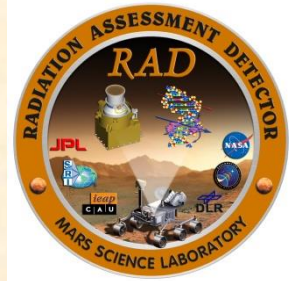




MSL/RAD Radiation Measurements on the Surface of Mars on the Way to Solar Maximum – New Findings & Updates

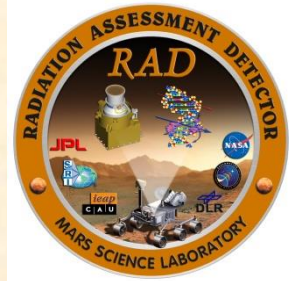


Bent Ehresmann, D. M. Hassler, C. Zeitlin
and the MSL/RAD Team

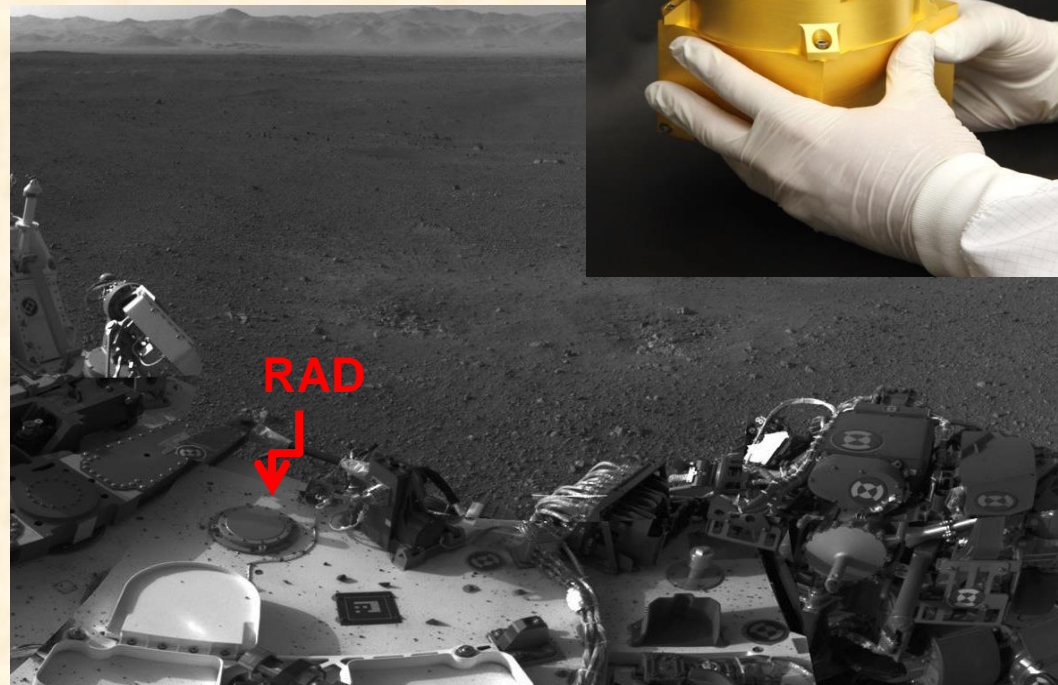
27th WRMISS Workshop
5 September, 2024



The Radiation Assessment Detector (RAD)

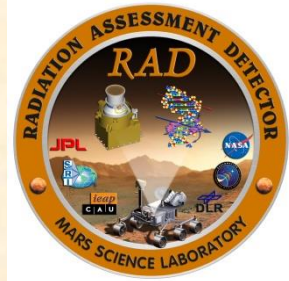


- RAD has been measuring the radiation environment on Mars on board NASA's Curiosity rover since August 2012, now spanning **one whole solar cycle**
- Knowledge of the radiation environment is crucial for the planning of **future human exploration of Mars**
- Exposure to radiation from Galactic Cosmic Rays (**GCRs**) and Solar Energetic Particles (**SEPs**) remains one of the major risks for manned space flight

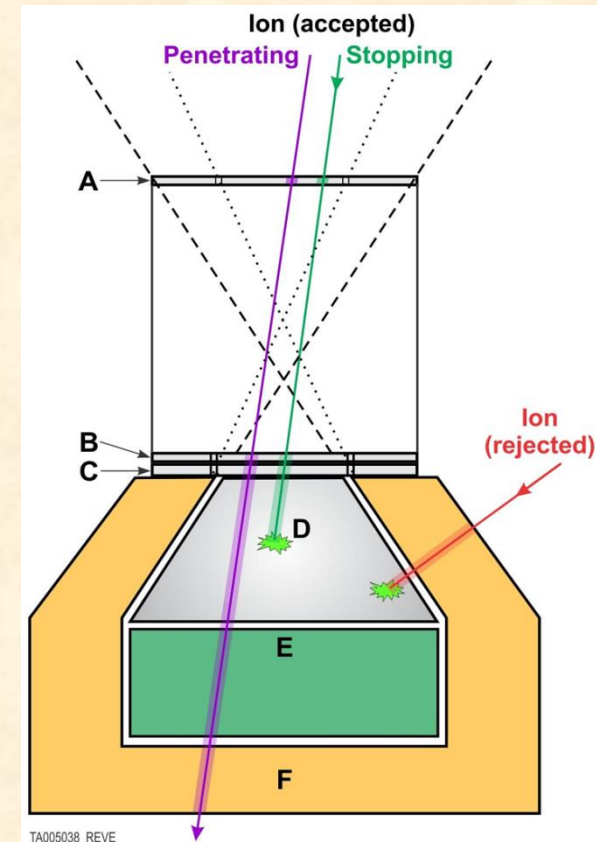




The Radiation Assessment Detector (RAD)



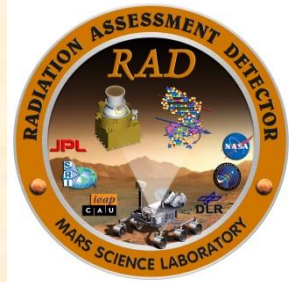
- The RAD Sensor Head consists of the following detectors
 - Three 300 μm thin Si detectors (**A**, **B**, **C**),
 - CsI scintillator (**D**),
 - Plastic (tissue-equivalent) scintillator (**E**),
 - Plastic scintillator (**F**) used for anti-coincidence
- The coincidence of detectors **A&B** defines the acceptance angle for charged particle detection
 - A&B FOV $\sim 30^\circ$ from zenith
- **D & E** are used to detect *neutral particles* (in anti-coincidence with detectors C & F)



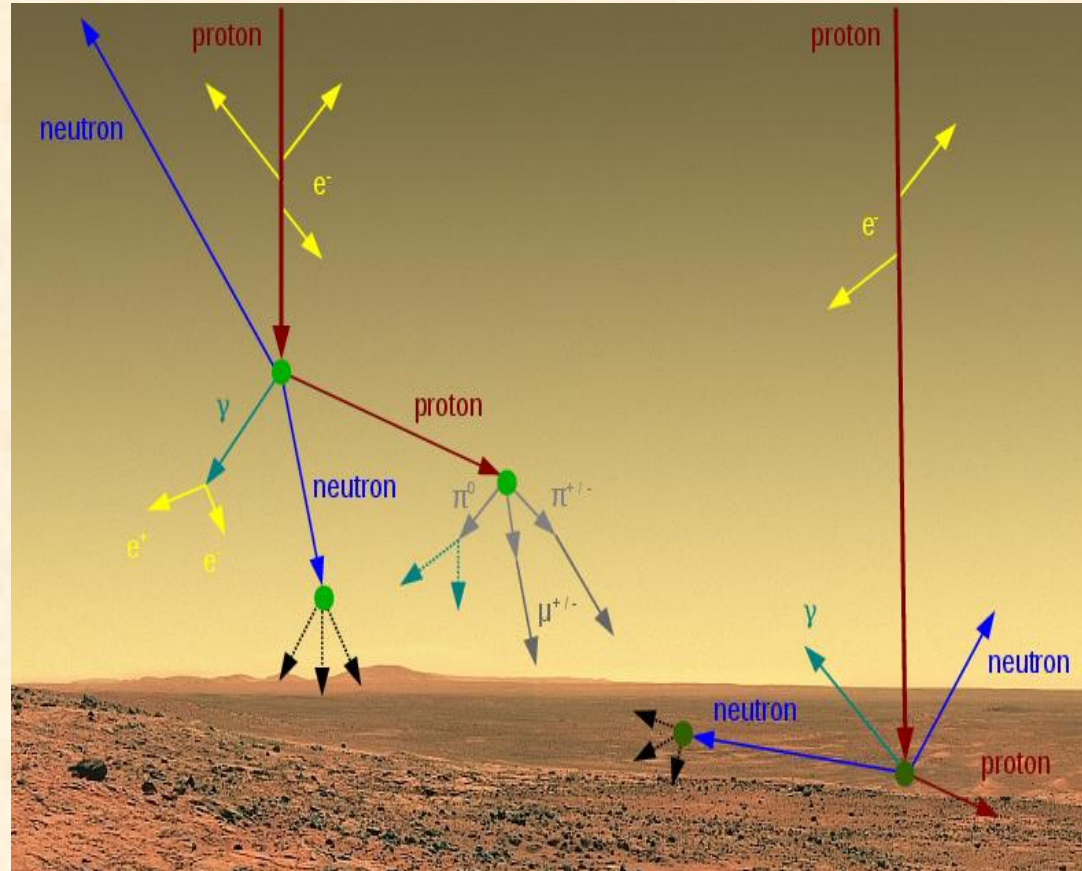
Schematic of the RAD sensor head



Radiation Environment on Mars



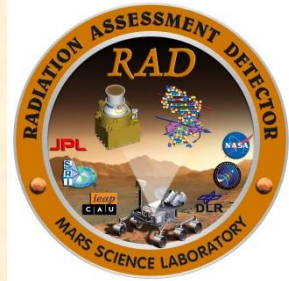
- Radiation environment on the surface of Mars differs greatly from Earth
- Mars lacks a global magnetic field and its **atmosphere** is very **thin**
- **GCRs** and high-energy **SEPs** can propagate deep into the atmosphere and soil and interact with the nuclei → intense radiation field of **charged and neutral particles** on the surface
- Protons need ~**150 MeV** to reach the surface!



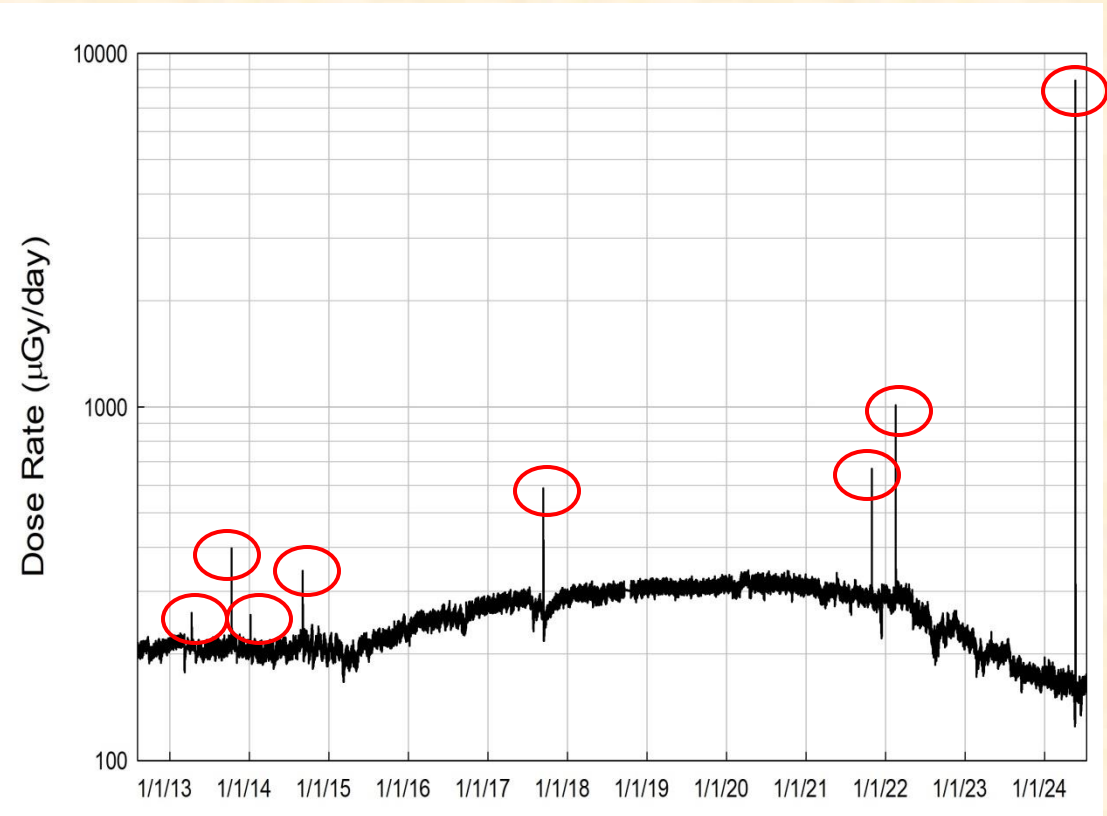
Schematic of GCR protons propagating through the atmosphere and interacting with the nuclei of soil and atmosphere.



RAD Dose Rate Measurements over one Solar Cycle



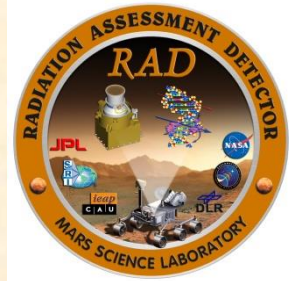
- Dose rates range from ~ **340 $\mu\text{Gy/day}$** (solar min) to ~ **160 $\mu\text{Gy/day}$** (current solar max)
- Previous solar max dose rates **deeper** than during end of last solar max (2012)
- Dose rates stem mainly from **GCR**...
- ... and vary with **solar modulation & Martian pressure** cycle
- RAD has detected **eight SEP events** so far (red circles)



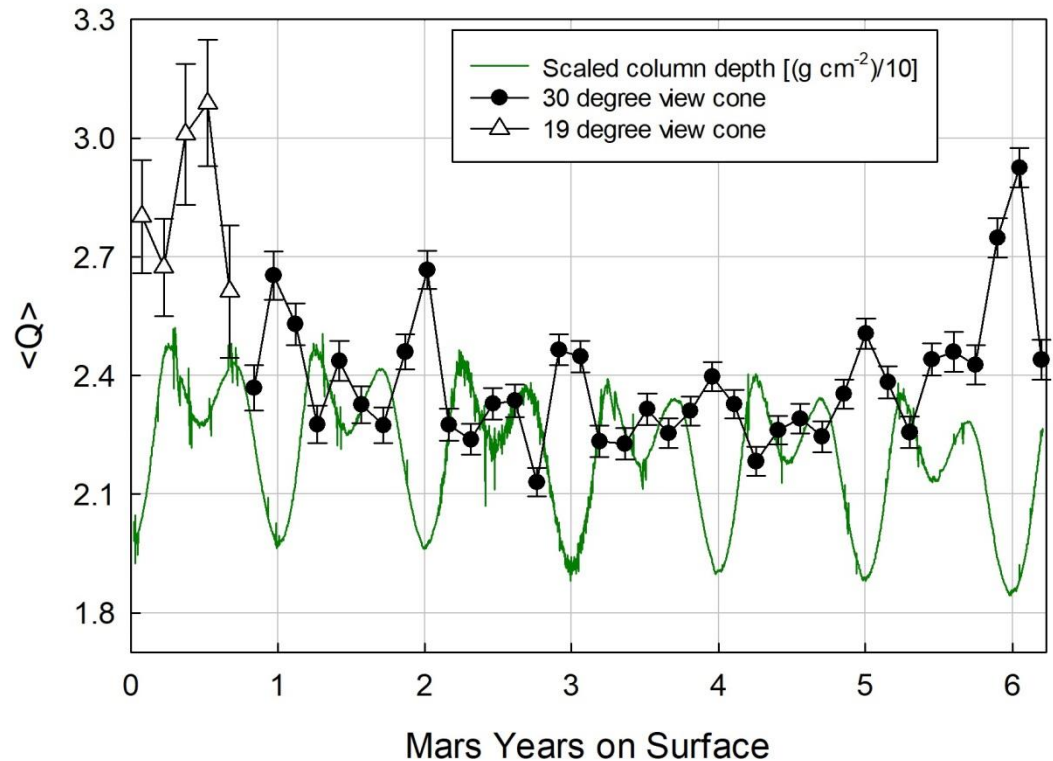
RAD dose rate measurements in tissue-equivalent plastic scintillator E. The eight solar events detected by RAD are highlighted in red.



Radiation Quality Factor $\langle Q \rangle$



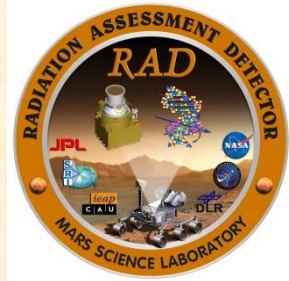
- $\langle Q \rangle$ is a measure for the **biological effectiveness** of the radiation field
- ... calculated from RAD **LET spectra** and **$Q(LET)$**
- $\langle Q \rangle$ is affected by both **solar modulation** and **Martian pressure**
- Decrease in pressure leads to increase in $\langle Q \rangle$
- Increase in solar modulation (solar max) leads to increase in $\langle Q \rangle$
- $\langle Q \rangle$ ranges from ~ 2.1 (solar min) to **3.1** (solar max)



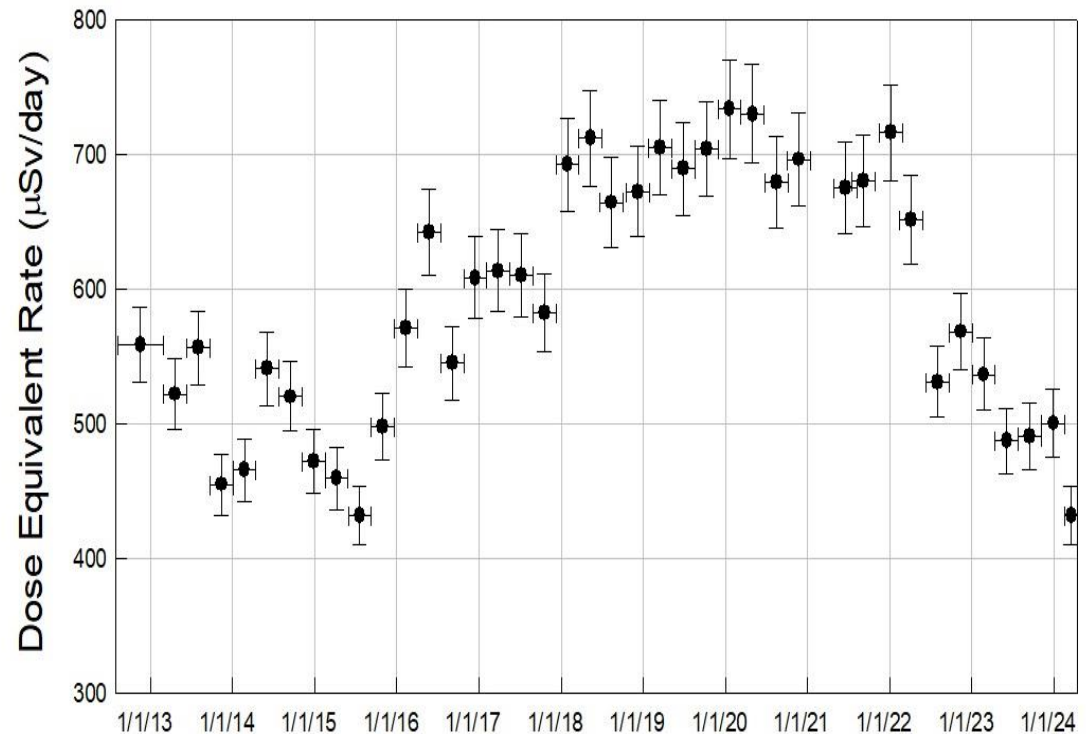
(Green) Vertical column depth of Martian atmosphere. (Black) Average $\langle Q \rangle$ of the Martian radiation field.



Dose Equivalent Rate H



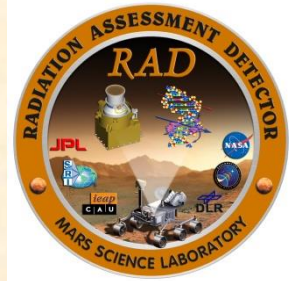
- Dose equivalent is the product of measured **absorbed dose** $\times \langle Q \rangle$...
- ... and provides a better insight into the radiation exposure **relevant to astronaut health**
- Dose equivalent rates highest during solar min (**725 $\mu\text{Sv/day}$**)...
- During solar max: **425 - 450 $\mu\text{Sv/day}$**



RAD dose equivalent rates calculated from RAD absorbed dose rate measurements and $\langle Q \rangle$ of the radiation field.



RAD Acknowledgments



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