Anomaly, Protons

400-km Solar Min. Flux (greater than 100 MeV), protons/(cm²-sec)



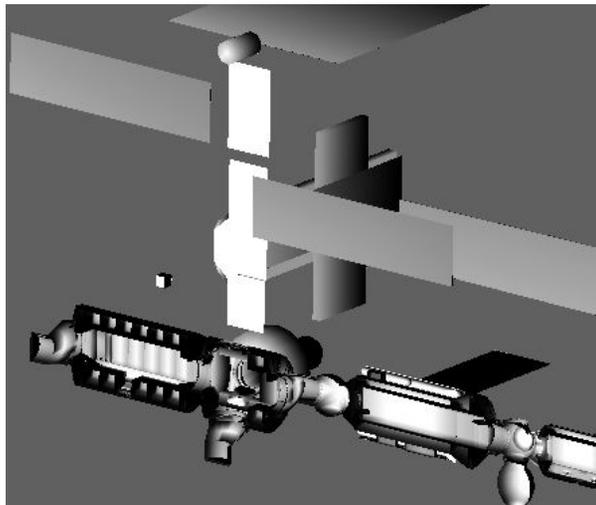
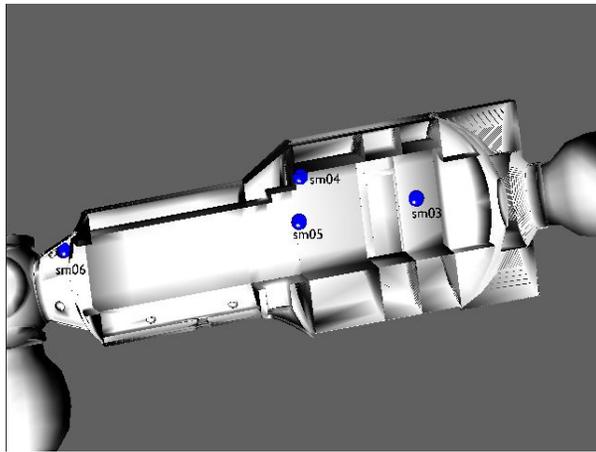
ISS descending pass through center of SAA



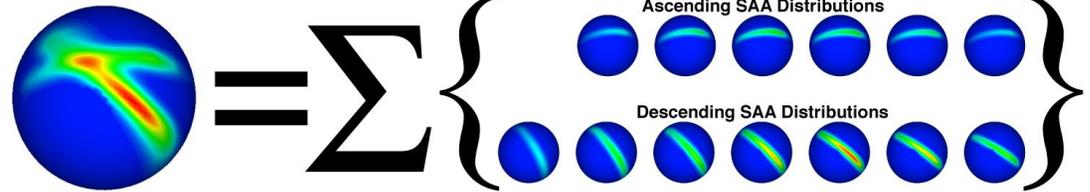
Directional dose distribution due to protons within the SM starboard Crew Quarters

Flight Validation in LEO

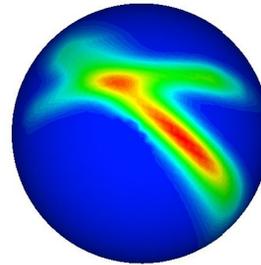
ISS 11A



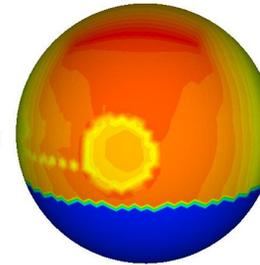
Cumulative Directional Proton Distribution



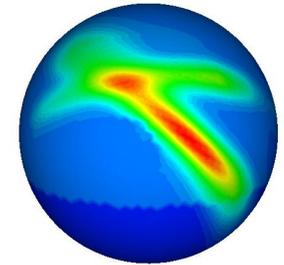
Cumulative Directional Proton Distribution



Earth-shadowed GCR Distribution



Total Daily Directional Dose Distribution



Sileye-3 detector on ISS
(particle telescope spectrometer)



Summary

- **Accelerator-based measurements produce detailed data on the radiation transport properties of materials for significant components of the GCR HZE flux.**
- **These data are being used to quantify the radiation shielding effectiveness of the candidate materials and to improve the accuracy of the models.**
- **Ultimately the models, in combination with data from radiobiological experiments now in progress, will comprise a tool for use by spacecraft and instrument designers, mission planners, flight surgeons and radiation health specialists.**

Collaborations

- **NASA SRSP Transport Consortium**
- **Flight instruments and detectors**
 - ISS partners (ICCHIBAN)
 - EriL Research/OSU (passive)
- **DSTB (NASA-MSFC)**
- **Robotic missions (MARIE)**
- **Physics support for radiation biology and bioastronautics (NSRL, LBNL NSCOR, LLUPTC, NASA-JSC)**

The Measurements Consortium

- **LBL**
Charged particle and neutron measurements and data analysis
- **NASA-MSFC**
ZDDS detector development and operation
- **NIRS**
Charged particle and neutron measurements and data analysis
- **NSCL**
Neutron measurements
- **BNL**
NSRL experiment support
- **Tohoku University**
Neutron measurements, modeling and data analysis
- **Chalmers University of Technology**
Modeling and data analysis
- **The University of Houston**
Charged particle measurements, modeling and data analysis