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Neutron Spectrometry and Dosimetry on spacecraft with passive detector system

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Radiation environment inside the aircraft

- is the result of the interaction of primary cosmic rays (p, α , e, γ rays, HZE) with shielding materials
- The evaluation is very complex because it is constituted by different components (p, e, α , n, HZE) in a wide energy range
- The most dangerous particles are HZE and Neutrons, because of their High LET, both for the astronaut health and for instrument safety

Neutron Spectrometry

- Neutron energy range covers ten orders of magnitude from thermal energies (0.025 eV) to hundred of GeV
- Different detection techniques are required

Advantages of passive detector system:

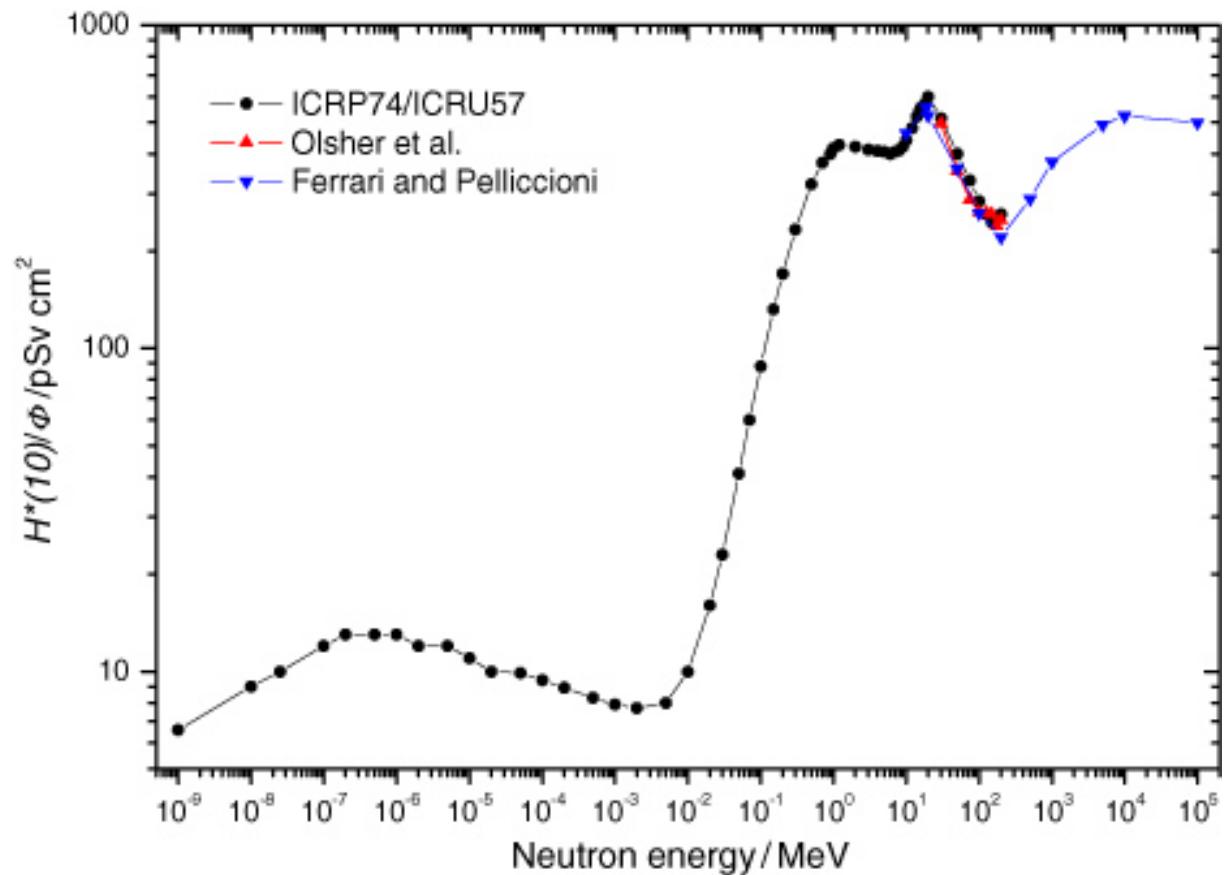
- Small size and weight
- No power supply
- No interference with electronic and instrumentation

Neutron Dosimetry

Neutron Ambient Dose equivalent $H^*(10)$ is evaluated from the energy spectra by using the ICRP 74 conversion factors H/Φ

Neutron Spectrometry and Dosimetry

Types of radiations and energy intervals	Weighting factors of radiation, W_R
Photons, all energies	1
Electrons, all energies	1
Mesons, all energies	1
Neutrons energies	5
< 10 keV	10
10 – 100 keV	20
100 keV – 2 MeV	10
2 – 20 MeV	5
> 20 MeV	
Protons	5
Heavy Ions	20



ICRP-74 - Energy dependence $H^*(10) / \Phi$ conversion coefficients

Neutron Spectrometry and Dosimetry

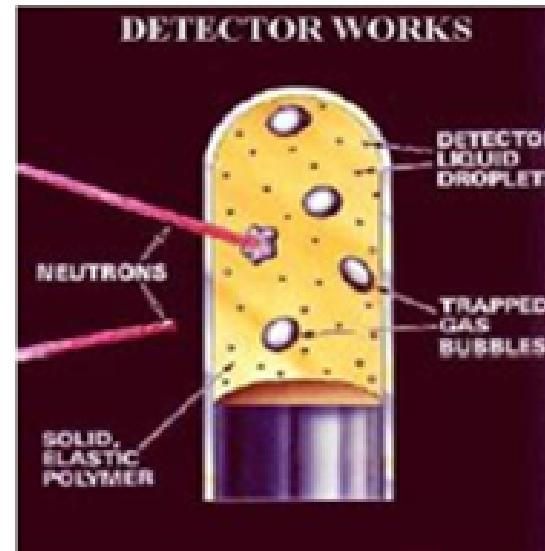
BUBBLE DETECTORS

(BTI , Ontario, Canada)

1-BDT (0.025 eV- 4 eV)

2-BDT shielded (2 eV- 20 keV)
(Cd 2 mm, Borate Polyethylene 5 mm)

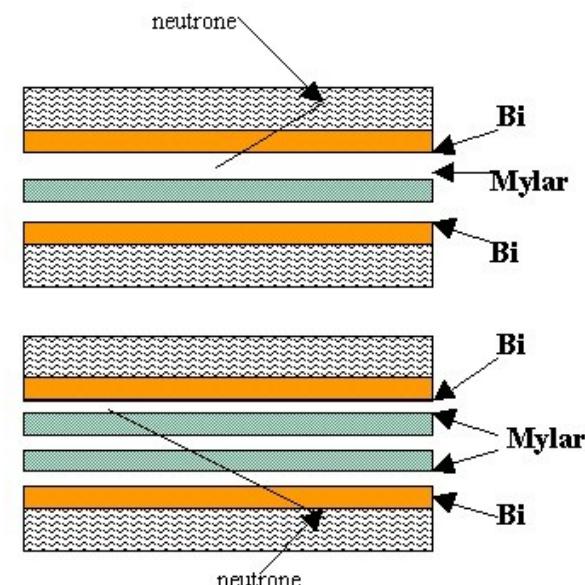
3-BDS-10 (10 keV-20 Mev)

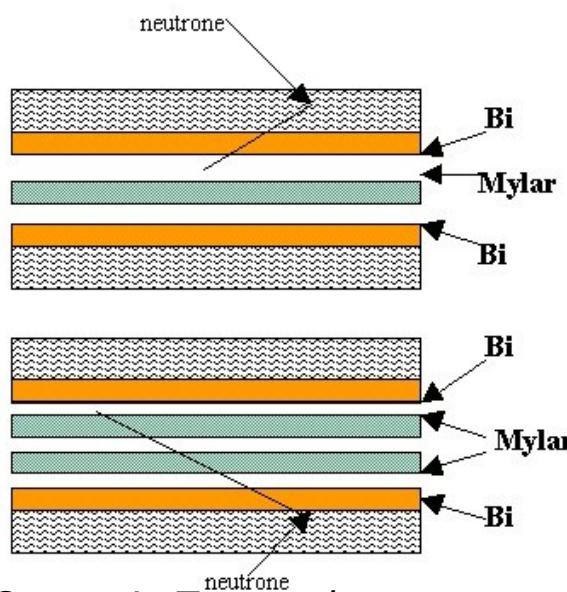
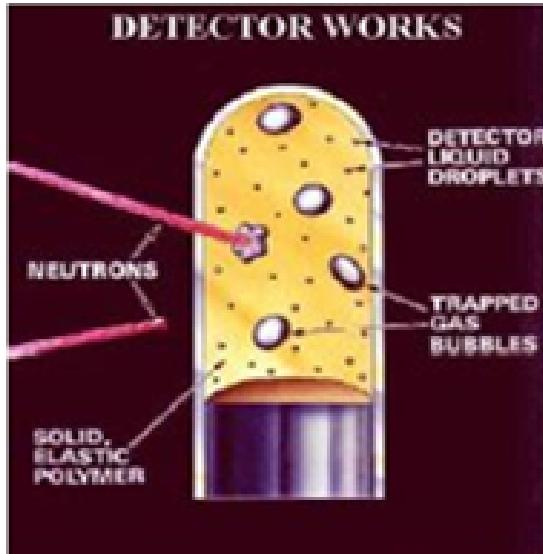


TRACK DETECTORS

4-Polycarbonate foils (Curbell Plastic, USA)

5-Bi209 stack (50 MeV-100 GeV)





W. G. Cross L. Tommasino

Dosimetry of High Energy Neutrons and Protons
by 209Bi Fission. Radiation Protection Dosimetry
70(1) April 1997

Bubble detectors consist of polycarbonate vials ($d=1.8$, $h=8$ cm) filled by tissue-equivalent gel (H_2O, C, F, N) where superheated freon drops, in a metastable state, are dispersed.

Neutrons interact with gel creating charge secondary particles (α and p) which transfer their energy to the gel producing a state change with bubbles formation.

BDT - $0.025 \text{ eV} < E_n < 4 \text{ eV}$

BD-PND – $100 \text{ keV} < E_n < 20 \text{ MeV}$

BDS 10 - $10 \text{ keV} < E_n < 20 \text{ MeV}$

- **Bi209 Stack detector** is based on bismuth fission.
- Fission fragments produced by the interaction of high energy neutron with Bismuth nuclei deposit their energy in the Mylar detector producing tracks that become visible after a chemical attack.
- The analysis of tracks permit to determine the incident neutron fluence.

A passive detector system for neutron spectrometry

$0.025 \text{ eV} < E_n < 100 \text{ GeV}$

BUBBLE DETECTORS (BTI , Ontario, Canada)

1-BDT ($0.025 \text{ eV}- 4 \text{ eV}$)

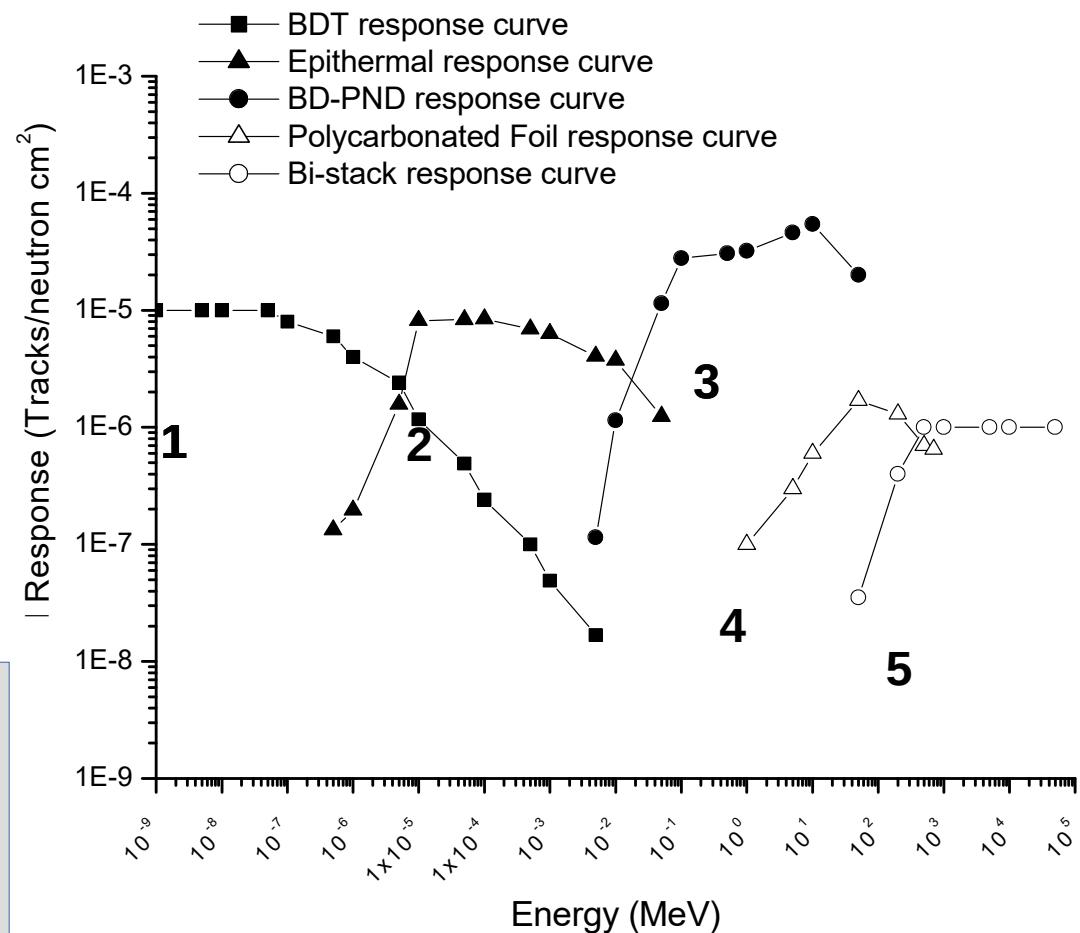
2-BDT shielded ($2 \text{ eV}- 20 \text{ keV}$)
(Cd 2 mm, Borate Polyethylene 5 mm)

3-BDS-10 ($10 \text{ keV}-20 \text{ Mev}$)

TRACK DETECTORS

4-Polycarbonate foils ($10 \text{ MeV}-100 \text{ MeV}$)

5-Bi209 stack ($50 \text{ MeV}-100 \text{ GeV}$)



BUNTO unfolding code

$$M_j = \int_{E_{\min}}^{E_{\max}} R_j(E) \Phi_E(E) dE \quad j = 1 \dots m$$

$$M_j = \sum_{i=1}^n R_{ij} \Phi_i \quad i = 1 \dots N$$

M_j : detector response

$R_j(E)$: response curve values

$\Phi_E(E)$: differential fluence
distribution of neutron energy

m : number of energy thresholds

n : number of energy intervals

- It is based on BUNKI's algorithm (SPUNIT), which finds the non negative solution through a iterative perturbation procedure.
- It can find a solution using a starting information on the spectrum shape, but it can also work in lack of information on the initial spectrum.
- The solution $\Phi_E(E)$ is the calculated mean from a number of spectra obtained by a random generation of M_j sampled on a normal distribution, whose parameters m and s are the experimental reading and the associated uncertainty of the j^{th} detector.
- It can be used with different spectrometry systems (like the extended energy and the BDS), if the response matrix of the detectors is known.

Experimental test

Low Energy Neutrons
(Thermal -20 MeV)

Medical e-linac accelerators

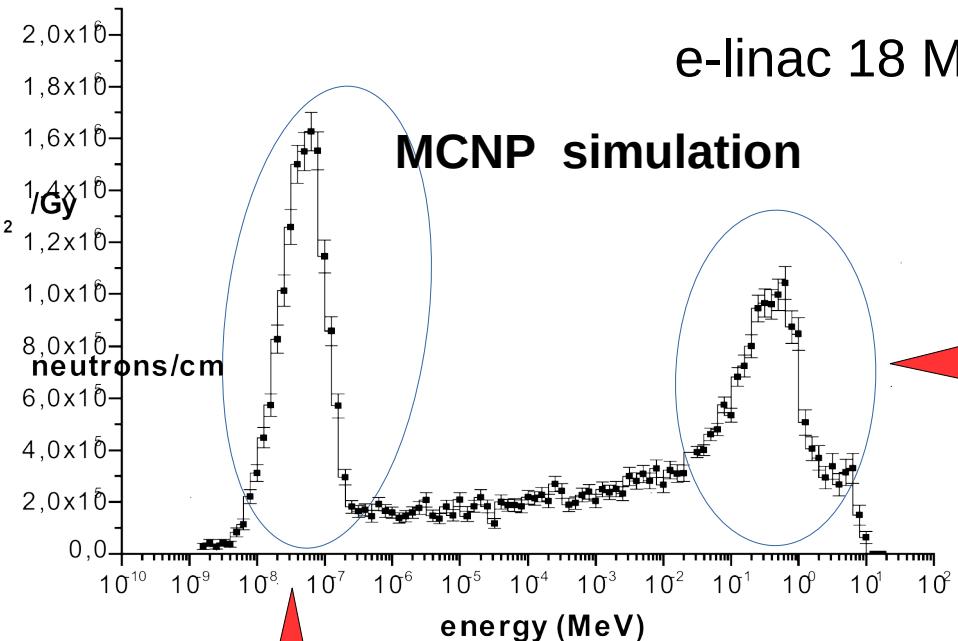
Bubble detectors

High Energy Neutrons
(20 MeV-100 GeV)

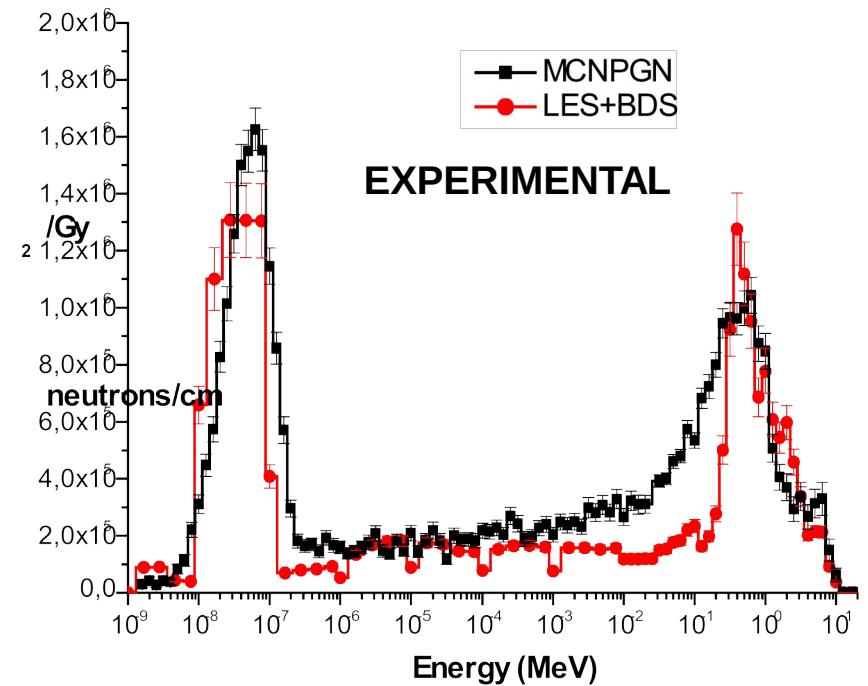
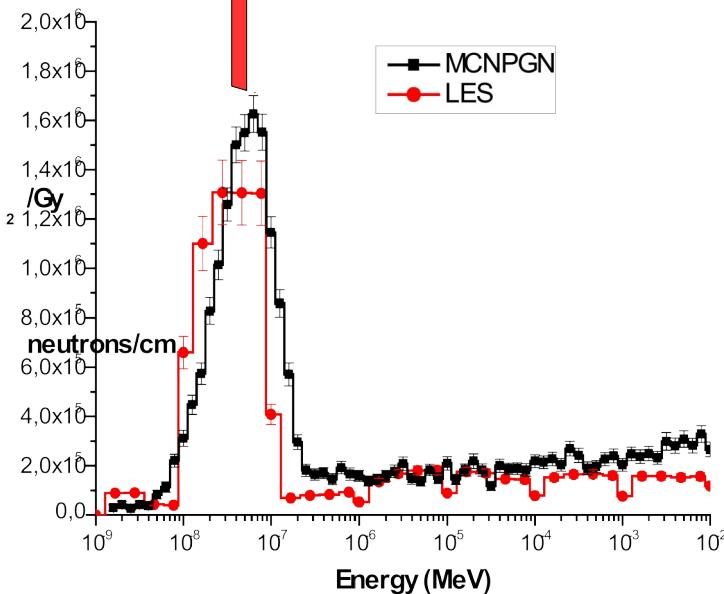
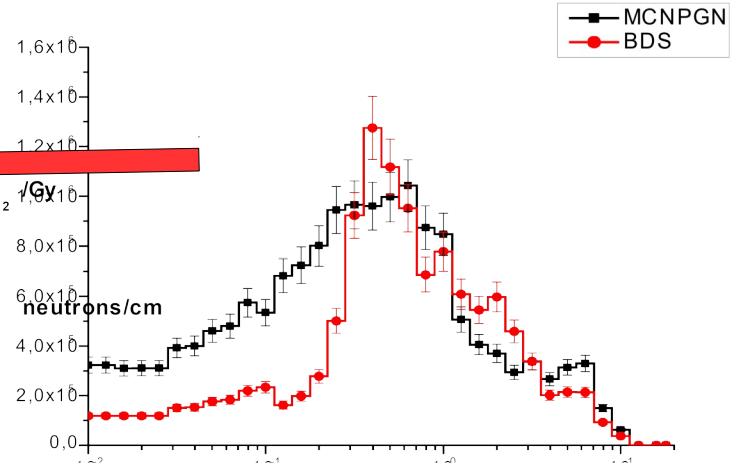
- 1- CERF Facility
- 2- High altitude Laboratories (3 -5 km)
- 3- High altitude flights (10-12 km)
- 4- Stratospheric balloon flights (20-40 km)

Track detectors

Test neutron spectrometer – Low Energy (0.025 < En < 20 KeV



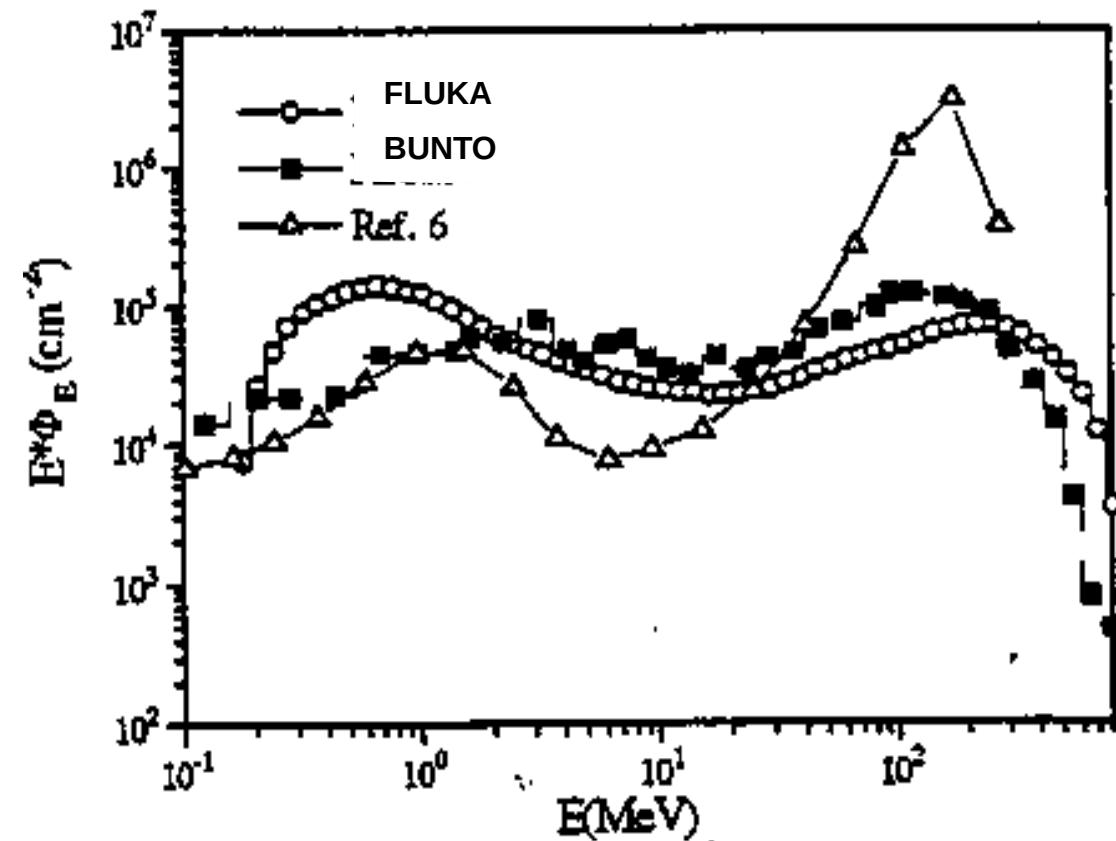
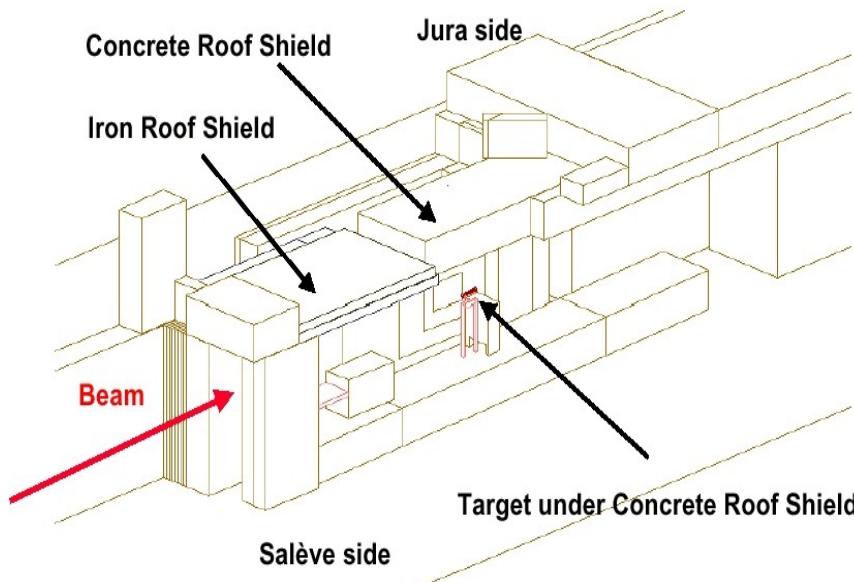
e-linac 18 MV- Elekta precise- prostate treatment
Neutron spectrum in the patient body



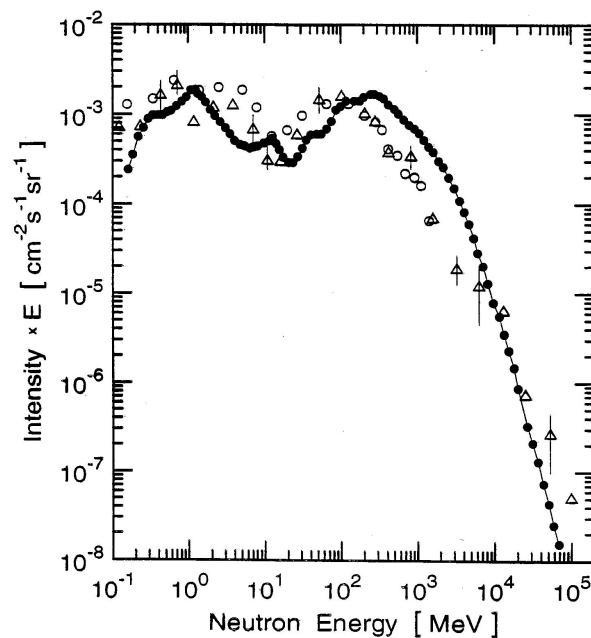
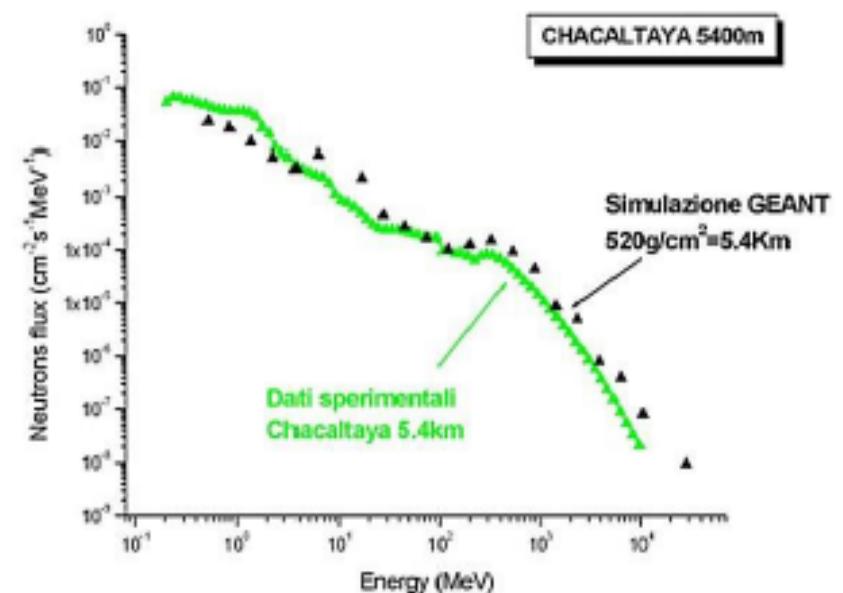
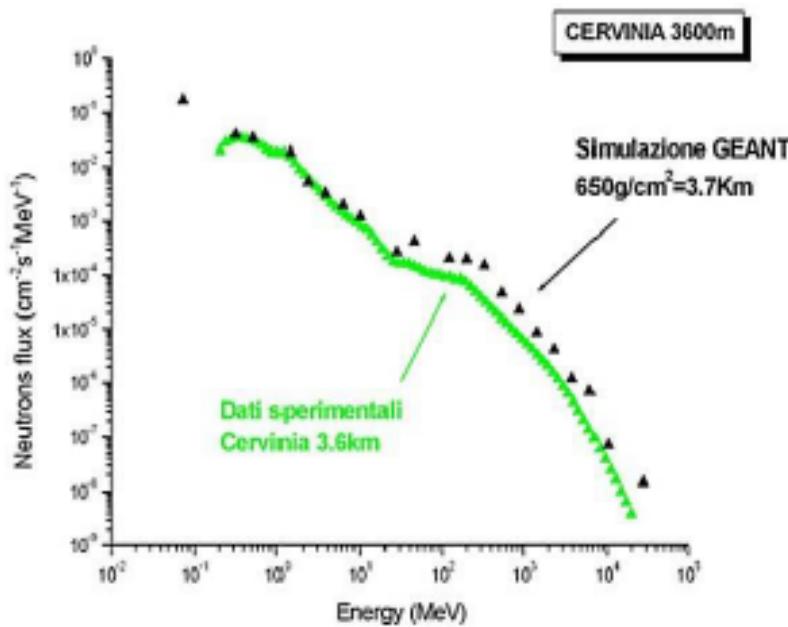
Test at CERF facility

High energy neutrons 20 MeV-100 GeV

- The calibration of the passive detector system has been performed at CERN (T14 position, H-6SPS beam): this facility is a reference field for the calibration of neutron detection systems to be used in the cosmic ray field.
- The passive detector results, unfolded with the BUNTO code, are compared with the FLUKA simulation of the experiment



Neutron spectra evaluated by passive spectrometer



Manfredotti C, Zanini A., et al. *Cosmic Rays at Earth*
Editor P.K.F. Grieder pp 207-209
ISBN: 9780444507105
Elsevier Science, July 2001



Neutron detector set

- Bubble detector BDT thermal neutrons
- Bubble detector BDT shielded epithermal neutrons
- Bubble detector BDS-10 intermediate neutrons
- Polycarbonate foils fast neutrons
- Fission Bi209 stack relativistic neutrons



Experimental measurements

SHUTTLE STS134 MISSION

GROUND CONTROL
2 BDT
2 BDPND



FLIGHT CONTROL
3 BDT
3 BDPND



NEUTRON DETECTORS INSIDE ENDEAVOUR SHUTTLE



- Bubble detectors

- BDT

- BDT shielded

- BDS-10

- thermal neutrons

- epithermal neutrons

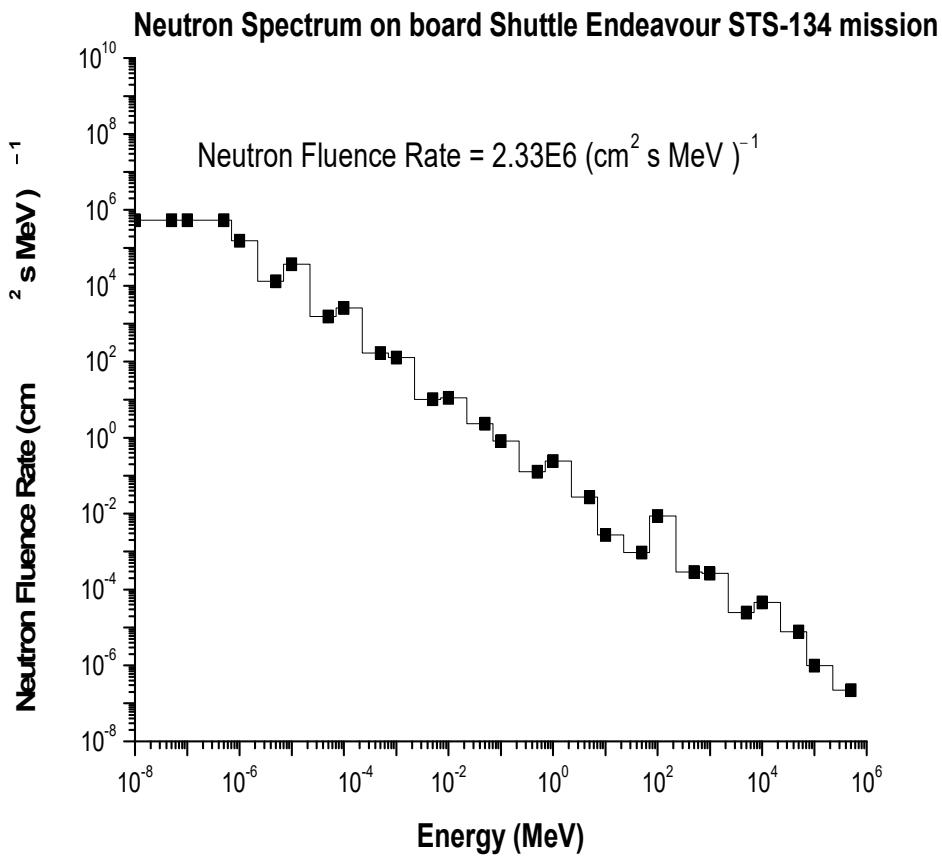
- fast neutrons

- Fission Bi209 stack

- relativistic neutrons

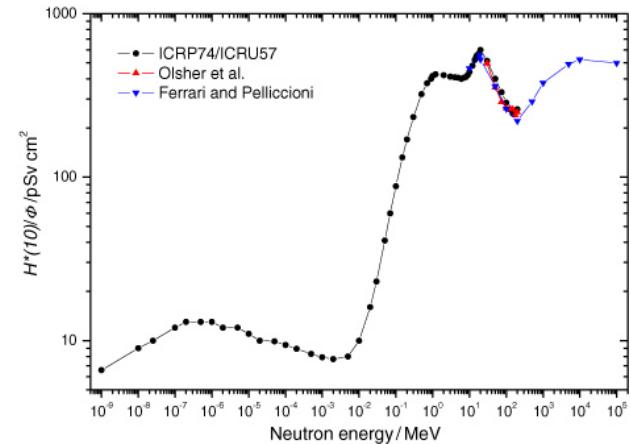


RESULTS

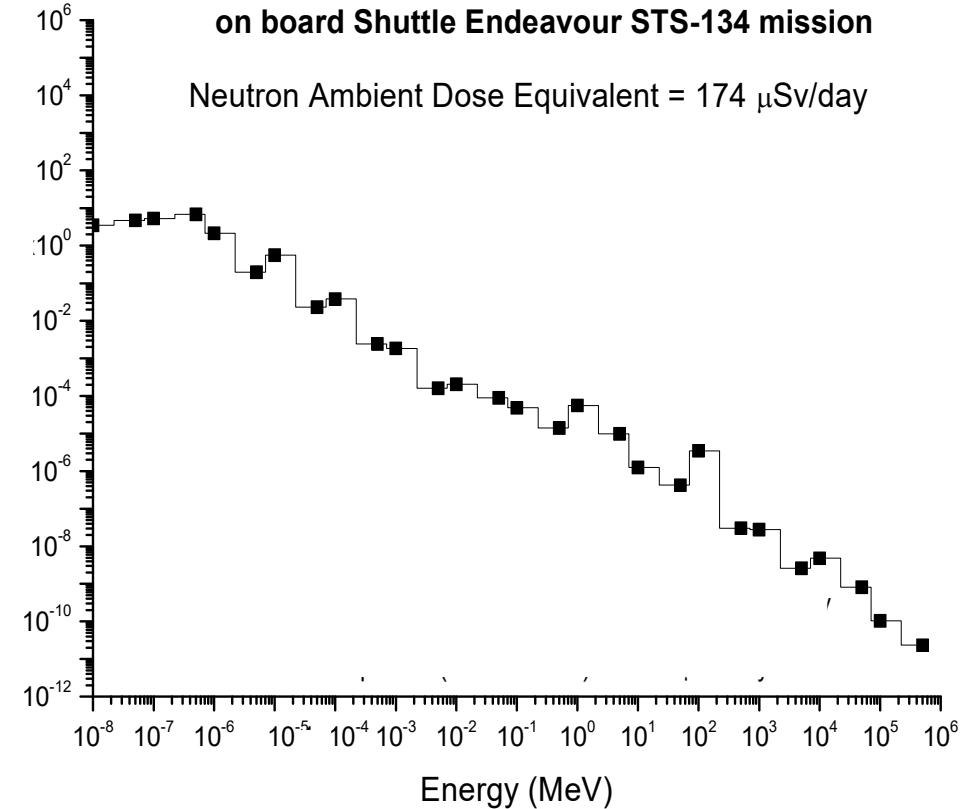


Neutron Ambient Dose Equivalent rate:
 $174 \pm 38 \mu\text{Sv/day}$

ICRP 74 conversion factors
 $H^*(10) / \Phi \text{ [pSv cm}^2]$

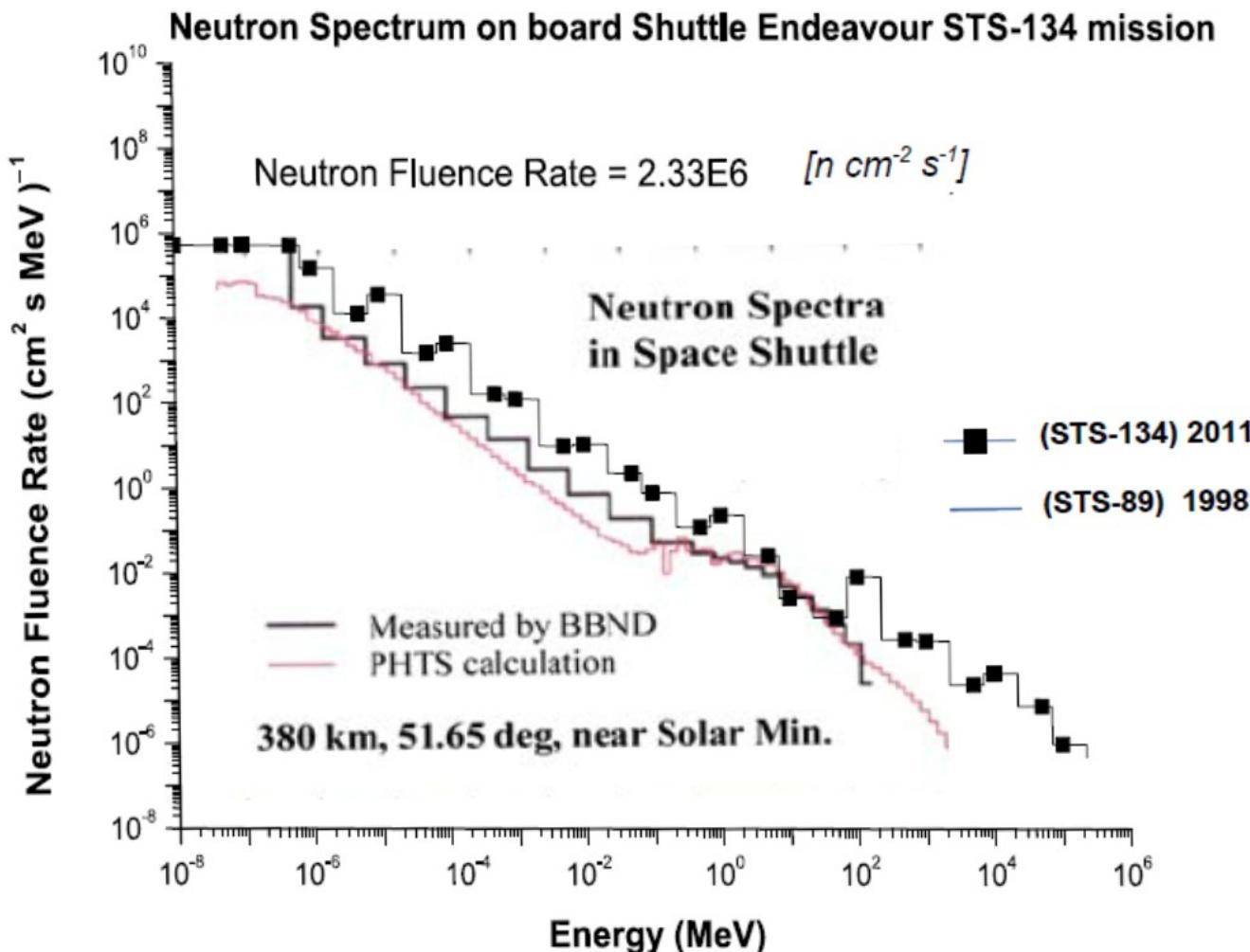


**Neutron Ambient Dose Equivalent
on board Shuttle Endeavour STS-134 mission**





Neutron spectra comparison STS-134 STS 89



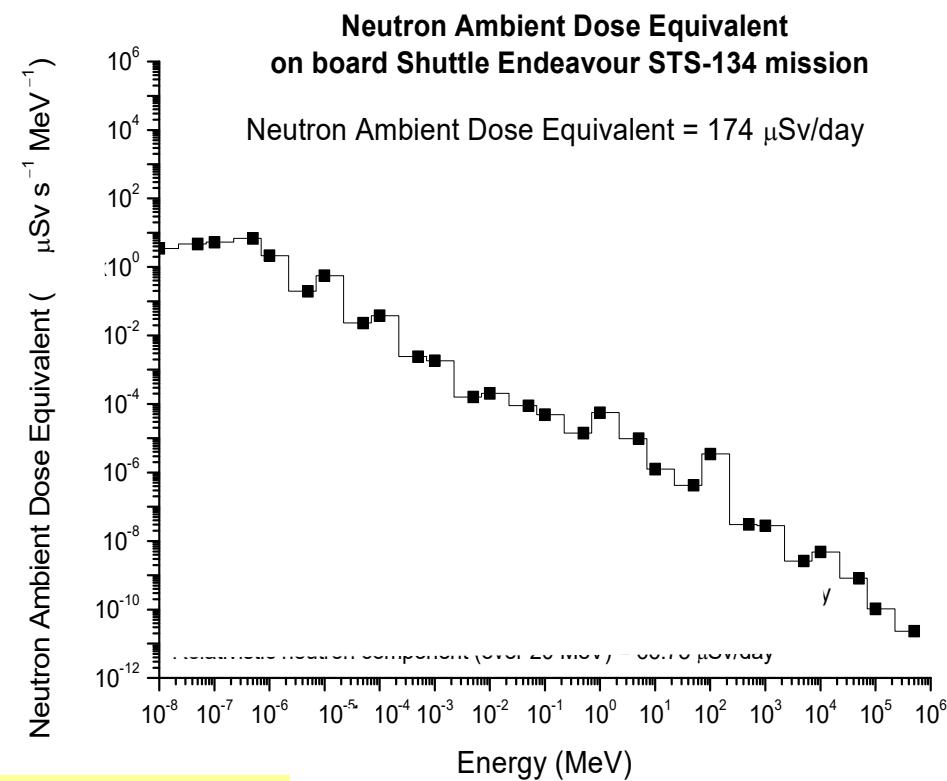
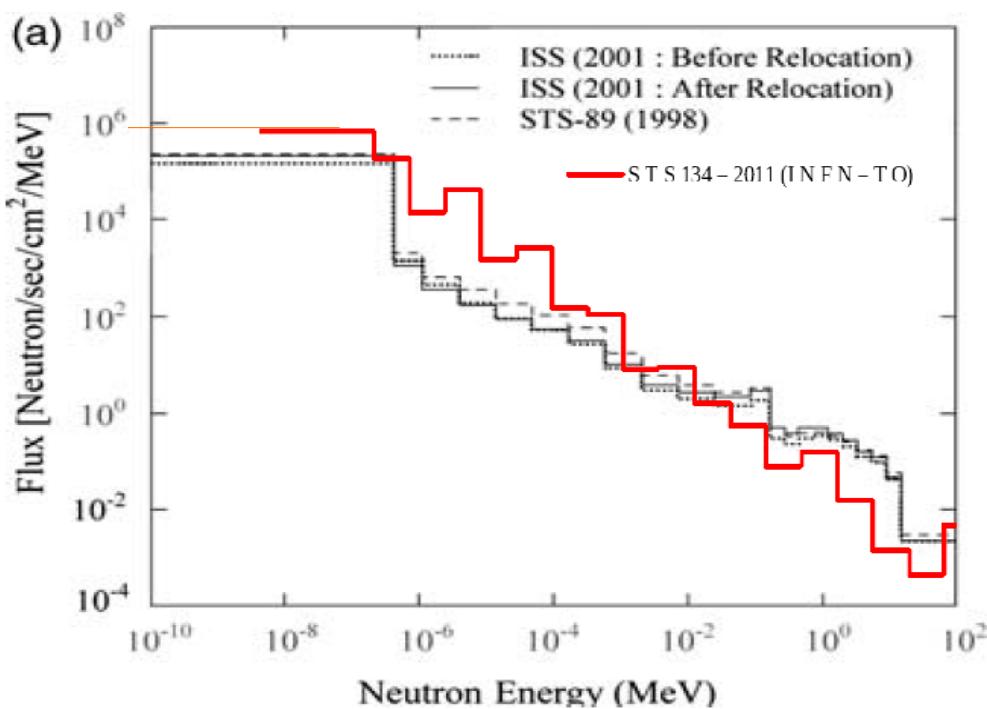
STS-134 (16/5-1/6/2011)

51,65°
ISS station

STS -89 (22/1 – 31/1/1998)

51,65 °
MIR station
(1,2)

- 1) Koshiishi, H. Evaluation of the neutron radiation environment inside the International SpaceStation based on Bonner Ball neutron detector experiment. Radiat. Meas. 42, 1510–1520 (2007).
- 2) B.J. Lewis et al., Review of bubble detector response characteristics and results from space Radiation Protection Dosimetry 150(1):1-21 · September 2011

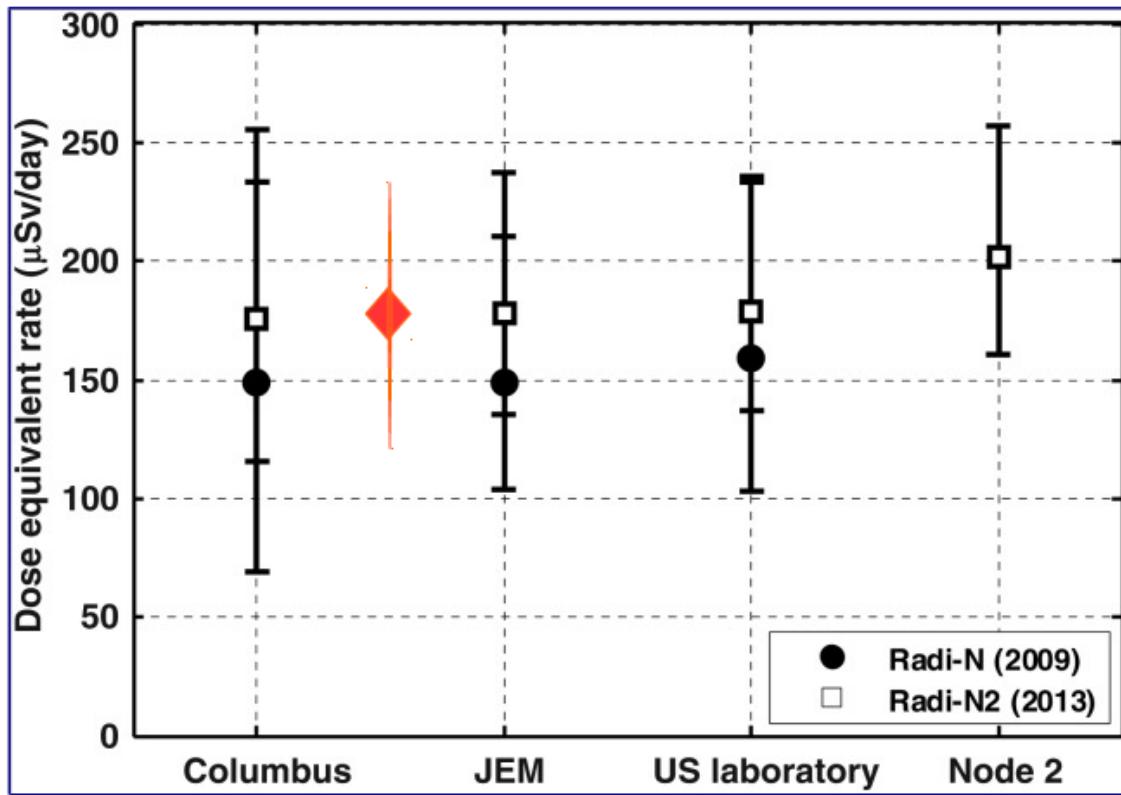


SHUTTLE ENDEAVOUR STS 134

H*(10) 0.025 eV -100 GeV

BDT	(0.025eV-10keV)	14 +/- 3 $\mu\text{Sv/d}$
BDPND	(10keV-20MeV)	93 +/- 20 $\mu\text{Sv/day}$
Stack Bi209	(20MeV-200GeV)	67+/- 15.4 $\mu\text{Sv/day}$

H*Total = 174 +/- 38 $\mu\text{Sv/day}$



Measurements performed in various positions of ISS with space bubble-detector spectrometer (SBDS).

December 2012 -October 2013.

M.B. Smith et al., Bubble-detector measurements of neutron radiation in the international space station: ISS-34 to ISS-37. 2016 Feb; Rad. Prot.Dos. 168(2): 154–166.

◆ Endeavour STS 134

CONCLUSIONS

A passive detector system , based on bubble detectors and track detectors, Has been used to evaluate the Neutron Ambient Dose Equivalent in a wide energy range (thermal- 100 GeV) on board of Endeavour shuttle STS 134.

The obtained values for $H^*(10)$ are in agreement both with previous neutron energy spectra measurements (Koshiishi, H., 2007) and with recent integral measurements

(M.B.Smith et al., 2016)

Thank you for your attention