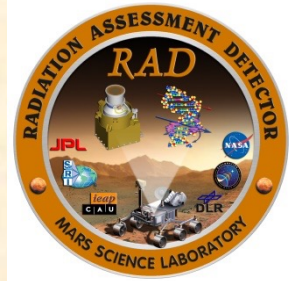




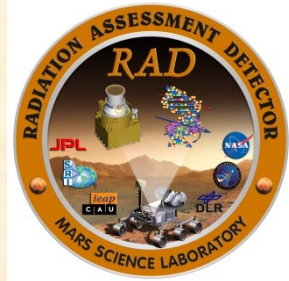
Recent RAD Results: The Oct 28, 2021 & Feb 16, 2022 Solar Storms & Radiation Shielding at Paraipetuy Pass



Bent Ehresmann (SwRI),
Don Hassler (SwRI), Cary Zeitlin (NASA), Robert F. Wimmer-
Schweingruber (CAU), & the RAD Science Team

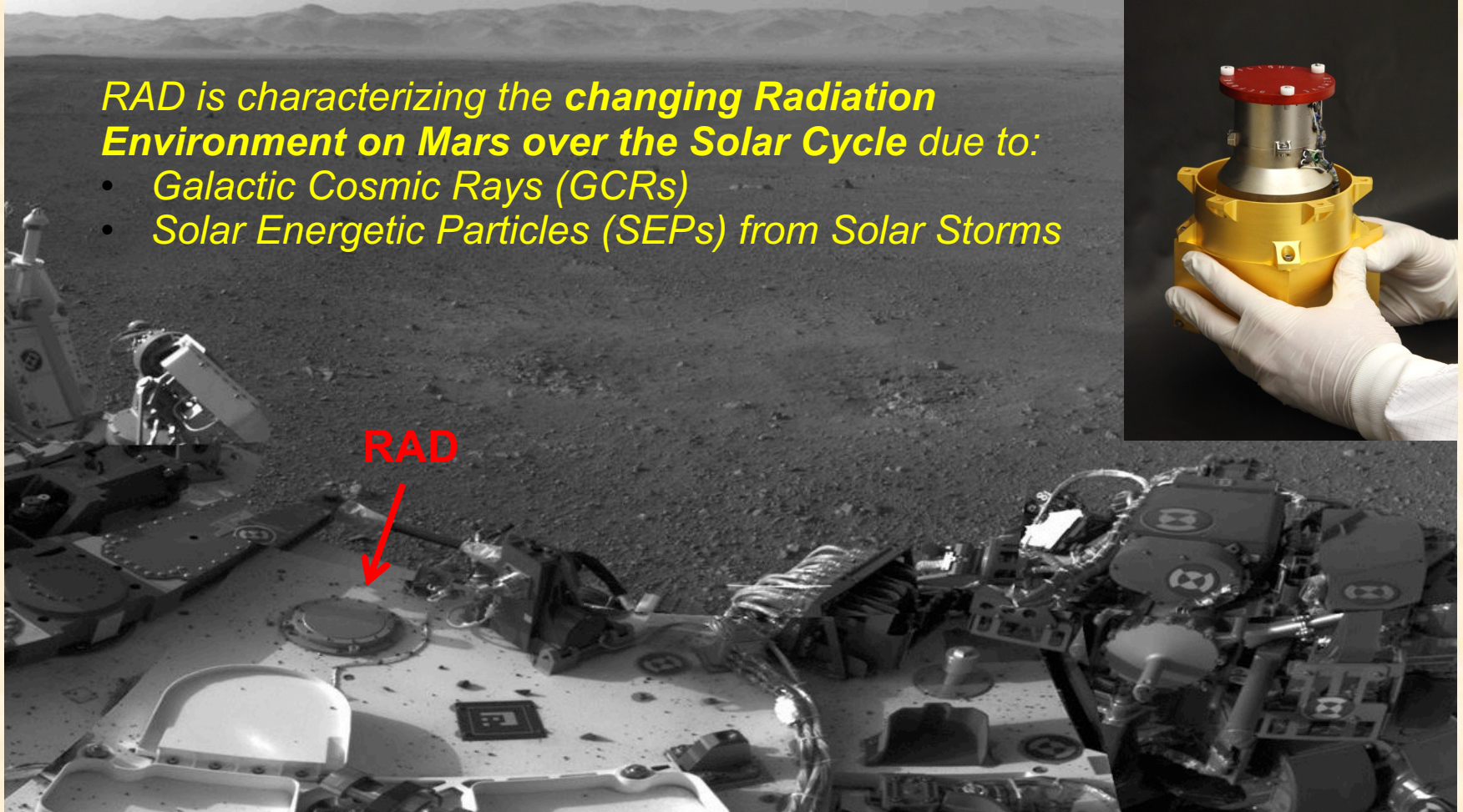
7 September 2023

RAD on the Mars rover *Curiosity* is a working Asset on the Surface of Mars...



*RAD is characterizing the **changing Radiation Environment on Mars over the Solar Cycle** due to:*

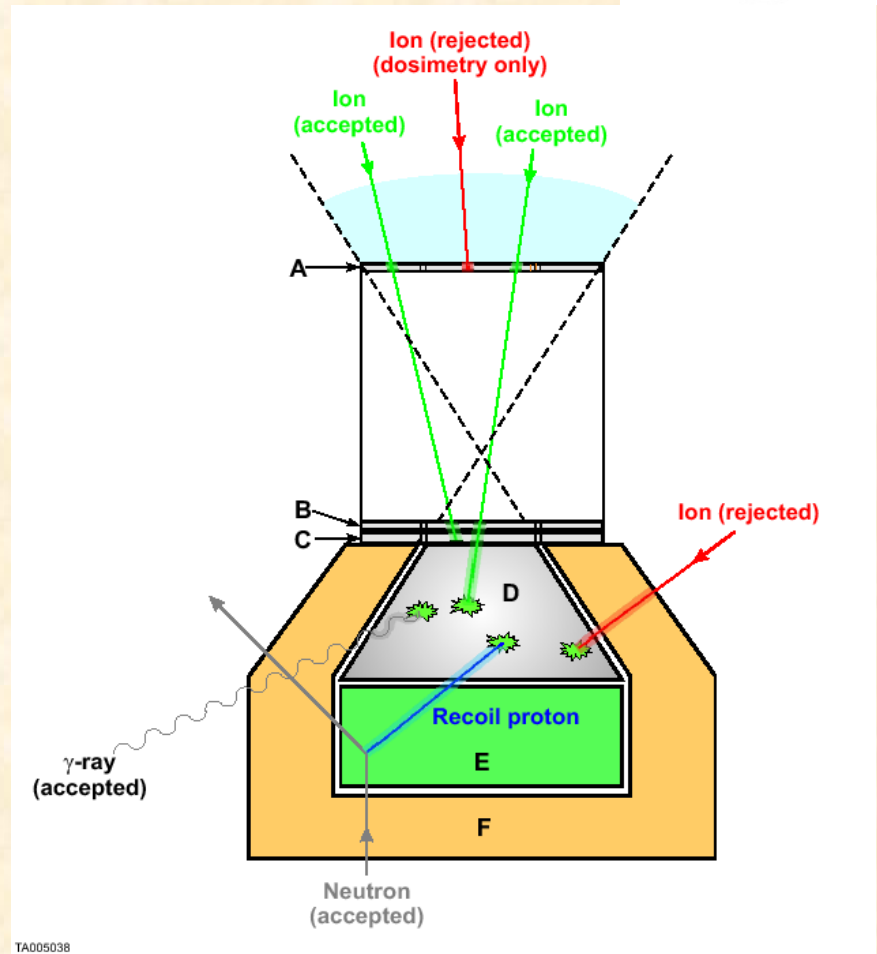
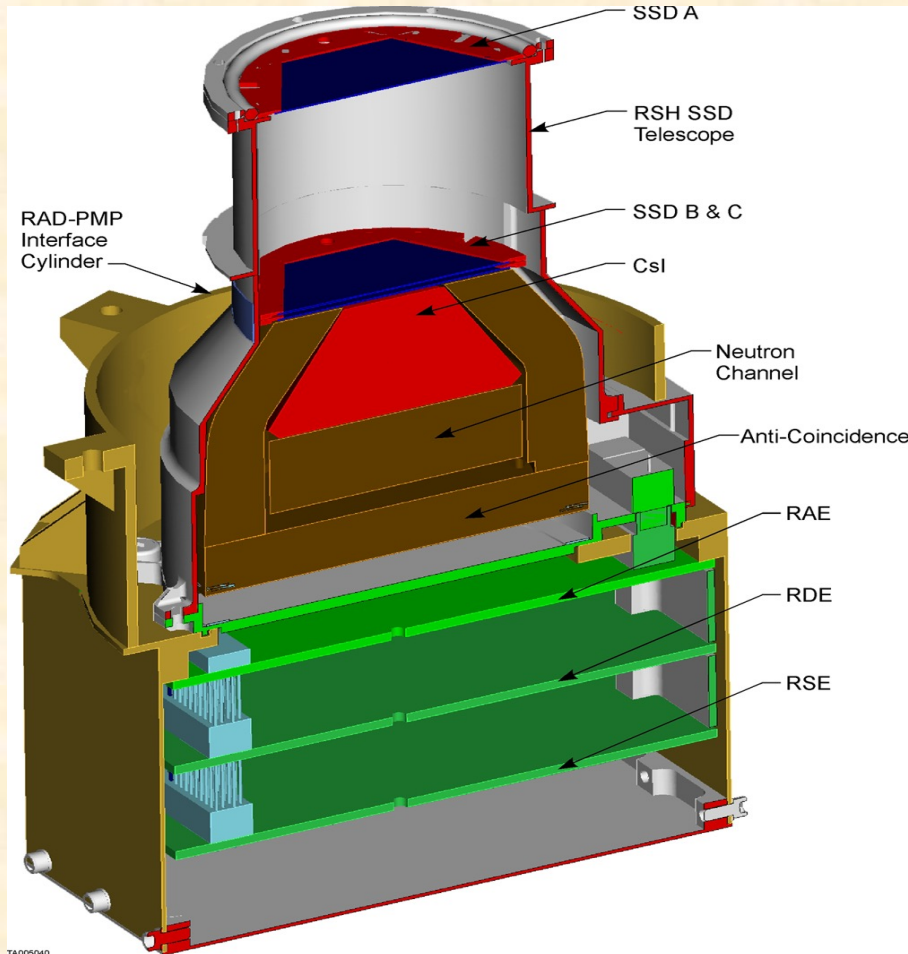
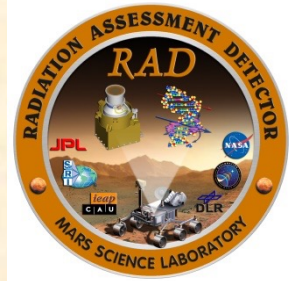
- Galactic Cosmic Rays (GCRs)*
- Solar Energetic Particles (SEPs) from Solar Storms*





C | A | U

RAD Cut-away View and Principle of Operation



A Solid State Detector (SSD) A

B SSD B

C SSD C

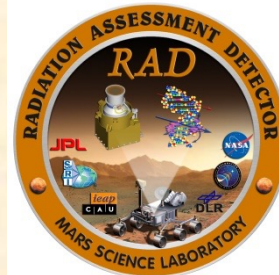
D Cesium Iodide (CsI)

E Neutron Channel (Bicron 432M scintillating plastic)

F Anti-coincidence Shield



MSL RAD: NASA's Newest Member of the Heliophysics System Observatory (HSO)



HELIOPHYSICS SYSTEM OBSERVATORY

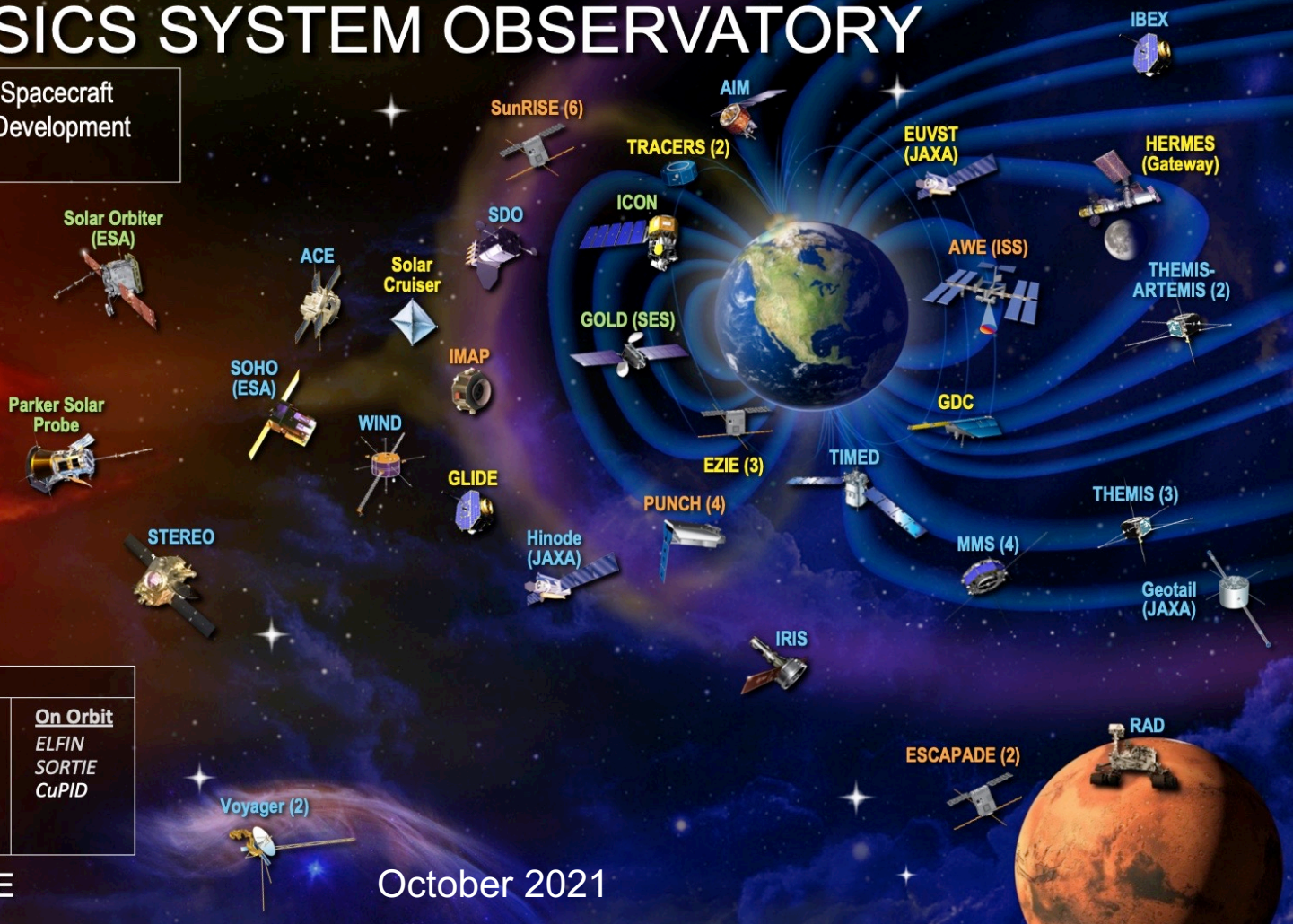
- 20 Operating Missions with 27 Spacecraft
- 12 Missions in Formulation or Development
- 6 Under Study

■	FORMULATION
■	IMPLEMENTATION
■	PRIMARY OPS
■	EXTENDED OPS

CubeSats		
In Development		On Orbit
AEPEX	CuSP	LLITED
AERO / VISTA	DAILI	MinXSS-3
CIRBE	Dione	petitSat
CODEX	GTOsat	REAL
CURIE	LAICE	SPORT

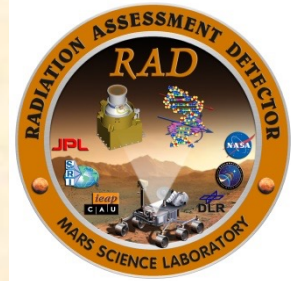
OPERATING & FUTURE

October 2021

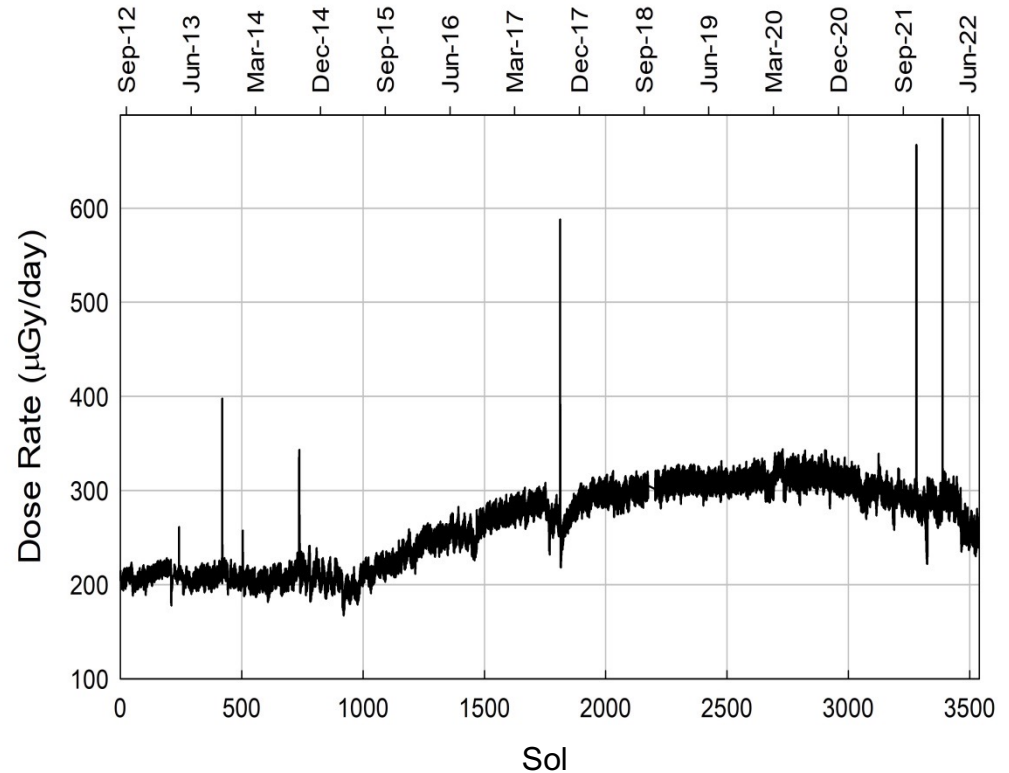




RAD Radiation Dose Rate Measurements over the Course of the MSL Mission



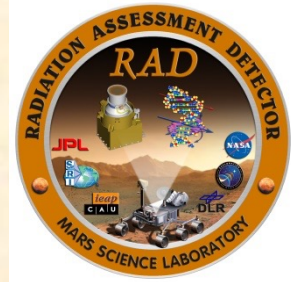
- We are now past Solar Minimum and on our way into the Maximum of Cycle 25
- Peak GCR dose rates were observed in March / April of 2020
- As the sun gets more active on its way to Solar max, large solar storms are more likely
- RAD saw its two biggest solar events to date in October 2021 & February 2022



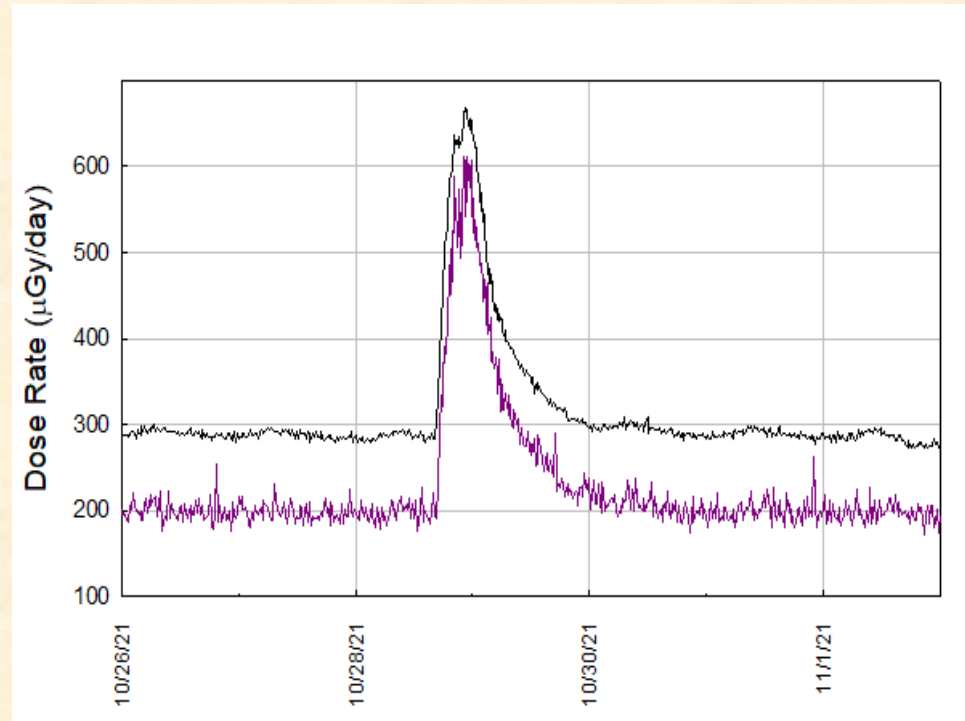
RAD Outlook for Curiosity's fourth extended mission (2023-25):

- complete an entire 11-year solar cycle of observations
- characterize potentially large/extreme events which can occur during Solar Max

The Solar Storm of Oct 28, 2021



- Dose rates in **E** increased by ~350 $\mu\text{Gy/day}$ (factor of ~2.25) over the course of a few hours, reaching a peak dose of ~670 $\mu\text{Gy/day}$
- **B** dose rates increased by ~400 $\mu\text{Gy/day}$ (factor of ~3)
- **B** is less shielded, so more lower-energy protons can reach the detector (compared to **E**)
- *This shows the event was stronger than the biggest previous event seen by RAD in Sep 2017*



Black: Dose Rate in E (tissue-equivalent)
Magenta: Dose Rate in B (Silicon)

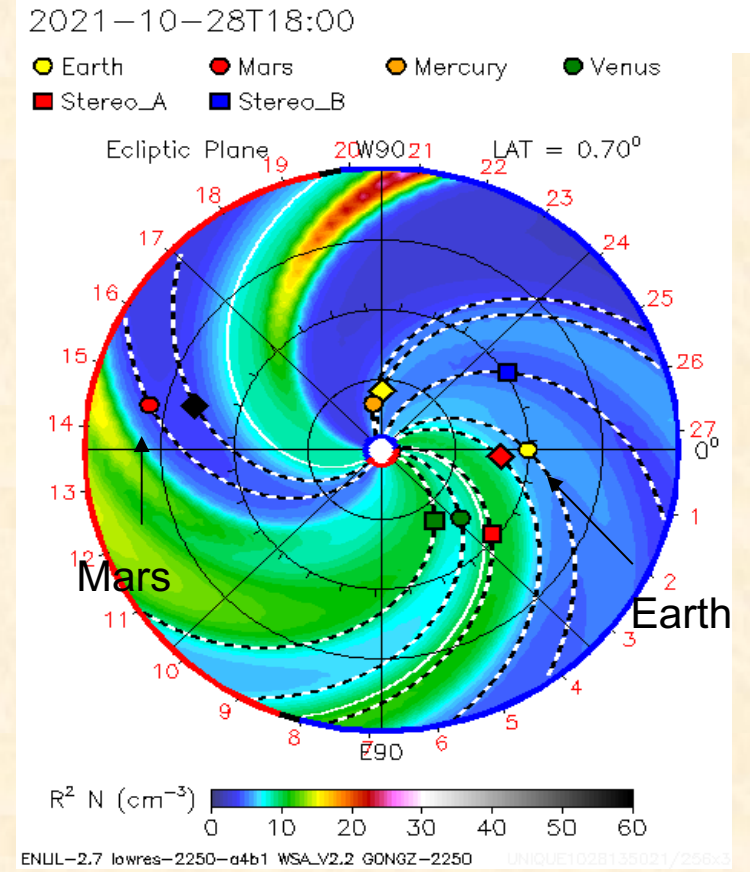
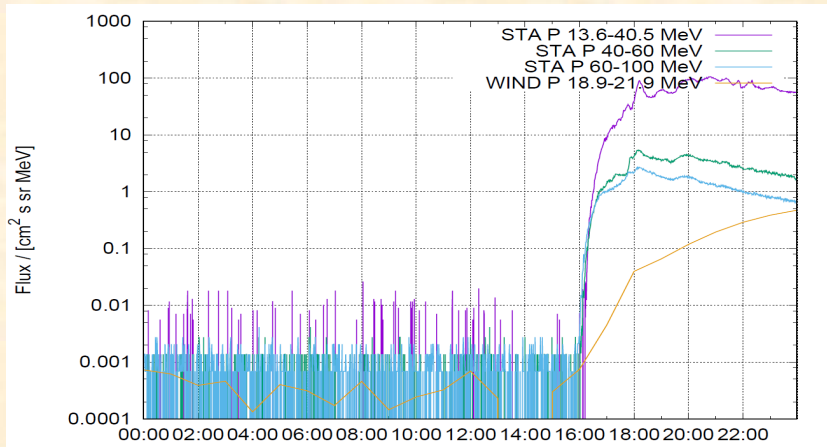


C | A | U

The Solar Storm of Oct 28, 2021



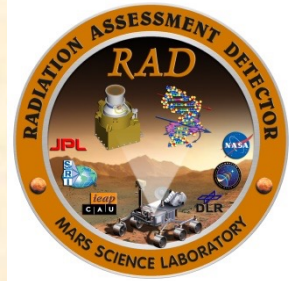
- The event was caused by the first X-class flare of the solar cycle, followed by a halo CME
- Mars & Earth were magnetically well-separated by $\sim 180^\circ$, but the event was still seen at both planets
- ENLIL forecasting didn't predict such a widespread event
- Flux increases were seen by several spacecraft throughout the heliosphere
- Understanding these widespread events is crucial for space weather prediction



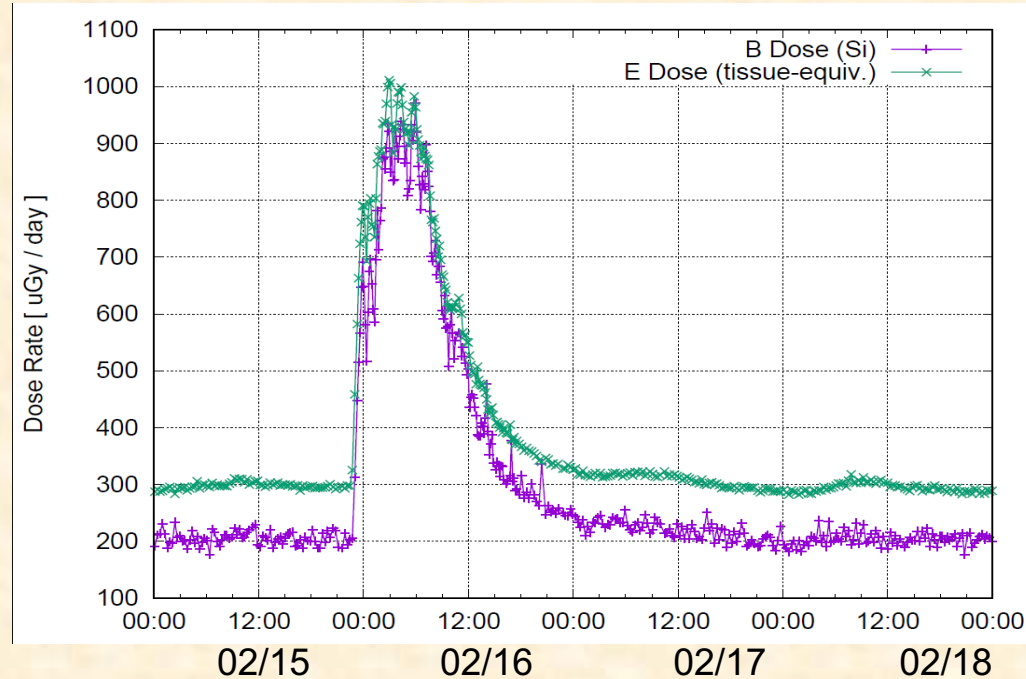
- RAD provides highly important measurements from Mars to cross-compare with the fleet of particle detectors in the heliosphere



The Solar Storm of Feb 16, 2022



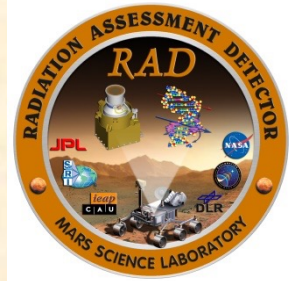
- The 16 February solar storm was the biggest event detected by RAD to date
- It marks the first time RAD measured dose rates **>1000 uGy/day** on the Martian surface
- Dose rates increased by factors of x3.4 (in the E detector) and x4.5-5 (in the B detector)



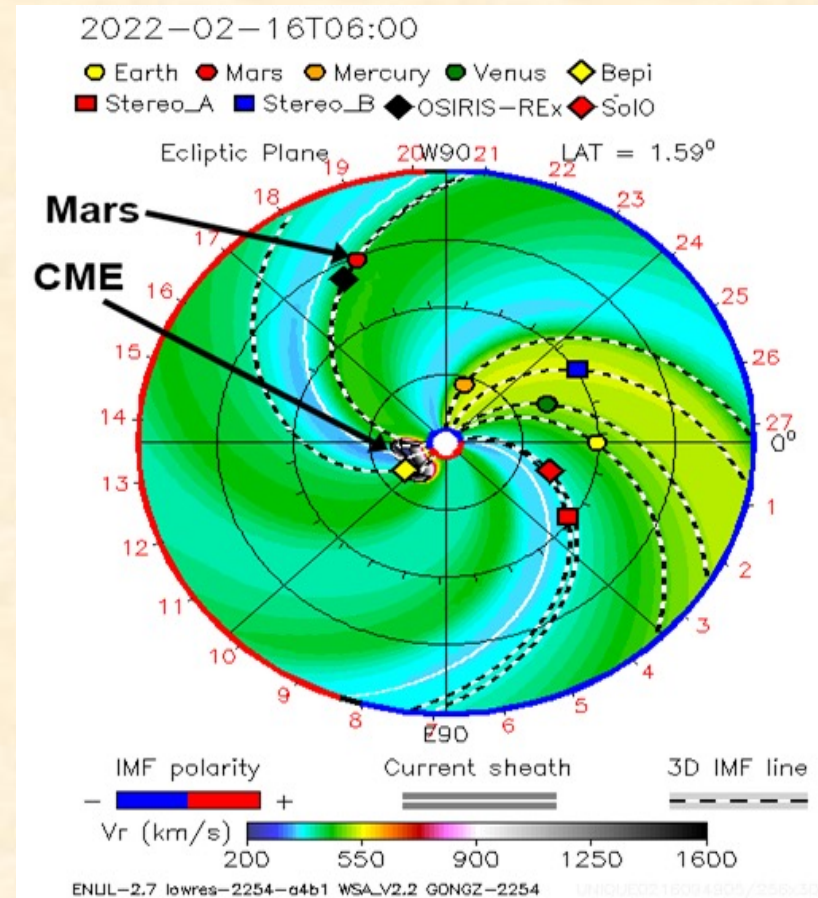
- The event was not seen at Earth (no particle flux increase)
- Other spacecraft such as STEREO-A only saw small increases in flux



The Solar Storm of Feb 16, 2022

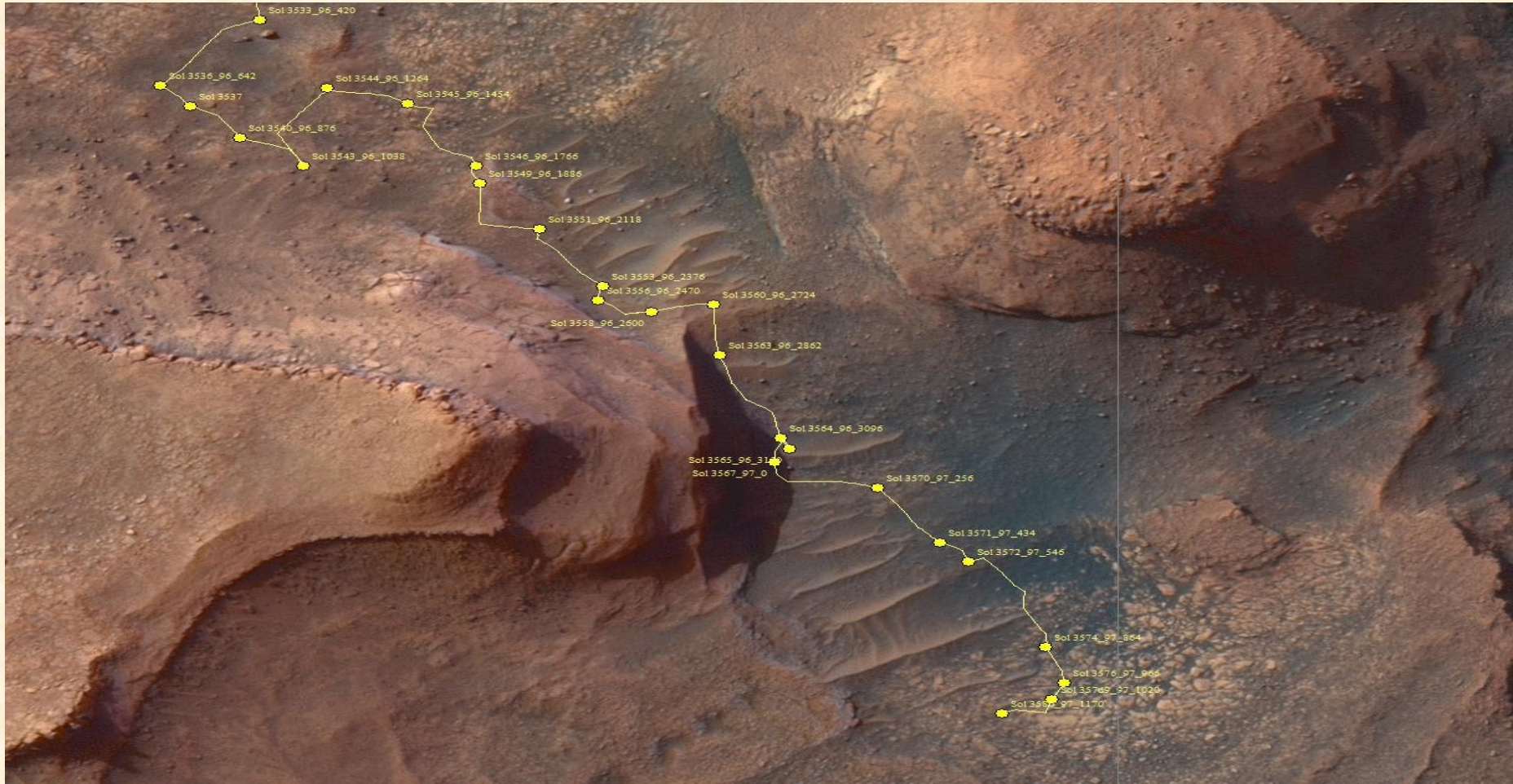
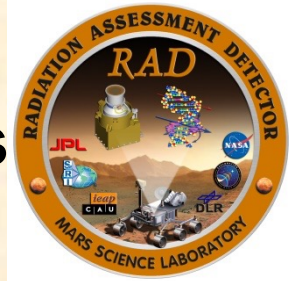


- Mars & Earth were separated by $\sim 135^\circ$
- The event was caused by a CME launched from the backside of the sun
- Mars was still connected to the flank of the CME
- SEP measurements spread-out throughout the heliosphere are highly important to understand the longitudinal spread of these events
- Space agencies planning space station EVAs or off-base Lunar or Martian “sorties”, need to understand if any active regions on the far side of the sun could develop into solar storms that might still reach the astronaut on the near side...
- ... to provide a reliable All-Clear forecast





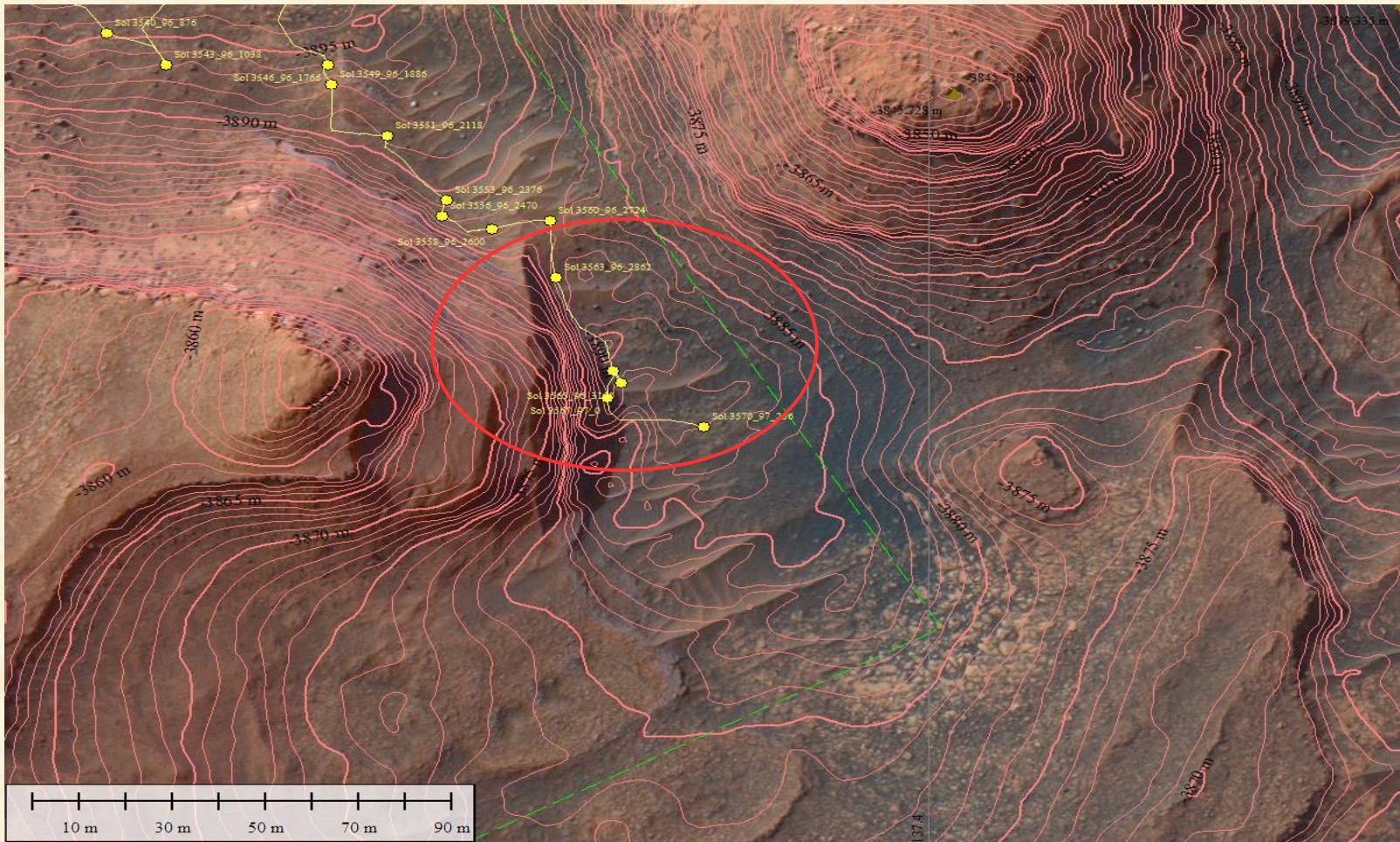
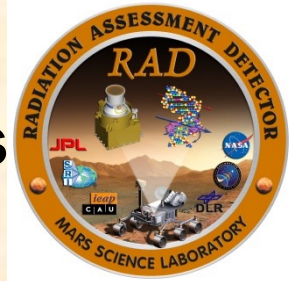
Radiation Shielding at Paraitepuy Pass



The drive through Paraitepuy Pass & particularly the close approach to Bolivar offered a new & exciting opportunity to measure the radiation shielding effect



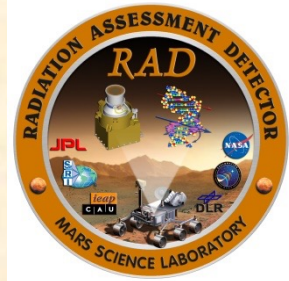
Radiation Shielding at Paraitepuy Pass



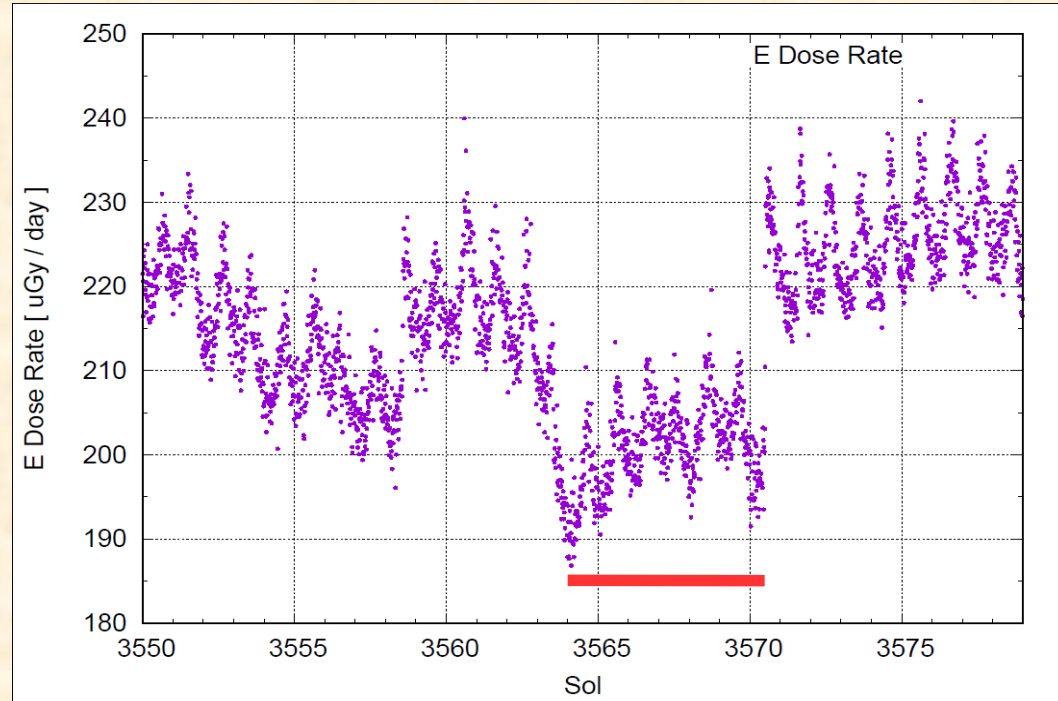
From sol 3563 to 3570 in the Pass & in close proximity to Bolivar, RAD measured a distinct decrease in dose rate



Radiation Shielding at Paraitepuy Pass



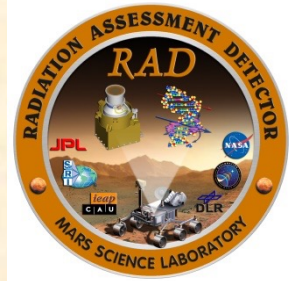
- Dose rate decrease was >11-12%
- Analysis needs to take into account that Curiosity was already traversing close to Bolivar & Deepdale for >20 days before the closest approach...
- ... to correctly estimate the expected / unshielded dose rate and thus the decrease from the shielding effect



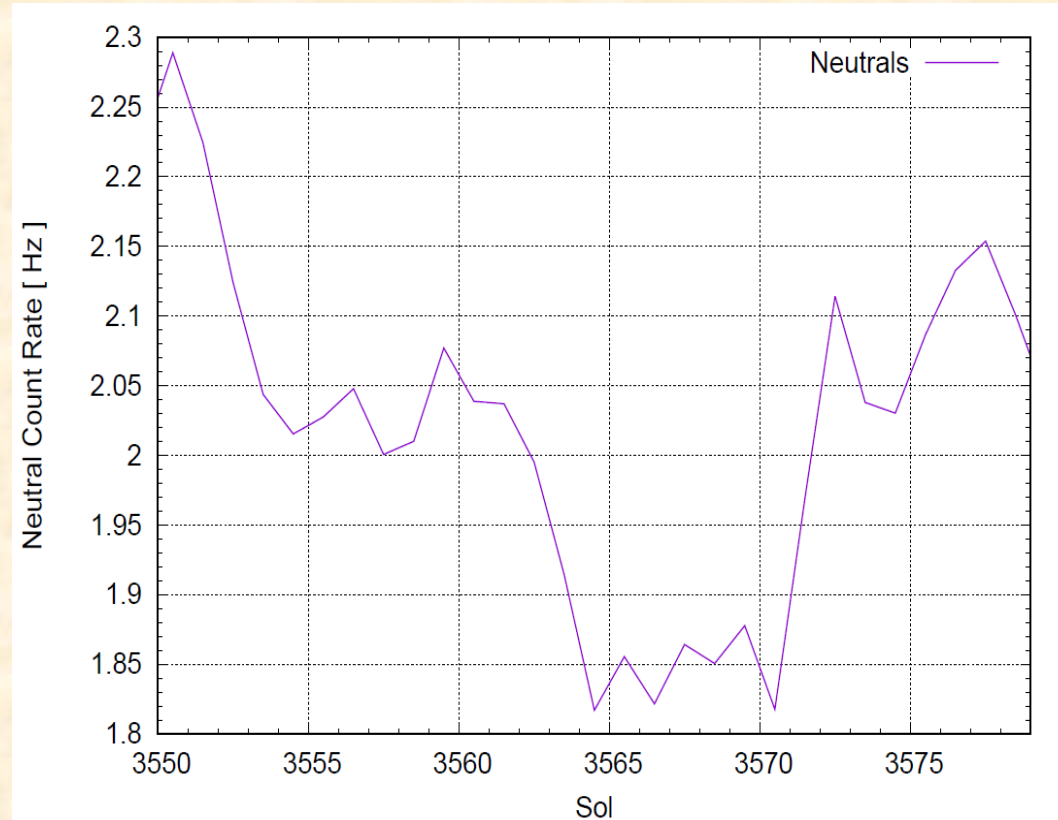
- Decrease in dose rate in the same order of magnitude (at least ~ 10%) as previous dedicated measurement campaigns at Maria Gordon Notch at Mt Mercou...



Radiation Shielding at Paraitepuy Pass – Neutral Particles

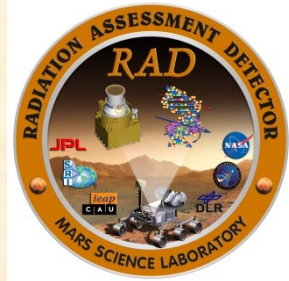


- Decrease in the neutral particle radiation was also >10-11%
- Neutrals here are gamma-rays and neutrons with energies >10 MeV
- (Important when comparing to shielding effects seen by DAN Passive which measures much lower-energy thermal neutrons)





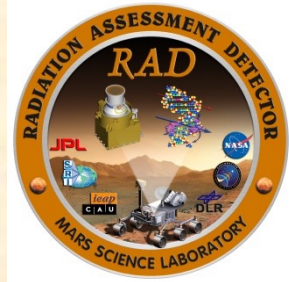
Radiation Shielding at Paraitepuy Pass – Summary



- Making use of available in-situ shielding could be an important aspect to reduce mass & cost requirements for future Mars missions
- Measuring the radiation shielding effect provides input for the planning of radiation storm shelters for future human explorers of Mars
- RAD measurements under different shielding conditions (different occlusion angle & azimuth range) are useful to validate angular-dependent radiation transport models used to estimate the effectiveness of storm shelters
- Additionally, shielding measurements let us also calculate the changing contribution of the MMRTG to the RAD dose rate...
- ... by comparing the dose decreases in detectors B (sensitive to RTG) & detector E (insensitive to RTG contribution)
- The RAD team is currently preparing a manuscript summarizing all observed radiation shielding measurements to date...
- ... and would like to thank the project for accommodating these highly interesting & important shielding measurements!



Thank you!



- RAD is supported by NASA SMD/Heliophysics and HEOMD/AES under JPL subcontract #1273039 to SwRI.
- ...and by DLR in Germany under contract with Christian-Albrechts-Universität (CAU).

